Evaluating the Consistency of Ice Thickness Models Across Regions in Predicting Subglacial Overdeepenings

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Submitted in accordance with the requirements for the degree of

Master of Science by Research

University of Leeds

School of Geography

Acknowledgements

This Master's by Research thesis was supervised by Dr. Duncan J. Quincey and Dr. Jonathan L. Carrivick of the University of Leeds. I extend my deepest gratitude to them for their mentorship, which went far beyond academic guidance. Their advice, expertise, and unwavering support helped me overcome challenges and successfully complete this work. They are a great combination of supervisors to work with. It would be a pleasure to work with them again sometime in the future.

I am also deeply grateful to Dr. Jenna Sutherland of Leeds Beckett University for providing the ice thickness data for OGGM in the central Himalayas, which was integral to this research.

My heartfelt thanks go to my family and friends for their support and encouragement during difficult times. A big thanks to everyone in Office 10.08, Garstang, for always being there to help with doubts and for brightening up the space with treats.

I am thankful to all the members of the Ice Club and the River Basin Processes and Management Cluster for their constructive feedback, which contributed valuable ideas to this thesis. Further, I would like to thank all the postgraduate researchers and staff of the School of Geography at the University of Leeds for providing a supportive and inspiring academic environment.

Abstract

Glaciers in high mountain regions are retreating rapidly due to climate change, exposing subglacial environments sculpted over millennia. Overdeepenings are a common part of the post-glacial landscape, and their development impacts landscape evolution, proglacial sediment dynamics, glacial lake formation, and associated hazards like Glacial Lake Outburst Floods (GLOFs). While Geographic Information System (GIS)-based approaches integrating Digital Elevation Models (DEMs) and ice thickness models have become critical for estimating overdeepening locations and characteristics, there has been little assessment of the extent to which the results are dependent on the input data that are used. To address this, the study first evaluated the relative influence of DEMs and ice thickness models on overdeepening predictions in the Central Himalayas. It compared results from five ice thickness models: Farinotti's Ensemble, Huss and Farinotti, GlabTop2, OGGM, and Millan; paired with four DEMs: Copernicus, ASTER GDEM, SRTM, and ALOS PALSAR. The analysis revealed that ice thickness models exert a greater influence than DEMs on overdeepening predictions. The combined framework was then applied to other regions, including the Peruvian Andes, Alaska, the European Alps, and New Zealand's Southern Alps, to further assess overdeepening variability and its dependency on the choice of ice thickness dataset. While no consistent trends emerged across regions and models, certain ice thickness models, such as OGGM and GlabTop2, were found to overestimate overdeepening parameters in many instances. Despite some commonalities, the maximum overlap in overdeepening predictions between any two models was 53.6%. To address the uncertainty in using a single ice thickness dataset to predict overdeepening locations, a potential framework that integrates multiple datasets is therefore proposed. In the absence of robust validation data, this approach could be used to provide confidence bounds on the estimation of overdeepening size and volume, helping to communicate uncertainties to stakeholders. More generally, the findings highlighted the variability in overdeepening parameters, emphasizing the need to account for this range when using overdeepenings as a foundation for further analysis. Establishing a clear purpose before estimating overdeepenings is recommended, as it can guide the selection of the most appropriate ice thickness model. This study provides insights into the consistency among ice thickness models and underscores the importance of multi-model approaches for robust and reliable assessments.

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List of Abbreviations

GLOF: Glacial Lake Outburst Flood DEM: Digital Elevation Model OGGM: Open Global Glacier Model RGI: Randolph Glacier Inventory SRTM: Shuttle Radar Topography Mission ASTER: Advanced Spaceborne Thermal Emission and Reflection Radiometer ALOS: Advanced Land Observing Satellite PALSAR: Phased Array type L-band Synthetic Aperture Radar GLIMS: Global Land Ice Measurements from Space HIGTHIM: Himalayan Glacier Thickness Mapper VOLTA: Volta and Topography Estimation IDW: Inverse Distance Weighing GIS: Geographic Information System

1. INTRODUCTION

Changes in both temperature and precipitation over the past four decades (Allen et al. 2018) have led to accelerating glacier retreat (Davaze et al. 2020; Hugonnet et al. 2021) in high mountain regions. This retreat can lead to changes in geomorphology (Prothro et al., 2018) and glacier dynamics (Chandler et al. 2016), as well as revealing the subglacial environment that has been sculpted by processes of erosion and deposition for thousands of years. Overdeepenings are subglacial depressions that are usually formed because of the erosive nature of the glacier (Hooke, 1991; Cook and Swift, 2012). The overdeepenings thus exposed can form glacial lakes, and can also trap sediments that would previously have been delivered to downstream areas (Magrani et al. 2020), influencing the landscape evolution.

Given these overdeepenings cannot be directly observed, GIS based approaches that estimate the subglacial topography are required. There has been a recent surge in studies that have used high-resolution DEMs combined with ice thickness models to understand the morphometry and overdeepening locations (Patton et al. 2016; Haeberli et al. 2016a; Magrani et al. 2020; Schlunegger et al. 2024). The vast majority of such studies rely on a single input DEM and a single ice thickness model to evaluate the subglacial topography. However, the diversity of available terrain datasets and ice thickness models introduces variability in overdeepening predictions (Gharehchahi et al. 2020; Magnin et al. 2020; Otto et al. 2022), and this variability is yet to be fully quantified, or resolved.

Differences in underlying principles, data sources, and times of acquisition between DEMs and ice thickness models complicate direct comparisons. The lack of systematic intercomparison between models and DEMs further limits understanding of any regional variability and precludes a full assessment of the robustness of predictions. Glacier morphology varies significantly across regions (Haeberli et al. 2016), largely because of the combination of spatially variable climatic settings and tectonically controlled topographic characteristics, influencing the accuracy of overdeepening estimates and underscoring the need for context-specific approaches.

To address these gaps, this study aims to evaluate the consistency of different DEMs and ice thickness models in predicting overdeepening characteristics and to assess their applicability across diverse mountain regions. By comparing outputs from multiple datasets and combining model results, this study seeks to develop a more robust framework for overdeepening detection. The thesis is structured as follows: first, a review of the current literature summarises the key studies that have considered GIS-based approaches for estimating subglacial overdeepenings, and assesses their limitations (Chapter 2) providing the foundation to identify current gaps (Chapter 3); the methods used in the current study are then introduced (Chapter 4); followed by an explanation of the analysis results (Chapter 5); the underlying reasons for the observed patterns and findings are then discussed in detail (Chapter 6) and then the entire study is summarised in the concluding chapter (Chapter 7).

2. LITERATURE REVIEW

2.1. OVERDEEPENINGS: FORMATION, AND RELEVANCE IN DEGLACIATING LANDSCAPE

Overdeepenings are topographic depressions located beneath the ice in a glacier (Penck and Brückner, 1909; Alley et al. 2019). They could be entirely erosional or non-erosional and be present as a part of the bedrock, and they may be of glacial (Figure 2.1) or non-glacial origin. Once exposed at the surface they can often fill with water to form glacial lakes. When located in sensitive and dynamic regions of the glacier (which is usually at the snout of the glaciers) they can affect the glacier

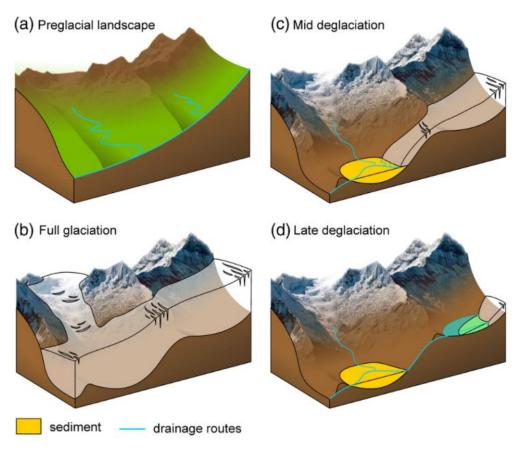


Figure 2.1: Schematic diagram of the subglacial topography (overdeepenings in yellow and green) being exposed at the surface (Cook and Swift 2012)

dynamics and hydrology to a great extent (Cook and Swift, 2012).

Overdeepenings can occur in any part of the glacier, but in a valley, glacier setting they are mostly found along the main tongue (Viani et al. 2020) in areas where there is a sudden change in the bed gradient (Emmer et al. 2015a) and the rate of velocity is high (Lloyd et al. 2023). A possible way of overdeepening formation can be seen in Figure 2.1. Large overdeepenings tend to develop where

the rocks are least resistant (Magrani et al. 2020), and they play an important role in storing, releasing and transferring sediments to the downstream areas (Liebl et al. 2021).

A study in the Austrian Alps has shown that 33% of all lakes are present within the terminus area of these glaciers (Emmer et al., 2015b). Although overdeepenings represent a small fraction of total glacier volume - approximately 3% in the Swiss Alps, 3 - 4% in the Himalayan-Karakoram region (Haeberli et al., 2016), and about 0.5 -1% in the Peruvian Andes (Colonia et al., 2017) - their downstream impacts are disproportionately significant. They influence sediment transport, hydrology, and glacial lake formation, highlighting their importance in both glacial and post-glacial environments.

2.2. PREVIOUS WORK TO IDENTIFY SUBGLACIAL OVERDEEPENINGS

The glaciers in the Andes (Colonia et al. 2017; Drenkhan et al. 2018, 2019; Tweed and Carrivick, 2015), High Mountain Asia (Kapitsa et al. 2017; Furian et al. 2021, 2022), the Himalaya-Karakoram (Linsbauer et al. 2012; Allen et al. 2016a; Haeberli et al. 2016; Bhushan et al. 2017; ; ; Sattar et al. 2019; Majeed et al. 2021a; Majeed et al. 2021b; Zhang et al. 2022; Rana et al. 2023; Srinivasalu et al. 2024) and the European Alps (Frey et al. 2010; Haeberli et al. 2016; Magnin et al. 2020; Gharehchahi et al. 2020; Otto et al. 2022 and the New Zealand Southern Alps (Carrivick et al. 2022) have previously been studied for overdeepening locations and characteristics. The third pole and the Himalayas have been studied extensively, and within the Himalayas in particular, the western Himalayas and the Karakoram region have been studied most frequently (Table 2.1).

Paper	Region	Model	DEM	Glacier outlines	Area Threshold	Major Findings
Frey et al. 2010	Swiss Alps	Shallow Ice Approximati on	Swiss ALTI		6250 m² for glaciers	Introduced the method in which the ice thickness is subtracted from the DEM to produce the overdeepenings. Suggested that SRTM and ASTER can be used for these analysis for other regions of the world.
Haeberli et al. 2016	Swiss Alps	GlabTop	Swiss ALTI	Swiss Glacier Inventory	10000 m^2	The modelled overdeepening in the Swiss Alps were 3% of the total glacier volume. The modelled overdeepening in the Himalayan Karakoram region
	Himalayan- Karakoram	GlabTop2	SRTM	RGI (Pfeffer et al. 2014)		was 3-4% of the total glacter volume. The overdeepenings in the Himalayas have area more than 50 km².
Gantayat et. 2017	Indian Himalayas	нідтнім	ASTER , SRTM	RGI (Pfeffer et al. 2014)		They suggested that a glacial lake can form on the snout of Drang Drung glacier and the already existing moraine dammed lake near the snout of Samudra Tapu glacier might increase in area and volume.

Table 2.1: Literature Review of studies on overdeepenings

Paper	Region	Model	DEM	Glacier outlines	Area Threshold	Major Findings
Linsbauer et al. 2012	Himalayan- Karakoram	GlabTop2	SRTM	Global Land Ice Measurements from Space(GLIMS) database and Randolph Glacier Inventory (RGI)	10000 m²	The modelled overdeepenings covered 3.8% of total volume of the glaciers in the Himalayan Karakoram region. These overdeepenings were shallower in Karakoram and Central Himalayas as compared to the deeper overdeepenings in western and eastern Himalayas.
Magnin et al. 2020	Western European Alps	GlabTop	IGN, RGD, ASTER	Glacier outlines drawn within GlaRiskAlps project	10000 m²	The depth of the overdeepenings varies according to the DEMS, IGN DEM predicted the deepest, largest and most voluminous GBOs whereas ASTER produced shallow GBOs. The GBOs in the region are similar to those present in the Peruvian Andes. Only 2 lakes were likely to form in the region given the retreat rate then.
Srinivasalu et al. 2024	Western Himalayas	Gantayat et al., 2017	ASTER	RGI (Pfeffer et al., 2014), GAMDAM	10000 m²	A total 132 glaciers were taken to estimate the overdeepenings in the Sutlej and Beas basin, which resulted in 118 potential glacial lakes.
Drenkhan et al. 2018	Andes	GlabTop	SRTM	Landsat 5 and Sentinel-2, NDSI		Total of 219 overdeepenings were identified in the glacier bed, where 49 met the slope threshold and out of which 20 sites were selected as potential future lake formation.

Paper	Region	Model	DEM	Glacier outlines	Area Threshold	Major Findings
Drenkhan et al. 2019	Andes	GlabTop				Future lake formation shows that they can form important water resources though they might not compensate the water loss by the glacier retreat.
Kapitsa et al. 2017	Tien Shan	GlabTop2	SRTM	Serveskiy et al. 2016	11000 m²	They identified 513 overdeepenings in 2000 glaciers, 67% of which matched with the present glacial lakes. If the area threshold is increased to 20000 m ² , the match increases to 85%. The location of the lakes is quite robust but the morphometry of the lakes is that robust.
Colonia et al. 2017	Peruvian Andes	GlabTop	SRTM	National Glacier Inventor y of Peru	10000 m²	There are higher chances of lake formation in Cordillera Vilcanota than Cordillera Blanca. They estimated that 201 overdeepenings were formed that consist of 0.5 to 1% of the total glacier volume there. Most of the lakes are likely to form within the next few years.
Viani et al. 2020	Aosta Valley	GlabTop2	Regional DEM by RAVA		10800 m²	The volume of overdeepenings is 0.8% of the total glacier volume which is like that found in the Peruvian Andes by Colonia et al., 2017. This value is less when compared to the percentage of total volume in Swiss Alps and Himalayan Karakoram.
Furian et al. 2021	High Mountain Asia	Farinotti's Ensemble	ALOS PALSAR		10000 m²	They predicted 25285 overdeepenings in the region but if the threshold is increased it produces 2700 overdeepenings in the 669 glaciers. The size and location of the overdeepenings is in good agreement with the studies.

Paper	Region	Model	DEM	Glacier outlines	Area Threshold	Major Findings
Furian et al. 2022	High Mountain Asia	Farinot ti's Ensem ble	ALOS PALSAR		100000 m ²	The future lakes are most likely to develop in the first half of the 21st Century the increase in the volume happens in the second half of the century.
Otto et al. 2022	Austrian Alps	GlabTo p2, HF Model	ALOS PALSAR	Glacier outlines were drawn from orthophotos and airborne lasers can data with varying acquisition dates.	Glaciers with area greater than 10000 m ²	The study area has 525 glaciers which were taken and between 86 and 198 overdeepenings were predicted in the region. Both the models performed better for the valley glaciers than for the smaller glaciers.
Gharehchahi et al. 2020	Swiss Alps	VOLTA	SwissA LTI and DHM25 L2			The large glaciers which have an area greater than 5 $\rm km^2$ are better compared to the influence of DEMs on the overdeepenings. They estimated the formation of 125 overdeepenings using Swiss ALTI and 113 using DHM25 L2. The location, area and depth of the overdeepenings were influenced by the choice of the DEM. Out of the 171 overdeepenings that were identified 100 of them were classified as highly likely to form a glacial lake in the future.
Carrivick et al. 2016	South American Glaciers	VOLTA	ASTER			A total of 617 were considered for the study and 282 overdeepenings were identified.

Paper	Region	Model	DEM	Glacier outlines	Area Threshold	Major Findings
Bhushan et al. 2017	Gangotri Glacier	Gantayat et al., 2017	Cartosat -1			They identified 8 possible sites for formation of supraglacial lakes. The overdeepenings present at the snout of the glacier are most likely to form glacial lakes in the future.
Majeed et al. 2021b	Gya Glacier	GlabTop	ALOS PALSAR	Glacier outlines were taken from satellite data		They evaluated the evolution of Gya glacier and how the lake has evolved. The proglacial lake was spreading over ~11 ha, delineated from satellite data.
Rana et al. 2023	Parkachik Glacier	Gantayat et al., 2017	CORONA KH4, Landsat and Sentinel- 2A			They found that nearly 74 potential future lakes would form. They examined the expansion of lakes from 1964 to 2017 and found that their total area increased.
Rashid et al. 2022	Pangong Region	GlabTop	ALOS PALSAR			They used this study to show how much lake could expand in the future.

Paper	Region	Model	DEM	Glacier outlines	Area Threshold	Major Findings
Sattar et al. 2019	Dhauliganga Basin	Gantaya t et al., 2017	ASTER	RGI (6.0)	Area of glaciers > 1km ²	They found out that 54 overdeepenings were estimated for the glaciers that have an area of 2.85 km ² .
Carrivick et al. 2022	New Zealand Southern Alps	Farinotti 's Ensembl e	ALOS PALSAR	RGI (6.0)	100000 m²	This paper provides the inventory of the ice- marginal proglacial lakes across the New Zealand Southern Alps and showed a positive relationship between glacier mass balance and lake growth. It was also concluded the number and size of the ice marginal lakes would increase in the future which would affect the drainage.
Zhang et al. 2019	Central Himalayas	GlabTop	TanDEM- X, SRTM	Using LANDSAT images	10000 m²	They found that nearly 74 potential future lakes would form. They examined the expansion of lakes from 1964 to 2017 and found that their total area increased. They compared their findings for the region using the GPR readings and found that the location of the overdeepenings was robust.
Allen et al. 2016b	Himachal Pradesh	GlabTop 2	ASTER		10000 m²	They found out that 279 lakes would form across the entire Himachal Pradesh. This study shows that the potential risk of GLOF hazard would increase in future.

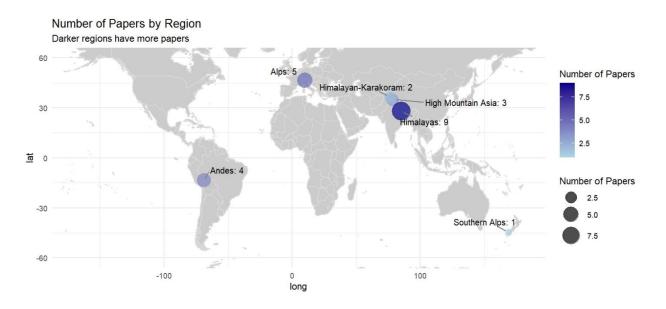


Figure 2.2: Map of the regions that were studied

There have been numerous studies (Figure 2.2) on overdeepenings using the method given by Frey et al., (2010). Studies have relied on a single DEM and ice thickness model to detect overdeepenings and project future scenarios, often without addressing the substantial uncertainties that are inherent in such predictions (Drenkhan et al., 2019; Furian et al., 2022). Most of these are based on High Mountain Asia and Himalayas combined (Bhushan et al. 2017; Majeed et al. 2021a; Furian et al. 2022; Rashid et al. 2022; Rana et al. 2023) as seen in Figure 2.2. This is of some concern where study outputs are classified and then communicated in risk terms, given that communities live within the downstream areas that could be affected (Drenkhan et al. 2019). It also means that broad-scale analyses of potential hydropower that could be available from future overdeepenings are also highly uncertain even on a regional scale (Farinotti et al. 2019b).

The choice of DEM has been shown to have a small influence on determining overdeepening locations, but a greater impact on the morphometric parameters of the overdeepenings (Zhang et al. 2019; Gharehchahi et al. 2020). The most commonly used DEMs are ALOS PALSAR, ASTER and SRTM, given that they are widely available and offer relatively high (~30 m) spatial resolution. Studies like Frey et al. (2010) and Gharehchahi et al. (2020), focussing on the European Alps use Swisstopo, which is a local DEM for the region. Other parameters that vary between studies include the size threshold at which the overdeepenings are considered to be significant – the large majority adopt a threshold of 10,000 m² (0.01 km²) to remove individual pixels and small clusters of pixels that most likely arise from artefacts of the processing the ice thickness datasets (see Table 2.1).

Studies employing multiple DEMs or ice thickness models, such as Magnin et al. (2020) for Peruvian Andes and Gharehchahi et al. (2020) for the Swiss Alps, have highlighted that there is uncertainty of \pm 10 m in determining the depth of the overdeepening, and of \pm 24% in determining the ice thickness of the glaciers, respectively, both within ice thickness models, considering their absolute accuracy with reference to measured bedrock elevations. This emphasises on the fact that irrespective of region, DEM and ice thickness model used, the uncertainties remain an inherent part of the analysis.

Otto et al. (2022) combined two ice thickness models (Huss and Farinotti and GlabTop2) with three glacier outline datasets, enhancing confidence in their estimates through a multi-model approach, which seems to show promising application for a more robust evaluation of overdeepening locations and characteristics where a number of independent evaluations are available. Out of all the overdeepenings, 70% overlapped with the already present lakes in the glaciers. However, none of these studies have sufficiently explored any regional variability among models or the choice of DEM.

The accuracy of the glacier outlines used in GIS-based workflows has also been regarded as important because many valley glaciers, especially the ones in the Himalayas, have proglacial lakes that develop at the terminus of the glaciers and hence the dimensions of the lake, and any identified overdeepenings, change every time the outline changes (Paul et al. 2019). Further, several approaches for estimating ice thickness (such as VOLTA; James and Carrivick, 2016, HIGTHIM; Gantayat et al. 2017), use the algorithm to reach an ice thickness of zero at the terminus, which can be erroneous where glaciers terminate with a steep ice face into a proglacial lake (James and Carrivick, 2016; Gantayat et al. 2017). The most widely used glacier outlines have been from the RGI version 6.0 (https://www.glims.org/RGI/randolph60.html; Pfeffer et al. 2014); these are easy to access, generally considered to be one of the most accurate assessments of glacier extent, and can also readily integrate with existing GIS-based workflows. Some studies (Yang et al. 2016; Freudiger et al. 2018) have used satellite images to construct their own glacier outlines, but this is not common.

Studies present overdeepening estimates in terms of specific volume, and area percentages, without adequately addressing how to manoeuvre around the uncertainties inherent in these predictions (James and Carrivick, 2016; Drenkhan et al. 2019; Gharehchahi et al. 2020). An intercomparison of different ice thickness models and the DEMs could help address a major gap and improve the reliability of future hazard assessments

2.3. ICE THICKNESS MODELS

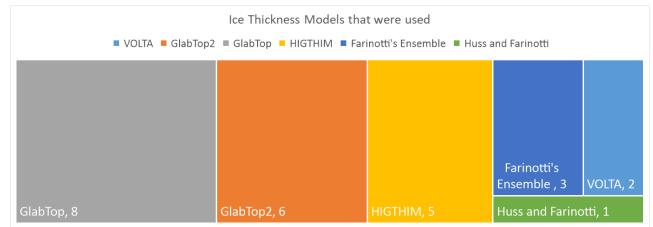


Figure 2.3: Ice thickness models being used in different studies from Table 2.1 with the number of studies they have been used in

Most previous studies have used GlabTop (Paul and Linsbauer, 2012) for ice thickness estimation as seen in Figure 2.3; this was one of the first methods available and the also the programming code for the model is readily accessible. It was therefore integrated into many early analyses of glacier volume estimation, and then subsequently, studies seeking to characterise glacier bed topography. Recent studies have used ice thickness from Farinotti's ensemble (Carrivick et al. 2022; Furian et al., 2021, 2022) because the data are readily available and they are now considered to be the most robust, comprising the best estimations across a range of input models. Unlike global ice thickness models such as GlabTop and Farinotti's ensemble, which are designed for broad applicability, the approach developing by Gantayat et al. (2017), known as HIGTHIM, is specifically tailored to Himalayan glaciers that are valley type (Gopika et al. 2021). This method was employed to estimate ice thickness exclusively within the Himalayas, reflecting its region-specific focus and utility (Bhushan et al. 2017; Majeed, et al. 2021a; Sattar et al. 2019; Srinivasalu et al. 2024).

A small number of studies have shown that different ice thickness models show regional variability. For instance, the GlabTop model underestimated the volume of overdeepenings in the Argentiere glacier of the Swiss Alps (Magnin et al. 2020), by 2.8 to 3.7 Mm³ when compared to the measured bed topography by Vincent et al. (2019). Whereas, it overestimated bed topography by ~14% of volume in the Andes (Drenkhan et al. 2018). When compared with the Ground Penetrating Radar (GPR) readings or existing glacial lakes most of the ice thickness models accurately predict the location of known overdeepenings but the size and morphometry of the overdeepenings are not accurate (Zhang et al. 2019; Gharehchahi et al. 2020; Furian et al. 2021; Otto et al. 2022). More

generally, individual ice thickness estimates have been shown to vary considerably (10±24% from mean ice thickness), with the range among 17 different models being comparable to the magnitude of the ice thickness itself (Farinotti et al., 2017). Understanding different approaches for the models used in this study is therefore likely to be important for understanding any differences in their outputs.

In summarising each of the approaches below, the following symbols are used for the parameters within each of the models, with any deviation from this notation mentioned in the text.

 ρ : density of ice (kg m⁻³)

```
g : acceleration due to gravity (ms<sup>-2</sup>)
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f : shape factor

```
\tau : basal shear stress (Nm<sup>-2</sup>)
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n : Glen's flow law exponent (considered as 3 by default)

2.3.1. GlabTop (Linsbauer et al. 2012)

GlabTop follows the principle as given by Cuffey and Patterson, (2010); given the ice is assumed to be perfectly plastic, basal sliding is not considered, and the ice body is assumed to have almost 10 times greater extent than the vertical depth of the ice body. Then the ice thickness can be calculated assuming it depends on the slope and the basal shear stress using the following equation:

$$d = \frac{1}{(\rho g f sin \alpha)}$$
 Equ(1)

Where d is the ice thickness, ρ is set to 900 kgm⁻³, g is taken as 9.81 ms⁻² and f is set at 0.8 (Patterson, 1998). The value of τ is calculated using Haeberli et al., (1989) quadratic regression to all data points:

$$\tau = 0.005 + 1.598 \bigtriangleup h - 0.435 \bigtriangleup h^2$$
 Equ (2)

 τ is calculated for each grid cell of a glacier and smoothed using an appropriate kernel size of 25m. The values of τ are stored in a separate grid which is later used for calculating ice thickness. Δh is the elevation range. The model requires three inputs: a DEM, glacier outlines and centrelines that are manually digitized. The digitised branch lines are then converted to 100 m raster cells, and divided into sectors of 50 m elevation range. The slope is calculated and averaged for these sectors and this mean value is assigned to the centroid cell of the sector. The values of τ for the glacier and the mean slope value thus calculated are used for estimating the ice thickness for the centroid cell. The centroid cell is then converted into a vector (a point feature) that forms the base point for the calculation of ice thickness. Inverse Distance Weighting (IDW) is used to calculate the ice thickness for the cells between the base points on the branch lines of different sectors. These values are then interpolated using the topogrid to estimate the ice thickness for the continuous grid of the glacier.

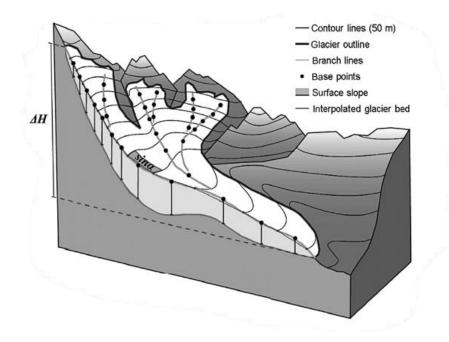


Figure 2.4: Parameters used for ice thickness estimation (Paul and Linsbauer, 2012)

2.3.2. Huss and Farinotti's (HF) Model (Huss and Farinotti, 2012)

The Huss and Farinotti approach require glacier outlines and a DEM as the primary inputs. First, the glacier outline is used to mask the DEM and derive mean slope $\tilde{\alpha}_i$ and area S_i for a given elevation band (i) of 10 m vertical spacing (dz = 10 m) for a given glacier. The width (w_i) of the elevation band is calculated using $w_i = S_i/I_i = \text{Si}$. tan ($\tilde{\alpha}_i$)/dz, where I_i is the elevation band's horizontally projected length. This way the 3-dimensional glacier as seen in Figure 2.4, is converted into a 2-dimensional (2D) shape and the calculations from hereon make use of the values from the 2D shape.

The apparent mass balance <u>b</u> of the glacier is calculated using the ice flux divergence, surface mass balance <u>b</u> and elevation change $\partial h/\partial t$ using the following equation:

$$\underline{b} = \overline{b} - \rho. \ \partial h / \partial t$$
 Equ (3)

The apparent mass balance is determined using accumulation and ablation area altitudinal gradients $d\underline{b}/dz_{acc}$ and $d\underline{b}/dz_{abl}$, respectively. These gradients are linearly related to each other by a constant factor $f_{db/dz} = 0.55$. Since maritime and continental glaciers have different gradients, the model defines a continentality index (C):

$$C = (\frac{ELA_{lat} - ELA_g}{f_{cont}}) + 1 \quad Equ(4)$$

Where f_{cont} = 2400 m is a constant parameter, ELAlat is latitude-dependent reference ELA and ELA_g is median altitude. After this C is used to adjust the apparent mass balance gradient to $d\underline{b}/dz_0$ local conditions:

$$d\underline{b}/dz = d\underline{b}/dz_{\circ}$$
. C Equ(5)

Once the apparent mass balance is estimated using the gradients and balanced mass budget, the ice volume flux for elevation band i of the glacier is obtained by integrating the apparent mass balance along the glacier. After this, Glen's flow law is used to obtain the ice thickness. It is assumed that some of the deformation in ice can be due to basal sliding and it is taken into consideration using the equation:

$$f_{sl} = c_1 + c_2$$
. (S - 10) + c3. (C - 1) Equ(6)

Where c1=0.2, c2=0.1 and c3=4 are constants, S is the glacier area (km²) and C is the continentality index from equation 4. The ice thickness for each band elevation band is calculated using the Glen's flow law (Glen, 1955):

$$h_{i} = \sqrt{\frac{(1 - f_{sl}) \cdot q_{i}}{2A_{f}(T)}} \cdot \frac{n + 2}{(F_{s,i}\rho gsin\tilde{\alpha}_{i})^{n}} \quad Equ(7)$$

Where f_{sl} is a basal sliding factor, q_i is the normalised ice flux, $F_{s, i} = w_i/(2h_i + w_i)$ is a valley shape factor (Nye, 1965). A_f is the flow law's rate factor that depends on the temperature of the glacier.

For temperate glaciers, A_f is assumed to be 0.075 bar⁻³ yr⁻¹. For non-temperate glaciers, A_f is reduced with constant factor $f_{A(T)} = 7^{\circ}C$ as:

$$A_{\rm f}(T) = A_{\rm f}(T=0). \frac{f_{\rm A(T)}}{f_{\rm A(T)} - T}$$
 Equ(8)

In the final step, the elevation band's ice thickness is extrapolated from the simplified 2-D shape of the glacier to a regular grid by using IDW.

2.3.3. GlabTop2 (Frey et al. 2014):

This model works on the same principles as GlabTop but is automated and does not require manual input of flowlines.

DEM and glacier masks are inputs to estimate the glacier's ice thickness. Parameters from volume area scaling and slope-dependent thickness estimation are used to calculate the ice thickness for a glacier complex and are pixel-based. The glacier is divided into different cells (pixels) for ice thickness estimation. First, the glaciers that share a margin are assigned a combined ID. The process is implemented for cells in one ID, setting aside the glacier cells of different IDs. Cells within the glacier boundary are termed glacier cells, whereas those outside the boundary are called non-glacier cells. Non-glacier cells adjacent to glacier cells are referred to as glacier-adjacent cells. Glacier cells at the boundary edge are identified as glacier marginal cells. The glacier cells inside the glaciers and next to the glacier marginal cells are called the inner cells.

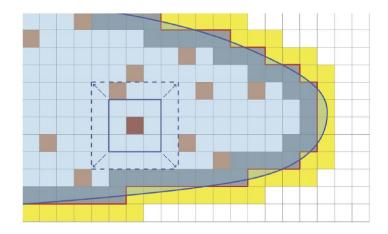


Figure 2.5 : GlapTop2 Workflow: Blue outline: the outline for glacier polygon; Red outline: the outline for glacier cells; Light blue cells: inner glacier cells; Powder blue cells: glacier marginal cells; Yellow: Glacier adjacent cells; White cells: non–glacier cells. Dashed blue square: randomly selected inner cells (Frey et al., 2014)

It is essential to set some parameters for GlabTop2 (t, f, h_{min} , r, h_{ga} and n) for the calculation of the ice thickness. τ is calculated using the following equation from Haeberli and Hoelzle, (1995) using the vertical extent ΔH :

$$\tau[kPa] = \begin{cases} 0.5 + 159.8 \,\Delta H - 43.5 \,(\Delta H)^2 &: \Delta H \le 1.6 \,km \\ 150 &: \Delta H \ge 1.6 \,km \end{cases}$$
Equ(9)

Where ΔH (km) is the elevation range of the glaciers. f is taken as 0.8 for all the glaciers. h_{min} is the minimum elevation, till which the buffer should extend. h_{ga} is the minimum ice thickness for the marginal cells of very narrow glaciers. The parameters n, r and h_{ga} are calculated by considering an idealised valley cross-section in a previously glaciated terrain. The values for n are the number of model runs and r is the percentage of inner cells that are set to fit best for all the glaciers. While n does not have much influence over the total ice thickness of glaciers, the value of r is important for the ice thickness's interpolation from marginal cells and random cells to all grid cells of the glaciers.

Random cells are selected from inner glacier cells and a buffer of 3x3 is laid around each cell, this buffer is then extended until the cell that has a difference in DEM equal to or greater than h_{min} , as seen in Figure 2.5. The mean surface slope α of the selected cells is then used in the following formula to calculate the ice thickness h_{f} :

$$h_{f} = \frac{\tau}{f \rho \sin(\alpha)}$$
 Equ(10)

The IDW method is used to interpolate the ice thickness generated for the random cells to all the cells in the glacier. This method is repeated n times. The ice thickness for each cell is then calculated n times and the mean of these calculations is used as the final ice thickness estimate.

2.3.4. HIGTHIM (Himalayan Glacier Thickness Mapper; Gantayat et al. 2017)

HIGTHIM is a semi-automated tool that can be used to calculate the ice thickness. It requires 6 inputs; glacier outline, glacier flowlines, surface velocity of glaciers, moraines outline, DEM and contour polygons. The glacier outlines can be either delineated manually or can be downloaded from open sources. The flowlines are constructed such that the difference between two adjacent flowlines is ~300 m and ~150 m distance of the flowline from the glacier's boundary. Surface velocity is calculated using LANDSAT images with minimal snow cover and the panchromatic band of 15m resolution. The moraines can be delineated using the LANDSAT imagery and False Colour Composite. The contour polygons are made at an interval of 100 m using the DEM.

The tool uses the laminar flow equation keeping Cuffey and Paterson's (2010) method as the base for ice thickness estimation:

$$H = \sqrt[4]{\frac{1.5 U_s}{A f^3 (\rho g sin \alpha)^3}}$$
 Equ(11)

Where the parameters are set for ρ as 900 kg/m³, acceleration due to gravity g as 9.8 ms⁻², creep parameter A as 3.24 x 10⁻²⁴ Pa⁻³ s⁻¹ (Cuffey and Paterson, 2010) shape factor f as 0.8 which is the standard value for valley glaciers (Haeberli and Holzle, 1995; Cuffey and Paterson, 2010; Linsbauer et al. 2012), and the surface slope angle a is determined across 100 m elevation contours.

The glacier polygon is split into different zones of 100 m each. The length of the flowline is determined for each of these zones. The slope of the contours is calculated using the length of the flowline for the individual zone. The surface velocity of the glaciers is used to extract the velocity for the flowline of the glaciers. This velocity is then used to calculate the ice thickness of pixels along the flowline using the laminar flow equation. The ice thickness for the rest of the glacier is calculated by interpolating the thickness from the flowline using thin plate spline interpolation.

2.3.5. Volume and Topography Estimation (VOLTA; James and Carrivick, 2016)

This model requires only a DEM and a glacier outline as input for ice thickness estimation. First, VOLTA uses a minimum bounding geometry (MBG) which creates a polygon around the glacier.

Following this, a glacier axis is created by joining the two farthest points of the glacier, within the polygon (Figure 2.6(a)). Perpendicular traverses are created on the glacier axis at the DEM's resolution (Figure 2.6(b)). These lines are then clipped to the glacier outline. After this, midpoints are placed on the clipped perpendicular traverses. These midpoints are then joined to construct a centreline, which is smoothed using the Polynomial Approximation with Exponential Kernel (PAEK) algorithm (Fig 3(c)). Glacier area controls the amount of smoothing, as per Kienholz et al., (2014):

$$I = \begin{cases} 2.10^{-6}.A + 200 & : l \le l \max \\ l \max & : l > l \max \end{cases}$$
 Equ(12)

Where A is the glacier area (m^2) and l max is 1000 m.

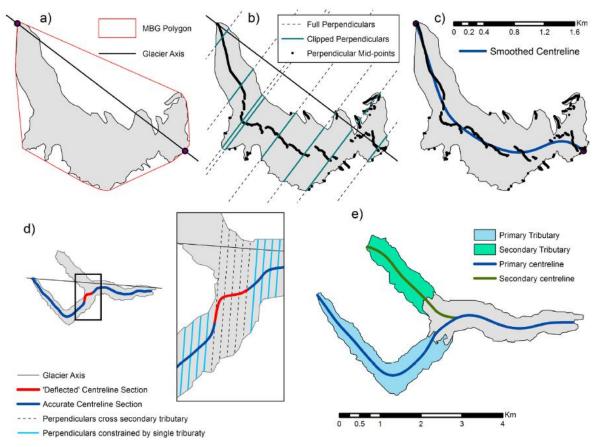


Figure 2.6: Method of centreline production in VOLTA (James and Carrivick, 2016)

For more complex glaciers, which have multiple tributaries and therefore more than one centreline, VOLTA follows the calculation of the upstream area at an interval of 1% of the total length of the centreline. The total upstream area increases down the centreline. If the increase in area between two successive points of the centreline is greater than 30%, a new tributary is identified there. Also, the newly identified tributary should have an area of greater than 20% of the total area. Once, the new tributary is identified, the centrelines are created in the same way as above (Figure 2.6(d) and

2.6(e)). Any perpendicular traverses of the previous tributary are not considered for midpoints. The initial centreline is considered as the primary, following the branch order by Kienholz et al. (2014).

VOLTA uses the following equation to calculate the ice thickness along the centrelines:

$$h = \frac{\tau_b}{f \rho g \tan \alpha} \qquad \qquad Equ(13)$$

A shape factor f is required for valley glaciers. VOLTA uses the approach given by Li et al. (2012) to set the shape factor dynamically. The only difference is VOLTA estimates the vertical ice thickness perpendicular to the horizontal x-axis,

$$h = \frac{0.9 w (\frac{\tau_b}{\rho g \tan \alpha})}{0.9 w - (\frac{\tau_b}{\rho g \tan \alpha})}$$
 Equ(14)

VOLTA has conditions for checking for any inaccurate values of f. Once these conditions are met, the ice thickness for those points is calculated using Equ(14). If the conditions are not met then the ice thickness is calculated using the average shape factor for the centreline and Eq(13). Then VOLTA interpolates the ice thickness using the ANUDEM 5.3 (Hutchinson et al. 2011) through the Topo to Raster tool, using ice thickness and glacier outlines as input. Basal shear stress is found using the formula:

$$\tau_{b} = 2.7 . 10^{4} \sum_{i=1}^{n} \left(\frac{A_{i}}{\cos \alpha_{i}}\right)^{0.106}$$
 Equ(15)

Where A_i is the elevation band area and $\cos \alpha_i$ is calculated over 200 m elevation bands.

Slope averaging distance α_d is the distance through which the centreline gradient is analysed. It is set to 10 times the average thickness \hat{h} , which is calculated using the volume area scaling method where A is the glacier area:

$$\hat{h} = \frac{0.2055A^{1.375}}{A}$$
 Equ(16)

A threshold is defined for minimum slope α_0 as otherwise the ice thickness may be overestimated for flatter regions of the glacier. It is set to 4° for VOLTA.

2.3.6. Open Global Glacier Model (OGGM; Maussion et al. 2019):

OGGM is a relatively recent ice thickness model. It works on a mass conservation approach. Firstly, gridded climate data for monthly precipitation and temperature are used to calculate the mass balance of the glacier using the formula:

$$m_i(z) = p_f P_i^{\text{Solid}}(z) - \mu^* \max (T_i(z) - T_{\text{Melt}}, 0) + \varepsilon \quad \text{Equ}(17)$$

where m_i is the monthly mass balance at an elevation z, p_f is a global precipitation correction factor, P_i^{Solid} is the monthly solid precipitation and is calculated as a fraction of the global precipitation ranging from 100% to 0%, T_i is the monthly air temperature, T_{Melt} is the monthly air temperature above which ice melt is assumed to occur, and ε is a bias correction term. T_{Melt} is set to -1°C and p_f is set to 2.5. μ^* is the glacier's temperature sensitivity, and it is calibrated using the approach by Marzeion et al. (2012). μ is calibrated across several years and it is assumed that the mass balance across these years sums to zero. Hence, the mass balance is taken as zero for 31 years and $\mu(t)$ is calculated. This $\mu(t)$ is then used to calculate the bias $\varepsilon(t)$. This calculation is done for more than one given year. The time with the least bias value is taken as the time for calculation and the value of $\mu(t)$ for this time is taken as μ^* . This value of t* is then interpolated for different glaciers which do not have observations using the inverse distance interpolation.

The flowlines for the model are constructed using the algorithm provided by Kienholz et al., (2014) (Figure 2.7(b)). Initially, centrelines are created which are then converted into flowlines. There is a main branch line and the rest of the centrelines are tributaries, the tails of the tributaries are cut before they join their respective descendants (Figure 2.7(b) and 2.7(c)). These flowlines are ordered using the Strahler, (1957) ranking system. This order helps in mass flow routing.

The catchment area for the centrelines is defined using the algorithm used to define the centrelines. Intersecting the normals to the flowlines with the catchments or the glacier's boundary provides estimates for the geometrical width (Figure 2.7(c)). These geometrical widths are then corrected by a factor to preserve the altitudinal distribution of the glacier (Figure 2.7(d)). To keep the total area of the glacier the same, these widths are then multiplied by another factor, specific to the altitudinal unit, which ensures that the area of the glacier remains the same as provided in the RGI dataset.

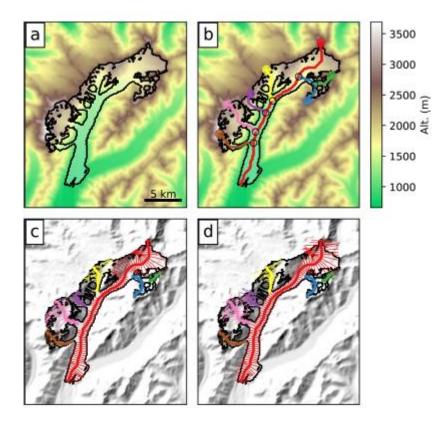


Figure 2.7 : Workflow of OGGM Model for New Zealand's Tasman Glacier; a) preprocessing of topographical data b) centrelines c) geometric width determination for the glaciers d) geometric width correction for the glaciers (Maussion et al., 2019)

The mass balance obtained from the above equation, the flowlines computed from the centrelines using an algorithm developed by Kienholz et al. (2014) and the geometrical widths along them, calculated from the catchment areas are used in the ice thickness inversion method. The method uses the following formula for ice flux:

$$q = uS$$
 Equ(18)

Where q is the ice flux (m³s⁻¹), u is the average ice velocity (ms⁻¹), and S is the cross-section of the area. S varies according to the shape of the glacier bed; for a parabolic bed, it is approximately 2/3 of the area of a rectangular bed. Shallow ice approximation is used to calculate the average ice velocity:

$$u = \frac{2A}{n+2} h\tau^n \qquad \qquad \text{Equ (19)}$$

Where *h* is the ice thickness, *A* is the ice creep parameter (s⁻¹Pa⁻³), τ^n is the basal shear stress and is calculated using the equation:

$$\tau = \rho g h \alpha$$
 Equ(20)

26

Where ρ is set to 900 kg m⁻³, g is set to 9.81 ms⁻² and α is the slope calculated along the flow line. Sliding velocity can be added to the equation using the following formula:

$$u_s = \frac{f_s \tau^n}{h}$$
 Equ(21)

Mass balance is then used in the equation of mass conservation where the glaciers are assumed to be in a steady state:

$$q = \int_{\Omega}^{n} \left(\dot{\mathbf{m}} - \rho \frac{\partial n}{\partial t} \right) d\mathbf{A} = \int_{\Omega}^{m} \dot{\mathbf{m}} d\mathbf{A} \qquad \text{Equ(22)}$$

 \dot{m} is the mass balance and \dot{m} is the apparent mass balance at a point in the glacier. For this model, it is assumed that the glacier is in a steady state and $\dot{m} = \dot{m}$. This mass balance is then used with the bed topography, flowline, slope and width:

$$\frac{\partial s}{\partial t} = w \dot{m} - \nabla . u \qquad \text{Equ(23)}$$

This equation is solved using a formula that ensures numerical stability and does not include h ($\Delta t = \gamma \frac{\Delta x}{\max(u)}$), where γ is a dimensionless courant number chosen between 0 and 1. This formula is governed by the shape of the glacier bed, unlike other diffusivity formulas which depend on the value of h and overlook the geometry of the bed. *S* at *t* + Δ *t* is used to calculate h (*t* + Δ *t*) for local bed geometry. This way h can be calculated for points along the flowline, which is later interpolated for the entire glacier.

2.3.7. Farinotti's Ensemble (Farinotti et al. 2019)

Farinotti's ensemble used five models to estimate the ice thickness of the glaciers. RGI version 6 was used for the outlines of the glaciers. Different DEMs were used for ice thickness models, based on the location of the glaciers. To estimate the final thickness, the ice thickness models were weighted according to inverse variance and bias. A composite solution û for the ice thickness h_i is computed assuming that the ice thickness models are independent. û is calculated as:

$$\hat{u} = \frac{\sum h_i w_i^{-1}}{\sum w_i^{-1}}$$
 Equ(24)

Where w_i is a weighting that represents both the variance σ_i^2 and the bias b_i . b_i is taken as an absolute value penalising both overestimation and underestimation. One-third of the data for ice thickness was used for cross-validation, to estimate the bias and the variance. The remaining two-

thirds of the data was used for model validation. The cross-validation experiment was performed thrice, and the deviations between modelled and measured ice thicknesses were stated relative to the mean ice thickness. The mean and interquartile range of the pooled deviations (from the three experiments) were estimated for b_i / h_i and 1.5 σ/h_i , respectively. The variance \hat{u} is given by:

$$\sigma_{\hat{u}}^2 = \frac{1}{\sum \sigma_i^{-2}} \qquad \qquad \text{Equ(25)}$$

The sensitivity of the weighting strategy was calculated by combining the results of the crossvalidation experiment for separate RGI regions. The results are not sensitive to this scheme.

2.3.8. Millan's ice thickness (Millan et al. 2022)

Millan et al. (2022) have mapped 98% of the world's glaciers for ice thickness. Satellite images were orthorectified and used to map the glacier velocity using the method described in Millan et al. (2019). This velocity was calculated for each mosaicked glacier of each sub-region. The outliers from the calculated velocity are removed and a Gaussian smoothing is performed for spatial adaptive filtering. This velocity is used in Shallow Ice approximation, where basal sliding is taken into consideration to calculate the ice thickness using the following equation:

$$v_{s} = v_{b} - 2 (\rho g)^{n} \parallel \nabla s \parallel^{n-1} A \frac{H^{n+1}}{n+1} \nabla s \qquad Equ(26)$$

Where v_s is surface velocity, v_b is basal velocity, ρ is set to 917 kg m⁻³, g is set to 9.81 ms⁻², n is taken as 3, s is the ice surface elevation (∇ s being the surface slope) and A is the creep parameter. The slope map of the entire region and the velocity were used to calculate the ice thickness:

$$H = \left(\frac{(v_s - v_b)(n+1)}{2A (\rho g)^n \|\nabla s\|^n}\right)^{\frac{1}{n+1}} \qquad \text{Equ(27)}$$

Since, the basal velocity is not known, a ratio between basal and surface velocity B is used where $v_b = \beta v_s$. This changes the above equation into the following equation:

$$H = \left(\frac{v_{s}(1-\beta)(n+1)}{2A(\rho g)^{n} \|\nabla s\|^{n}}\right)^{\frac{1}{n+1}}$$
 Equ(28)

A and β from the above equation needs to be calibrated. A is calibrated using the following equation:

$$\mathsf{A} = \left(\frac{v_{s}\left(1-\beta\right)\left(n+1\right)}{2(\rho g)^{n} \|\nabla s\|^{n} H^{n+1}}\right) \qquad \mathsf{Equ}(29)$$

2.3.9. Section Summary

Most of the models used in previous studies were included in the ITMIX calculation by (Farinotti et al., 2017) and can be broadly grouped into 4 categories based on their approach to estimating the ice thickness. These are:

- 1. Mass conservation approaches: HF, OGGM, VOLTA
- 2. Shear stress-based approaches: GlabTop, GlabTop2
- 3. Velocity-based approaches: Millan's ice thickness, HIGTHIM
- 4. Minimization approaches: Ensemble

Nb. Ensemble can fall into all the categories because it uses at least one model from the categories but it has been classified as a minimization approach here which is the main principle behind it

Though these approaches might seem different, several of them follow similar principles. For example, the shear stress-based approach does use velocity in its calculation even though it is not classified as a velocity-based approach. Parameterization becomes important for defining the differences in the ice thickness models. Setting parameters is one of the initial steps for ice thickness calculation. The most common parameters used are the basal shear stress, slope of the glaciers, shape factor, acceleration due to gravity, Glen's flow law constant and density of ice. Among these, definitions of basal shear stress and slope of the glaciers vary for different models. The shape factor is set as 0.8 for GlabTop, GlabTop2 and HIGTHIM, whereas for OGGM, VOLTA and HF model, it is calculated. These parameters are then applied with the required equation for further calculations.

The primary inputs required by all the models include a DEM and glacier outlines, though the specific sources might be different for each model. Apart from that, velocity is also a frequently required input for the models. Some models like HIGTHIM and Millan's ice thickness model calculate the velocity from satellite images by incorporating them in mathematical formulae whereas OGGM does not calculate the velocity though it uses mathematical formulae to estimate the ice thickness.

Over time the complexity of ice thickness models has increased leading to a greater variety of required inputs. Earlier models like the HF- model required only DEM and a glacier outline. In contrast, recent models like the OGGM use climate data, catchment area width and flowlines along with DEM and outline. While flowlines are not a required input for all the models, they still play an important role in determining the ice thickness. All models follow Kienholz et al. (2014) for the construction of flowlines though the method used for constructing them can be different. Models like GlabTop and HIGTHIM require manual construction of the flowlines, whereas VOLTA and OGGM can generate them based on DEM and the glacier outline provided.

The glacier flowline is also used as a base unit for ice thickness estimation. The base unit can be considered the smallest part of the glacier for which the ice thickness is calculated and later interpolated to generate the ice thickness for the entire glacier. HF Model has an elevation band as a unique base unit when compared to other models that use cells (pixels) as their base unit. These cells are along the flowlines for most of the models. The interpolation technique that is commonly used is IDW. VOLTA and HIGTHIM use spline for interpolation and Millan's ice thickness model has used kriging specifically for Antarctica.

Millan's model, OGGM and Farinotti's Ensemble have estimated the ice thickness of all the glaciers worldwide. HF model has estimated the ice thickness of 300 glaciers worldwide. These ice thickness outputs can be further used as foundational data for various glaciological studies. The parameters, methods, equations, and inputs differ not only between models but also within the same model across various environmental settings. As a result, analyses based on these ice thicknesses are likely to vary.

2.4. IMPORTANCE OF PREDICTING FUTURE OVERDEEPENING POSITIONS

While glacial lakes play a crucial role in glacier dynamics, they also pose hazards. In regions like the Himalayas and Peruvian Andes, where glacial hazards are common, early detection and monitoring of overdeepenings can enhance preparedness and mitigation efforts, particularly for hazards such as GLOFs (Sattar et al. 2019). Globally, the number and size of glacial lakes have been increasing (Carrivick et al. 2022; King et al. 2020; Majeed et al. 2021a), which is a trend that is expected to continue in future years (Furian et al. 2025).

There have been many glacial lake outburst floods in recent years. The recent GLOF in August 2024, due to a break in the dam of Thynabo Lake in Nepal, swept away the major infrastructure of Thame

village and displaced at least 135 people (Khatri, 2025). Such events underscore the importance of identifying and assessing hazard potential in downstream locations. Given that glacial lakes are often unstable during their early formation stages (Emmer et al. 2015b), it is crucial to predict and monitor the evolution of overdeepenings to anticipate future hazards.

Overdeepenings play a central role in predicting the formation of future lakes, though not all overdeepenings transform into lakes (Fountain and Walder, 1998; Evans, 2008). Nevertheless, they remain important because they influence sedimentation patterns and contribute to landscape evolution. Advanced detection techniques, particularly remote sensing, have proven invaluable in identifying overdeepenings in inaccessible regions (Haeberli et al. 2016; Magnin et al. 2020). This technological progress enhances our ability to manage GLOF risks and contributes to long-term hazard reduction strategies (Drenkhan et al. 2018, 2019; Sattar et al. 2019).

2.5. SUMMARY

Current and future overdeepenings have become a popular focus in the glaciological literature, mostly on a regional basis in the valley glacier setting, and in mountain regions where communities live within the downstream areas (Drenkhan et al. 2019). While overdeepenings in the ice sheet environment is crucial for understanding broader glaciological processes and the delivery of sediments to the oceans (Carrivick et al. 2022), those in valley glacier settings demand particular attention due to their direct implications for human safety and hazard management. Frey et al., (2010) first explained how ice thickness, glacier outline and DEM can be used for estimating overdeepenings. This method has since been used very broadly.

Despite the easy availability of DEMs and ice thickness data, most studies use either one ice thickness model and one DEM or at most two ice thickness models or two DEMs. The overdeepenings produced from just one combination of a DEM and ice thickness model are then used as a base to gain future perspectives on a hazard assessment (Drenkhan et al., 2019; Furian et al., 2025). Given that existing datasets often lack precision in determining the morphometry of overdeepenings, relying on them without accounting for their inherent uncertainties can amplify errors in future predictions. Accurate morphometry is crucial for assessing the potential downstream impacts of overdeepenings, particularly if they evolve into glacial lakes capable of triggering GLOFs. Therefore, addressing these uncertainties is essential for reliable hazard assessment and management.

The wide variety of datasets available for glaciological studies has led to inconsistencies in the data used across individual studies. This lack of standardization makes it challenging to compare results between different regions, even when the glaciers are of similar type. Studies that use multiple datasets often yield varying results, highlighting the inconsistencies from using different data sources. Given these variations and the availability of various DEMs and ice thickness models, a comparative study incorporating multiple datasets could offer valuable insights. Such an analysis would help clarify how different models and DEMs perform relative to each other, ultimately improving our understanding of their reliability and the potential biases they introduce. This is the primary motivation for the current study.

3. AIMS AND OBJECTIVES

Based on the literature review, it can be seen that only a few studies have used more than one DEM or one ice thickness model to detect subglacial overdeepenings. Those studies that do use more than a single approach are still limited by the fact they only implement their process within a single region. To date, there has not been an intercomparison between the models and the DEMs to assess whether there is regional variability among the models.

Considering the findings from the previous studies, the aims of this study are, therefore, to:

- 1. Determine the consistency of different DEM and ice thickness model combinations in predicting the depth, volume and area of subglacial overdeepenings.
- 2. Determine the extent to which the outputs from different ice thickness models vary across different regional settings.
- 3. Propose a method that can account for the uncertainties produced by different ice thickness models.

These aims will be met by addressing the following objectives:

- Generate quantitative estimations of overdeepening location, depth, area and volume using different combinations of DEM and ice thickness products for a single test area (Central Himalayas).
- 2. Investigate how variations in overdeepening predictions manifest across different mountain regions (in different climatic and topographic settings) to explore regional trends and inconsistencies.
- 3. Explore the effectiveness of combining outputs from multiple ice thickness models to improve the robustness of overdeepening assessments and reduce uncertainties.

4. METHODS

4.1. STUDY AREAS

Globally glaciers from five regions were chosen as study area, see Figure 4.1 and Figure 4.2 for reference. These regions were the Himalayas, Peruvian Andes, Alaska, New Zealand Southern Alps and the European Alps.

4.1.1. Himalayas

The Himalayas range from Namche Barwa in the east to Arkan Yoma in the West spanning over 2400 kilometres across five countries: India, China, Nepal, Pakistan and Bhutan. They are a major water source for Asian perennial rivers like Indus, Ganges and Brahmaputra. A recent glacier inventory for the Himalayan Karakoram glaciers shows that there are around 21,000 glaciers (Cogley, 2016). In the Himalayas, these glaciers cover an area of around 23,000 km², which has a volume of 1,071 km³ (Bolch et al. 2012). Although the rate of glacial retreat varies across different parts of the Himalayas, studies have shown that approximately 13% of a mapped 11,000 km² of the glaciated area has been lost due to glacier retreat (Kulkarni and Karyakarte, 2014). As these glaciers retreat, they expose new areas of the landscape leading to the formation of a large number of glacial lakes and increasing the risk of natural disasters such as GLOFs. This threat to mountain communities is additional to that posed by an increasingly uncertain water supply as glaciers reach and pass peak water (Huss and Hock, 2018), where temperature and precipitation patterns become more erratic and more extreme (Regmi and Bookhagen, 2022).

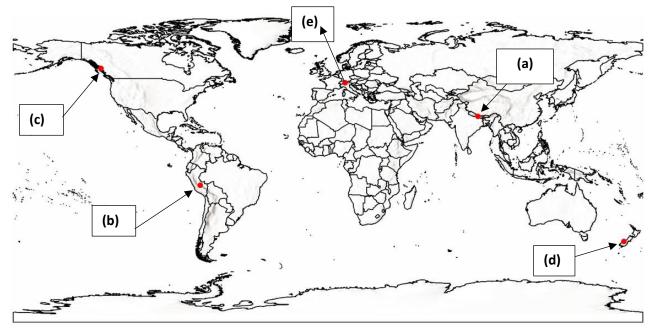


Figure 4.1: Study regions in the world map. Note the red dots represent the study region, and the letters refer to panels in Figure 4.2

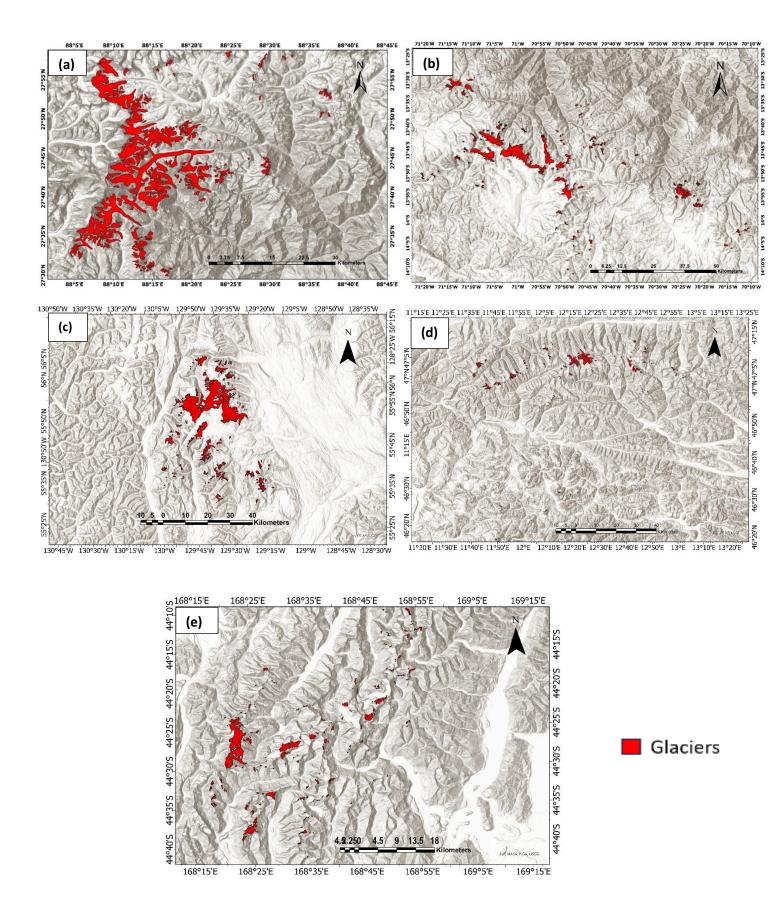


Figure 4.2: Selected glaciers in the (a): Himalayas (b) Peruvian Andes (c) Alaska (d) European Alps (e) New Zealand Southern Alps

4.1.2. Andes

The Andes are home to over 99% of the world's tropical glaciers (Kaser, 1999). These glaciers are distributed across several countries including Bolivia, Peru, Ecuador and Colombia-Venezuela and are found at elevation ranges from 4,000 m and, to 6,500 m in altitude. Collectively, they cover a total area of 2,560 km² (Rabatel et al. 2013). These tropical glaciers are particularly sensitive to changes in moisture and precipitation and act as early indicators of climate change (Jomelli et al. 2009). Glaciers in the inner tropics will be more sensitive to warming in the future (Vuille et al. 2018). Similar to the glaciers present worldwide, at present they are also undergoing retreat since the Little Ice Age. Glaciers in the Cordillera Blanca region have decreased in area by 46% from 1930 to 2016 (Torres and Jaquet, 2017) and in the Cordillera Vilcanota they have decreased in area by 54% from 1975 to 2020 (Taylor et al. 2022). This ongoing retreat poses risks, as it can increase the likelihood of various glacial hazards, including avalanches, rockfalls, and GLOFs (Chevallier et al. 2011).

4.1.3. Alaska

There were around 100,000 glaciers in Alaska that covered 74,705 km² (equal to 4.33% total area of present-day Alaska) in 1980 (Meier et al. 1985). Recently the glacier retreat has increased in Alaska, with it projected to contribute the highest amount to sea level rise from 2015 to 2100 (Rounce et al. 2023). An example can be seen in glaciers like Barry Glacier where glacier thinning has accelerated in recent years, from 2010 to 2017 (Dai et al. 2020). It has resulted in a horizontal slope displacement of 120m for the glacier.

The mass balance of glaciers here is complicated as there are glaciers that show retreat and there are few glaciers that show advance (Alifu et al. 2016; Johnson et al. 2013), like the Hubbard Glacier. Glaciers in Alaska, and the West Fork Glacier, have experienced surging in the past. Between 1987 and 1988, the West Fork Glacier displaced approximately 4 km of ice by surging (Harrison et al. 2003).

The rising temperatures and changes in precipitation patterns govern the glacial dynamics here. Alaska's glaciers, in particular, have played a substantial role in this process, accounting for a significant portion (0.4m of sea level rise equivalent) of the overall sea level increase during this period (Zemp et al. 2019). Previous work focussing on Alaskan glaciers has focussed on the formation of glacial lakes, an increase in risk due to glacial lake outburst flood and also changes in the nutrient content that can affect the aquatic life downstream (Dorava and Milner, 2000; Moore et al. 2009).

4.1.4. European Alps

The European Alps extend around 1,000 km across Germany, France, Austria, Italy, Slovenia, Switzerland and Liechtenstein. Rather than forming a continuous expanse of ice, the glaciers in the Alps are scattered throughout. The range has a few high summits, with Mont Blanc standing as the tallest peak, reaching an elevation of about 4,800 m (Deline et al. 2012).

These glaciers have lost roughly 0.5 a⁻¹ volume of ice thickness ranging in a span around 12 decades starting from 1850, this loss accelerated sharply after 1975 till 2000 with a change of 25% and further with a change of 10-15% in the early 2000s (Haeberli et al. 2007). As a result, the total glacier surface area fell to just over 2,000 km² by the early 2000s (Sommer et al. 2020).

This rapid decline in the glacier area (Zemp et al. 2006) has far-reaching consequences for the region including disruptions to the water cycle, increased risk of natural hazards and alterations to the landscape like landslides and formation of glacial lakes (Haeberli, 2013; Haeberli et al. 2017).

4.1.5. New Zealand Southern Alps

The New Zealand Southern Alps stretch over approximately 500 kilometres along New Zealand's South Island. In the mid-1970s, the region contained 3,132 glaciers, covering a total area of 1,139 km². Of these, only 702 glaciers were larger than 0.2 km², and they accounted for about 86% of the entire glaciated area according to Hoelzle et al., (2007). There has been a loss of 41% to 62% (of the Little Ice Age (LIA) volume) in the volume of the glaciers since LIA till 2019 (Carrivick et al. 2020). These glaciers are located in a cool temperate climate, characterized by significant variations in precipitation, ranging from about 1.5 meters per year to nearly 12 meters per year (Henderson and Thompson, 1999) across the topographic divide. Among these glaciers, only a small number terminate in lakes (Chinn, 1999). Despite their limited number, these lake-terminating glaciers play an important role in the region's tourism (Purdie, 2013).

The New Zealand Southern Alps experienced a positive mass balance for many glaciers in the 1990s, indicating a period of relative stability and growth. However, recent studies (Anderson et al. 2021; Carrivick et al. 2022) have highlighted a marked shift, with glaciers in the region now experiencing a more pronounced retreat. Along with the retreat, the expansion of ice-marginal lakes can cause changes in the glacial dynamics in the New Zealand Southern Alps (Carrivick et al., 2022).

4.2. DATA SOURCES

4.2.1. Glacier Outlines:

Most previous studies have used Randolph Glacier Inventory version 6 (RGI 6.0, Pfeffer et al., 2014) for estimating overdeepenings and ice thicknesses for valley glaciers worldwide. This study also used RGI 6.0 for estimating the overdeepenings to ensure uniform data processing. RGI is a global inventory of glaciers worldwide (excluding the Greenland and Antarctica ice sheets). It is a subset of the Global Land Ice Measurement from Space (GLIMS) initiative (Raup et al. 2007), which has more attributes than the RGI. RGI glacier outlines were first released in 2012 and have been continuously updated since then.

4.2.2. Ice Thickness Models:

Five ice thickness models were used in this investigation. Three of them (GlabTop2, OGGM and Huss and Farinotti) were selected from those included in Farinotti's Ensemble, with Farinotti's Ensemble itself being chosen for the analysis (Farinotti et al. 2019a). The fifth ice thickness was that of Millan et al. (2022). All ice thickness models require glacier outlines and DEMs as basic inputs to estimate ice thickness. Additional inputs may vary depending on the model. OGGM requires climate data, while the Millan model incorporates glacier velocity data. The ice thickness of OGGM for the Himalayas was estimated manually. The parameters were set to default for the estimation. The resulting ice thickness estimates are available in raster formats, with a resolution of 49.99 m for the Millan model and 30 m for other ice thickness models. Data were downloaded using the following links:

Millan's ice thickness: SEDOO/2022 https://doi.org/10.6096/1007

HF, GlabTop2, OGGM and Farinotti's Ensemble: ETH Zurich/2019 <u>https://doi.org/10.3929/ethz-b-</u> 000315707

4.2.3. Digital Elevation Models (DEMs):

Four DEMs were used in the study (Copernicus, ASTER GDEM, SRTM and ALOS PALSAR), to assess the extent to which the choice of DEM impacted overdeepening location and characterisation. Among these DEMs, Copernicus and ALOS-PALSAR have a global coverage whereas the SRTM and ASTER GDEM have a *nearly* global coverage. SRTM and ASTER GDEM have been used most widely to estimate the ice thickness. Recent studies (Furian et al. 2023) have acknowledged ALOS-PALSAR for being high resolution and it has been used for validation of the overdeepenings estimate. These particular DEMs were chosen to represent sources that were easily available (i.e. open source), have been widely used in other studies (so provide potential comparison with this study), have published accuracies that demonstrate their quality, and do not contain voids that would complicate and/or compromise the validity of the processing chain. Previous studies have indicated that the choice of surface DEM can be important for the identification of overdeepenings (Linsbauer et al. 2012; Magnin et al. 2020), but without assessment across a broad area.

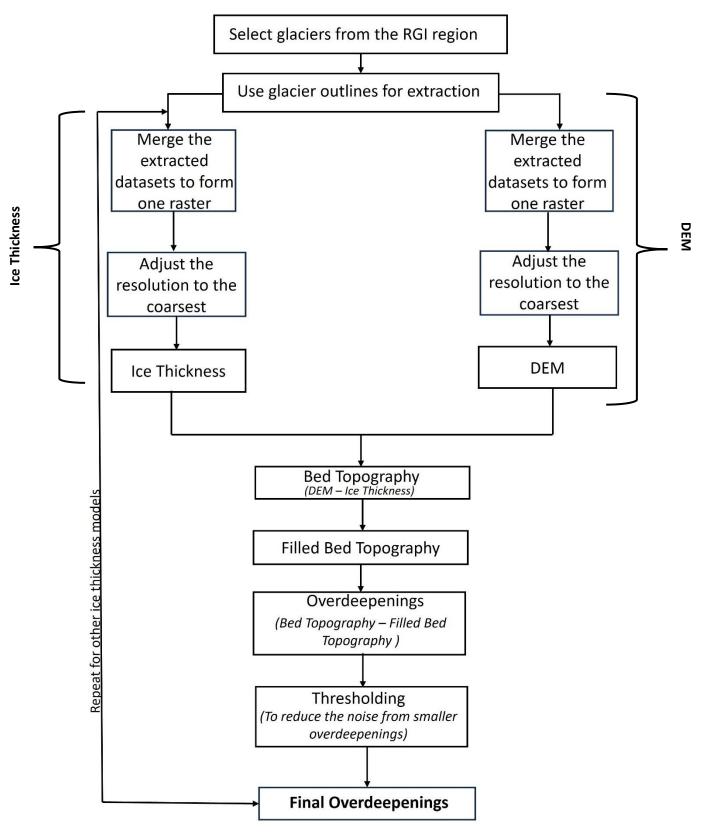


Figure 4.3: Flowchart of the methodology followed

4.3. DATA PREPROCESSING

Figure 4.3 represents the data processing followed in this study.

4.3.1. Selecting glaciers in one RGI region

There are 19 RGI regions for all glaciers worldwide in the inventory. Five regions containing valley glaciers were chosen for analysis in the current study. These regions are heavily populated (and therefore knowledge of overdeepening locations is of practical benefit to society) as well as representing a range of climatic conditions (monsoon, westerlies, tropical), geologies, and topographies.

The Central Himalayas were selected as the initial region for method development before extending the approach to other areas, given its stored water potential (Shruti et al. 2025). The region near the Zemu Glacier (27°45'51.19"N, 88°22'32.57"E) was chosen, as it recently garnered attention due to a major GLOF that occurred in October 2023 (Sattar et al. 2025). This event was triggered by intense rainfall in Sikkim, where the region experienced double its usual precipitation. The excessive rainfall caused the moraine dam of South Lhonak Lake to breach, resulting in a catastrophic GLOF. The flood had devastating downstream effects, claiming at least 93 lives and severely damaging infrastructure, including the Chungthang hydroelectric dam on the Teesta River. The other regions (as described above) were Alaska, New Zealand's Southern Alps, the Peruvian Andes, and the European Alps. In all the regions except for the Himalayas, area breaks were defined using the histograms observed in ArcGIS Pro. For the Himalayas, the subset was chosen randomly with glaciers including and around the Zemu Glacier. In each of these other four regions, area breaks were established, and 200 representative glaciers were chosen from these defined area ranges from the RGI region. These area breaks were determined by the histograms observed in ArcGIS Pro. Area was chosen as the basis to select the representative glaciers, to maintain the heterogeneity in terms of parameters like slope, debris cover, aspect without introducing a bias to any other factor, while also ensuring the sample size of 200 was representative of the entire RGI region (in terms of area breaks). The distribution of glaciers in the entire RGI region and the subset are given in Table 4.1 and Figure 4.5. The distribution of the glaciers in the subset is representative of the distribution in the RGI region (Table 4.1 and Figure 4.4, Figure 4.5), based on the defined area breaks. The median glacier area of the entire RGI region was then compared to that of the subset. As seen in Table 4.2, the median

value for the entire region and the subset is similar for all the regions, indicating that the findings

Table 4.1: Area breaks and distribution of glaciers in the subset and the RGI region

from the analysis could be broadly applicable to all glaciers within that region.

		Number of glaciers	
Region	Area Breaks (km ²)	RGI region	Subset
	0-0.1	5070	38
Alaska	0.1 –1	16357	120
AldSKd	1 –10	4924	36
	>10	757	6
	0-0.1	2181	112
Furancan Alas	0.1 - 1	1339	68
European Alps	1 – 3	271	14
	>3	136	6
	0 - 0.5	1921	116
Andre	0.5 – 0.9	292	20
Andes	0.9 - 1.4	218	13
	>1.4	460	51
	0-0.1	2120	120
New Zealand	0.1 - 0.5	1053	60
Southern Alps	0.5 – 0.9	185	10
	>0.9	179	10

Subsequently, a Wilcoxon test was conducted to more robustly assess the similarity of the distributions. A Wilcoxon test is a non-parametric test to check if two datasets are significantly different.

The p-value for all regions is greater than 0.05 except the Andes, indicating there is a significant similarity in distribution among the subset and the entire region (see Table 4.2). The p-value for Andean glaciers is 0.0266 but excluding the largest glacier from the Andean subset increases the p-value to 0.07, and it still does not affect the median as much (it changes from 0.2765 to 0.2760). Suggesting the distribution is resistant to extreme values even with presence of some bias. Hence, it can be assumed that the inclusion or exclusion of this glacier does not meaningfully alter the central tendency. Therefore, the glacier was retained in the subset, as its presence does not significantly impact the overall analysis.

	Median Value	Wilcoxon Test		
Region	RGI glaciers	Subset glaciers	p-value	
Alaska	0.313	0.311	0.5975	
European Alps	0.077	0.078	0.9782	
Andes	0.218	0.276	0.0266	
New Zealand Southern Alps	0.077	0.066	0.7207	

Table 4.2: Median values for the RGI glaciers and the subset with p-value of the Wilcoxon test

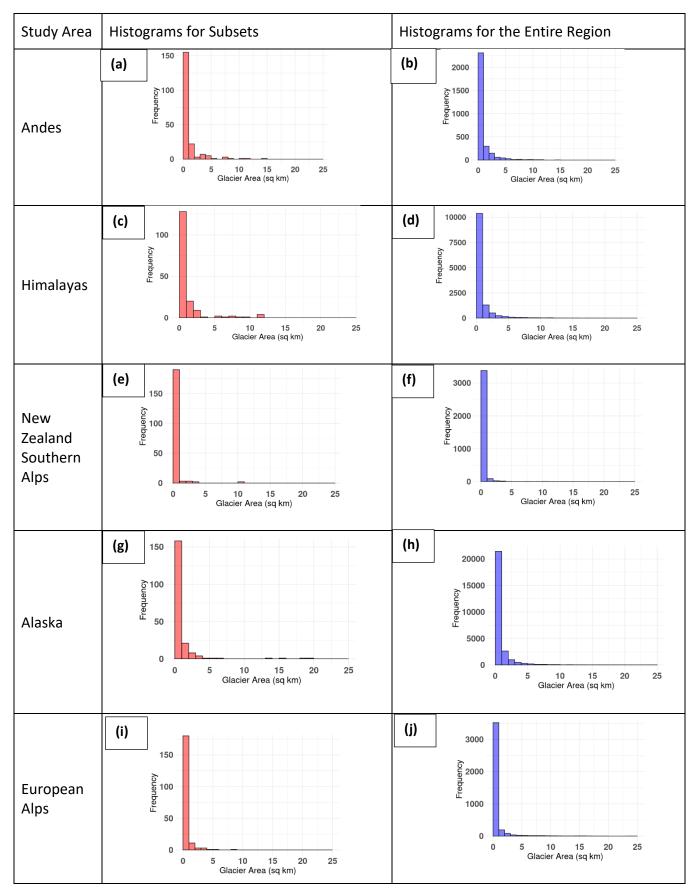


Figure 4.4: Histograms of area distribution of the RGI region glaciers and the selected subsets

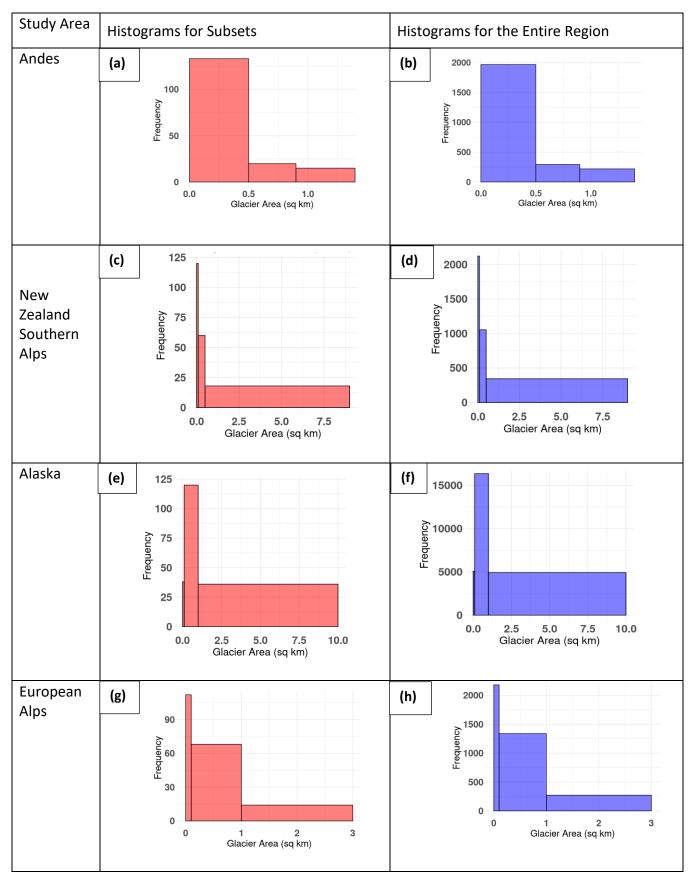


Figure 4.5: Histograms of area distribution of the RGI region glaciers and the selected subsets with defined area breaks

4.3.2. Selecting the ice thickness models

Five ice thickness models were chosen for the analysis. These were:

- *Farinotti's Ensemble (Farinotti et al. 2019a)*: This is one of the latest ice thickness datasets available. It has an input of 5 ice thickness models as an ensemble, suggesting it may be more robust than using any single model in isolation.
- *ii) Huss and Farinotti (Huss and Farinotti, 2012):* This is one of the oldest models and has been used in estimation of ice thickness for various studies. The ice thickness for this model is also available for all glaciated regions.
- *iii) GlabTop2 (Frey et al., 2010):* This is one of the widely used models for estimating overdeepenings in glaciated regions, as seen in Chapter 2, and has been frequently applied to estimate the ice thicknesses of valley glaciers.
- *Open Global Glacier Model (OGGM; Maussion et al. 2019)*: This is one of the recent ice thickness models, developed to simulate glacier evolution rather than ice thickness *per se*. OGGM requires inputs (like the precipitation and temperature data) and the data for this model are also available for all the glaciated regions.
- *v) Millan (Millan et al. 2022*): This is the latest ice-thickness product available for the glaciated regions.

4.3.3. Ice thickness grids were available on an individual glacier basis, except for the Millan dataset where the grid comprised the entire RGI region. In this particular case the thickness estimates for selected glaciers were clipped and merged before inserting into the workflow. Areas of no-data exist in several of the models, in particular around the glacier margins. Where necessary these pixels were removed by masking the rasters with the glacier outlines. The cleaned rasters were then merged to create an ice thickness dataset for the subset of the region.Selecting the DEM

The following DEMs were used in the initial analysis for the Central Himalayas:

i) Advanced Space Borne Thermal Emission and Reflection Radiometer (ASTER) DEM: It is obtained from the ASTER instrument (Fujisada et al. 1998) onboard the National Aeronautics and Space Administration's (NASA) Terra satellite, launched in 1999. It has a spatial resolution of 30 m and covers most of the Earth from latitudes 82° North to 82° South. This is one of the recommended DEMs for studying the Himalayan glaciers (Bhambri and Bolch 2009; Bolch et al. 2012; Ragettli et al. 2016).

- Shuttle Radar Topography Mission (SRTM) DEM: This mission was launched in 2000 as a joint venture of NASA, the National Geospatial-Intelligence Agency (NGA), and the German and Italian Space Agencies. It covers latitudes ranging from 60° North and 56° South (Rabus et al., 2003). The DEM has two resolutions one of 30 m and the other of 90 m. This study has used the resolution of 30 m. Also, it is one of the most widely used DEMs for the Himalayas (Fujita and Nuimura, 2011; Pandey and Venkataraman, 2013; Yao et al. 2012).
- iii) Copernicus DEM: This was programmed and launched by the European Union Earth Observation. This covers data from pole to pole. It is derived from the German Aerospace Center's (DLR) TanDEM-X mission data. This radar interferometry mission used twin satellites (TerraSAR-X and TanDEM-X) to generate a global digital elevation model of Earth's surface from 2010 to 2014. The 30m resolution was used for this study.
- iv) Advanced Land Observing Satellite Phased Array L-band Synthetic Aperture Radar (ALOS PALSAR) DEM: This DEM is generated by the PALSAR instrument onboard the ALOS satellite that covers the global region from 80° North to 80° South. The DEM is available in two main resolutions 12.5m and 30 m. This study has used the resolution of 30 m. This product has been used in several recent studies in the Himalayas (Furian et al. 2021, 2022).

4.3.4. Standardizing the resolution of all the datasets

Given that the DEMs and ice thickness models had differing spatial resolutions (Table 4.3), keeping them at their native resolution may have introduced some uncertainty in the comparisons between the model outputs. ALOS-PALSAR DEM was selected for the second part of the analysis because of its high resolution and worldwide void-free coverage. It was downsampled to 49.99 m for all of the analysis. This also ensured that the method this study follows, can be adapted easily. Therefore, all datasets were resampled to the lowest resolution using the 'Resample Tool' in ArcGIS Pro to minimize such errors. The resampling technique used was cubic convolution, producing a geometrically less distorted raster than that obtained by the nearest neighbour technique. Also, it uses values based on the smooth curve on the 16 nearest input cells. Since Millan's ice thickness model had the coarsest resolution of 49.99m, all other datasets were adjusted to match it. The rationale behind taking Millan et al. (2022) ice thickness data as one of the five datasets was to introduce at least one ice thickness model that is recent and is not a part of Farinotti's Ensemble

(Farinotti et al. 2019). This was done to check Millan's ice thickness performance (whether it produces any anomalies) in estimating overdeepenings in comparison with other ice thickness models across regions.

It is acknowledged as the data becomes coarser it might lead to change in size of the estimated overdeepenings. At the same time, it ensures that the overdeepenings thus generated might have less error in data as the predicted depth values become smoother with the use of cubic convolution while down sampling. Overdeepenings with at least 4 pixels were taken into consideration to ensure that area and volume generated by these would be high enough for inter-model and inter-regional comparison.

Dataset	Resolution	
Farinotti's Ensemble Ice Thickness	25m	
GlabTop2 Ice Thickness	25m	
Millan's Ice Thickness	49.99m	
OGGM's Ice Thickness	25m	
HF's Ice Thickness	25m	
ALOS-PALSAR DEM	30 m	

Table 4.3: Datasets and their resolution

4.4. ESTIMATING THE OVERDEEPENINGS

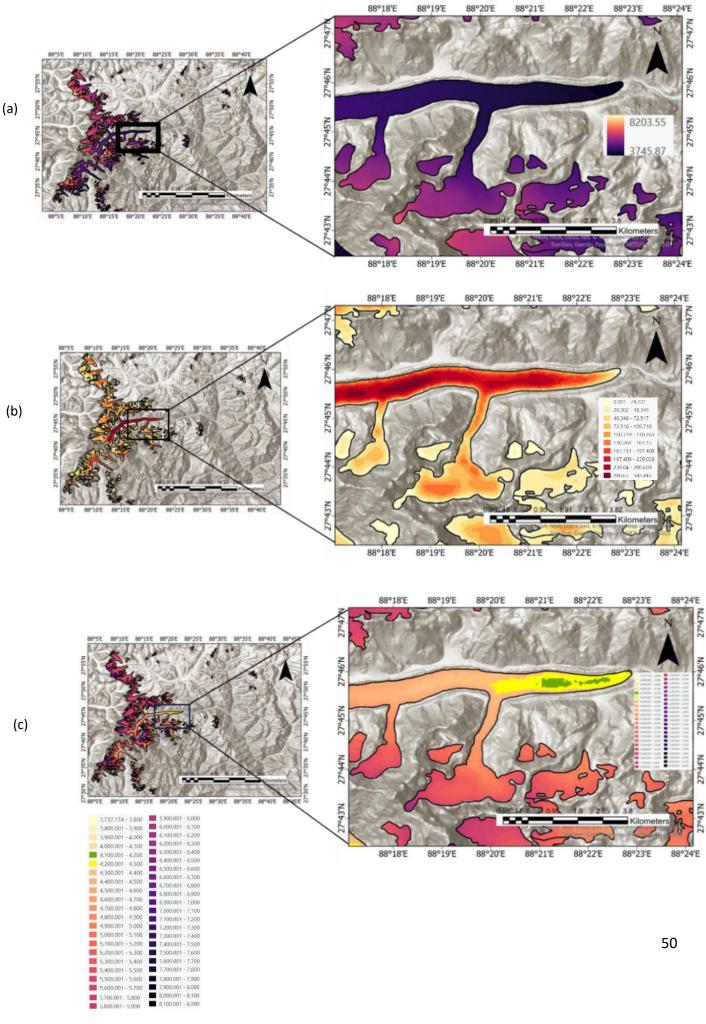
Frey et al. (2010) explained a multi-level strategy to investigate the formation of future glacial lakes in which they identified four levels. This study implements the first three levels of this strategy to determine the overdeepenings using the DEM and the ice thickness. The first two levels of identifying the study area with glaciers and selecting the appropriate datasets for analysis are explained in the section 4.4.1. and the third level of performing the analysis for estimating the overdeepenings is explained in 4.4.2. The aim of this study was to find consistency among the datasets, the fourth level of validating the datasets with ground data was avoided, as measuring the accuracy of the datasets was not in the scope of this study.

4.4.1. Estimating the Bed Topography

Understanding the bed topography beneath the glacier aids in a better understanding of the landscape evolution in the glaciated region (Linsbauer et al. 2012). This serves as the key to predicting overdeepening locations and characteristics. Since the bed topography beneath the glacier is covered by ice, and the DEM gives information about the elevation of the ice thickness and the bed topography, the bed topography was calculated by subtracting the ice thickness from the DEM.

4.4.2. Predicting the Overdeepenings

With a simulated bed topography, it was possible to automatically identify overdeepening locations by using hydrological tools to identify surface (glacier subsurface) depressions (Figure 4.6). These depressions were filled using the Fill Tool in ArcGIS to level the bed topography. The filled bed topography was then subtracted from the original bed topography (with sinks) to yield the overdeepenings.



(b)

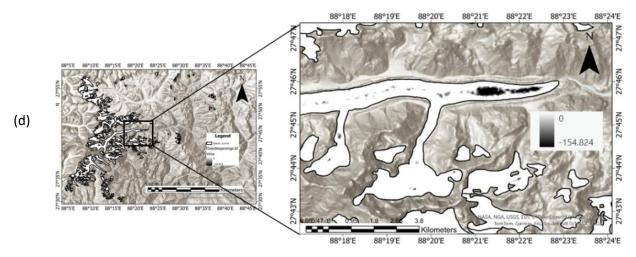


Figure 4.6 : (a) Elevation of the profile (m) (b) Ice Thickness of the region (m) (c) Bed Topography (m) (d) Overdeepenings (m)

4.5. THRESHOLDING

The overdeepenings obtained are large in number, and the smaller overdeepenings (with area less than 0.01 km²) can create noise and give false positives (Viani et al. 2020). Filtering out this noise, and improving the confidence in the locations of the overdeepenings that remain, is therefore important. Here, a thresholding approach was used to remove the smallest of the identified overdeepenings (with area less than 0.01km²).

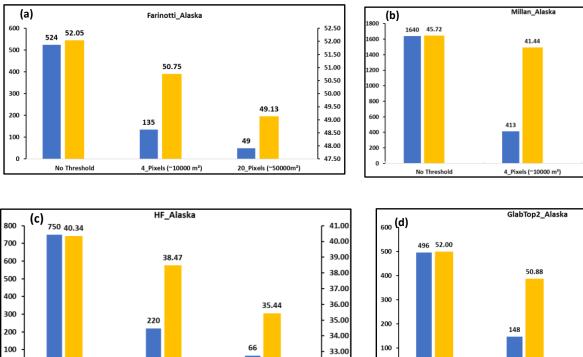
Two thresholds commonly used in previous studies were applied to the overdeepenings, and the resulting changes in their total area and the number of overdeepenings were analysed. Previous studies have typically used thresholds of 0.01 km² (Colonia et al. 2017; Linsbauer et al. 2012; Magnin et al. 2020; Srinivasalu et al. 2024) and 0.05 km² (Kapitsa et al. 2017; Viani et al. 2020). In the current study, thresholds of 4 pixels and 20 pixels were selected, as they closely approximate 0.01 km² and 0.05 km², respectively. The equation below shows the area of one pixel of the overdeepenings.

Area of one pixel = 49.999 $m * 49.999 m \approx 2499.901 m^2$

Since this study has five regions and five ice thickness models, it is important to assess the sensitivity of all different regions and models. Figures 4.7 to 4.11 (below) show the result of applying different thresholds to each of the studied regions. In each case, the y1-axis denotes the number of overdeepenings and the y2-axis denotes the total area of the overdeepenings in square kilometres. It can be seen that although the number of overdeepenings changes markedly once the 4-pixel threshold is applied, the change in the total area of the overdeepenings is relatively less. For 20 pixels the change in the number of overdeepenings reduces to as low as 4 overdeepenings for Farinotti's Ensemble and HF in the European Alps. This suggests that the number of overdeepenings

is highly sensitive to the 20-pixel threshold, leaving limited scope for further in-depth analysis. This can be noted from Appendix I, which shows that changing the threshold from no threshold to 4 pixels yields a maximum percentage decrease in number of overdeepenings of 78.33% (for New Zealand Southern Alps and Farinotti's Ensemble as model), whereas the decrease in area of overdeepenings is 8.94%. When the threshold changes from no threshold to 20 pixels, the percentage decrease of number of overdeepenings is 90%, whereas the percentage decrease in area is 15.08%.

A 0.01 km² threshold has been commonly used in studies of the Andes (Colonia et al. 2017; Drenkhan et al. 2018), Himalayas (Allen et al. 2016; Srinivasalu et al. 2024; Zhang et al. 2019), and European Alps (Magnin et al. 2020; Otto et al. 2022). Sensitivity analysis (Figure 4.7, 4.8, 4.9, 4.10, 4.11) also shows that a 4-pixel threshold works well to retain a substantial number of overdeepenings, ensuring good representation across all regions. Therefore, a 4 pixel-threshold was chosen for further analysis.



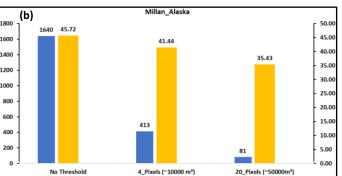
20_Pixels (~50000m²)

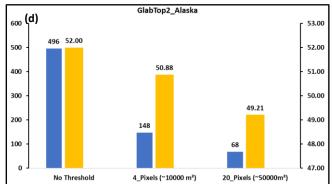


0

No Threshold

4_Pixels (~10000 m²)





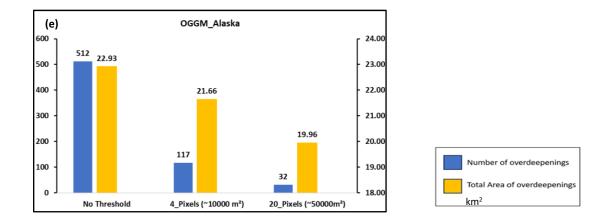


Figure 4.7: Sensitivity analysis for the threshold for selected glaciers in Alaska

32.00

European Alps

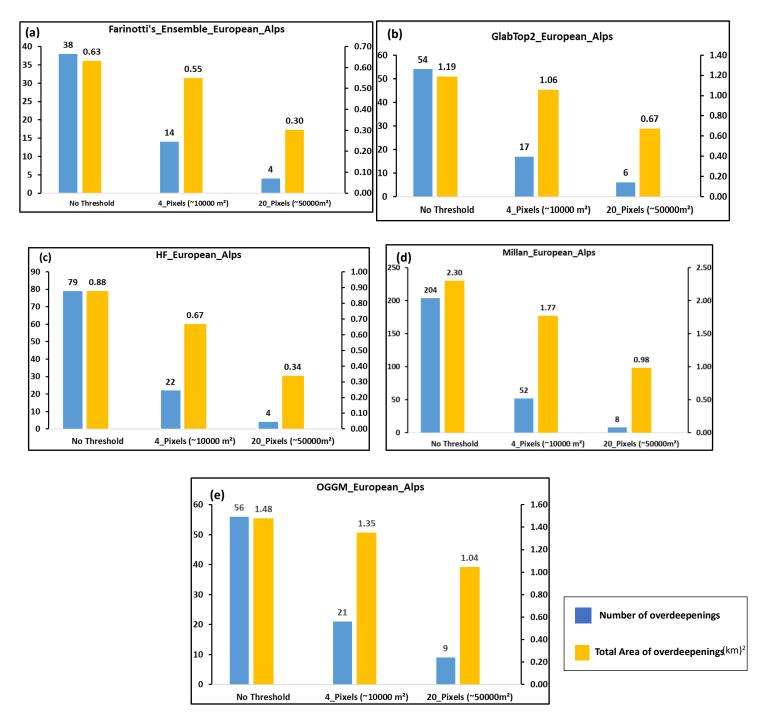
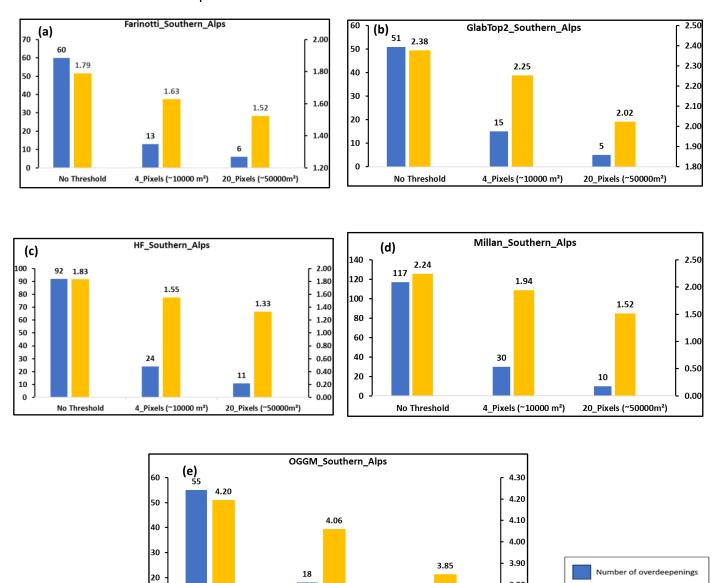


Figure 4.8: Sensitivity analysis for the threshold for selected glaciers in the European Alps



3.80

3.70

3.60

8

20_Pixels (~50000m²)

4_Pixels (~10000 m²)

Figure 4.9: Sensitivity analysis for the threshold for selected glaciers in the New Zealand Southern Alps

New Zealand Southern Alps

10

0

No Threshold

55

Total Area of overdeepenings Km²



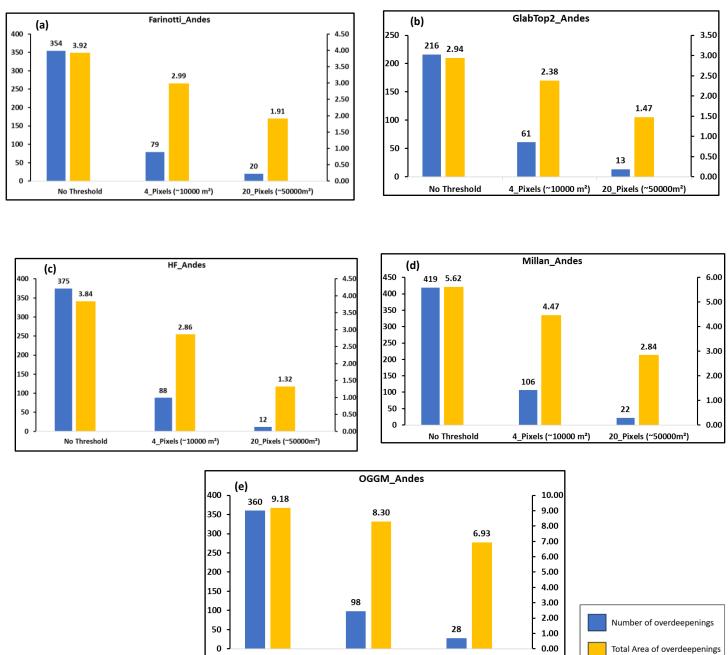


Figure 4.10: Sensitivity analysis for the threshold for selected glaciers in Andes

4_Pixels (~10000 m²)

20_Pixels (~50000m²)

No Threshold

km²

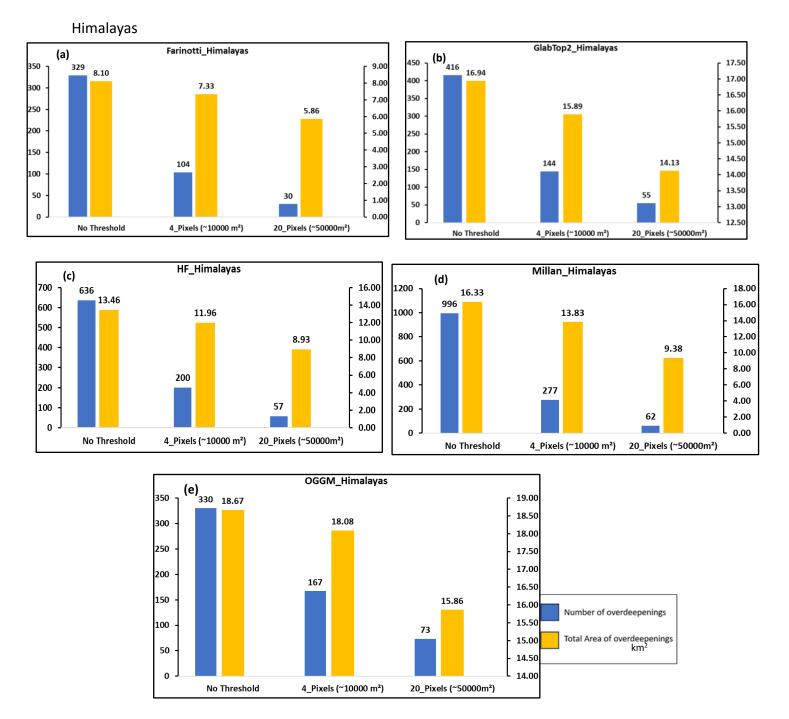


Figure 4.11: Sensitivity analysis for the threshold for selected glaciers in the Himalayas

4.6. ANALYSIS

4.6.1. Zonal Statistics

The overdeepenings were initially in raster format, with depth values assigned to individual pixels. For the analysis of each overdeepening, the pixels were first converted into polygons. Each pixel became an individual polygon. Adjacent polygons (until there was no adjacent polygon) were then merged using dissolve tool in QGIS to generate entire overdeepenings. These overdeepening polygons were used as zones in the Zonal statistics as a Table tool in ArcGIS, allowing for the calculation of various statistics including the mean, median, maximum, count, sum and minimum depth values. These statistics were then used to determine the area and volume of each overdeepening, as well as the total area and volume of all overdeepenings. The following equations state how area and volume of overdeepenings are calculated.

Area of an overdeepening = Count (the total number of pixels in an overdeepening) x Area of one pixel

Volume of an overdeepening = Sum (sum of the depth of the pixels in the overdeepening) x Area of one pixel

This was done for all the models and all the regions to estimate the area and volume of the overdeepenings and their distribution.

4.6.2. Combinations of ice thickness datasets

Since ice thickness models vary in their behaviour, they produce overdeepenings with differing depths, locations, and areas. To evaluate the relationship and performance of individual models with each other, the overdeepening rasters from each model were combined, with pixels assigned values based on their presence across models. This approach allows for further analysis of model comparisons, accounting for differences in the modelling approach and ice thickness estimates across models.

Overdeepening pixels were assigned values using the Con tool in ArcGIS: pixels with depth values less than 0, indicating depressions, were set to 1, while those with depth values of 0 or more were set to 0. This procedure was applied across all ice thickness models.

The overdeepenings from different ice thickness models were then combined using the raster calculator. Pixels were then analysed for their presence (and absence) within different model

combinations. Pixels were assigned a ranking number based on how many models they appeared in; for example, for a two-model combination a pixel that fell in both ice thickness models was given a rank of 2, whereas a pixel that occurred in just one of the models was given a rank of 1. If the ice thickness models are combined with each other in three different ways then it will give 25 model. In the first case, two different models are combined (${}^{5}C_{2}=10$), in the second case three different models are combined (${}^{5}C_{4}=5$). Adding all the possible model combinations, there was a total of 25 model combinations assessed for all given regions.

5. RESULTS

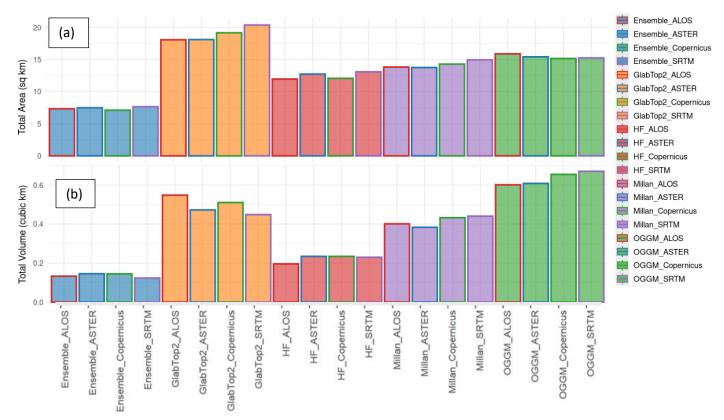


Figure 5.1: Total (a) area and volume (b) for all the model-DEM combinations for Himalayas

The results are structured in two main sections; the first section presents an analysis of the effects of using both different ice thickness models and different DEM combinations for the Central Himalayan region, as a proof-of-concept and to determine the focus of subsequent steps. Based on these findings, the second section presents a regional intercomparison of overdeepening characteristics using different ice thickness models but with a consistent DEM to isolate the ice thickness model effects.

5.1. MODEL AND DEM COMBINATIONS FOR GLACIERS IN THE CENTRAL HIMALAYAS

Four DEMs (ALOS, ASTER, Copernicus and SRTM) and five ice thickness models (Farinotti ensemble, GlabTop2, Huss and Farinotti, Millan and OGGM) were chosen to make an initial assessment of the extent to which the choice of input dataset is a determinant of the resulting overdeepening characteristics. Figure 5.1 shows the total area and volume of overdeepenings for the test region of the Central Himalayas.

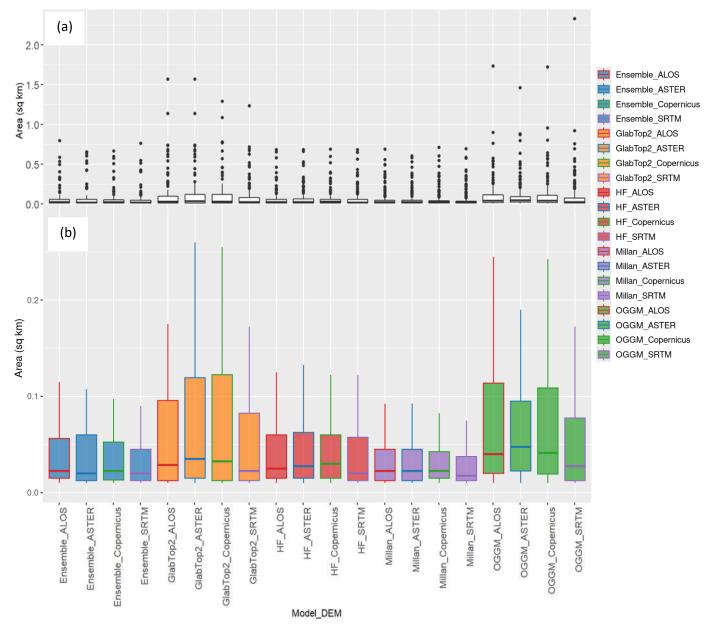


Figure 5.2: (a) Overdeepening areas for all the Model-DEM combinations (b) As a) but with outliers removed to improve visualisation. Note the box midpoint bars represent the median value and the whiskers represent the range.

5.1.1. Overdeepening Area

The total overdeepening area ranges from 7.12 km² (Farinotti's Ensemble-Copernicus model-DEM combination) to 20.37 km² (GlabTop2- SRTM; Figure 5.1). DEM choice seems to make more of a difference for some models than it does for others. For example, Farinotti's Ensemble shows a range of 0.53 km² depending on DEM choice, whereas for OGGM the range is much larger at 2.29 km². Among all the DEMs, SRTM gives the highest total area (highest total area as 20.37 km² for GlabTop2-

SRTM for all the models (see Appendix G for reference) except for OGGM, where ALOS gives the highest total area (15.88 km²).

Given this range, the GlabTop2 model paired with any DEM consistently results in the highest total area of overdeepenings, while Farinotti's Ensemble paired with any DEM produces the lowest total area.

Similarly, the range, distribution and median of the area of individual overdeepenings remain consistent irrespective of DEM choice, although there is again considerable variation between ice thickness models (Figure 5.1 and Figure 5.2). Here, GlabTop2 (gives an overdeepening with the highest area of 1.56 km² when combined with ALOS-PALSAR) and OGGM (with SRTM gives 2.32 km² as the highest area for an overdeepening) consistently produce the largest overdeepenings, Figure 5.1(a) and Figure 5.2 (b).

5.1.2. Overdeepening Volume

The total volume of the overdeepenings ranges from 0.123 km³ (Farinotti's Ensemble- SRTM) to 0.670 km³ (OGGM-SRTM; Figure 5.1 (b), Appendix G). OGGM follows the trend of predicting the highest volume across (0.670 km³) all the DEM combinations whereas Farinotti's Ensemble predicts the lowest volume. Again, DEM choice appears to affect some models more than others, for example, Farinotti's ensemble has a range of 0.022 km³ depending on the DEM, whereas GlabTop2 ranges by 0.099 km³.

Combined, Figures 5.1 and 5.2 show very clearly that the choice of DEM has little influence over the area and volume estimated by each combination, but there is large variability in the results depending on the ice thickness model that is chosen. Assessing the impact of ice thickness model choice across other glacierized regions therefore became the primary focus for the study.

5.2. INTER-REGIONAL COMPARISON USING FIVE ICE THICKNESS MODELS

The five regions considered for subsequent analysis were Alaska, the Peruvian Andes, the New Zealand Southern Alps, the Central Himalaya and the European Alps. ALOS-PALSAR DEM has been chosen for the analysis here.

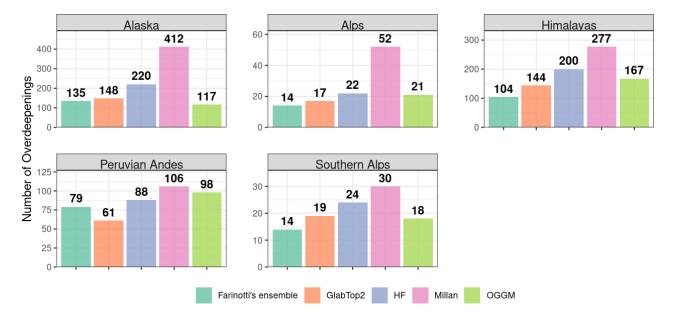


Figure 5.3: Number of overdeepenings across all regions

5.2.1. Number of Overdeepenings

Among all the regions Alaska has the highest number of overdeepenings and also has the widest range, with Millan predicting 412 overdeepenings versus OGGM predicting 117 (Figure 5.3). The

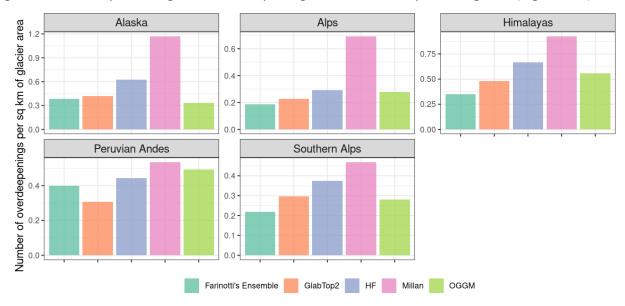


Figure 5.4: Number of overdeepenings per glacier area across all regions and all models. Note the y-axis varies between regions.

New Zealand Southern Alps contain the fewest overdeepenings and also the least variation among the model predictions, taking the difference between the highest and lowest number of overdeepenings as the basis for calculation. Millan consistently has the highest number of overdeepenings for all the regions and Farinotti's Ensemble has the least overdeepenings for the Himalayas, European Alps and New Zealand Southern Alps. For the Peruvian Andes, GlabTop2 predicts the lowest number of overdeepenings.

Alaska and the Himalayas have a higher number of overdeepenings per glacier area compared to the other regions. Among all the models Millan predicts the highest number of overdeepenings per glacier area for all the regions with the highest for Alaska, which is 1.17 overdeepenings per km² of glacier area (Figure 5.4).

5.2.2. Area of overdeepenings

As well as having the highest number of predicted overdeepenings, Alaska also has the highest total area among all the regions (Figure 5.5). In Alaska, GlabTop2 has the highest total predicted area (50.87 km²). In contrast, OGGM has the highest total predicted area for the Central Himalayas (18.08 km²), Peruvian Andes (8.29 km²) and New Zealand Southern Alps (4.06 km²). For the European Alps, Millan predicts the highest total area of 1.77 km².

Except for Alaska, Farinotti's Ensemble and Huss and Farinotti consistently show the lowest total area of overdeepenings

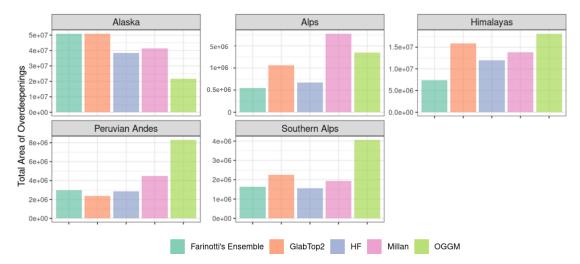


Figure 5.5: Total area of the overdeepenings across all the regions in m², note the values of y-axis differs based on the region taken into account

Alaska

All models predict the majority of overdeepenings to be smaller than 0.1km², while a small number of exceptionally large overdeepenings exceed 10 km² (Figure 5.6). Farinotti's Ensemble (0.0099 km² to 12.11 km²) and GlabTop2 (0.0099 km² to 8.8 km²) show the widest range of overdeepening areas. It is worth noting that several particularly large overdeepenings substantially increase the mean of both of these models.

More generally, the overdeepening area is highly consistent between three of the models, (HF, OGGM and Farinotti's Ensemble) with the median value ranging between 0.0225 km² and 0.0295 km²; GlabTop2 predicts larger overdeepenings compared to other models (median of 0.0412 km²), and Millan predicts smaller overdeepenings (0.017 km²).

European Alps

GlabTop2 and OGGM predict larger overdeepenings and a more varied range compared to the other models (Figure 5.6). As compared to OGGM, GlabTop2 has a higher number of large overdeepenings. Millan has the smallest range, as well as predicting generally smaller overdeepenings (median of 0.015 km²), albeit with several large (> 0.1 km²) overdeepenings.

Himalayas

There is reasonable consistency between the results of the five models, with four of the median values ranging from 0.02 km² to 0.03 km² (Figure 5.6). The exception is OGGM, which estimates larger overdeepenings with a median of 0.040 km². GlabTop2 (0.0099 km² to 1.567 km²) and OGGM (0.0099km² to 1.732 km²) show the most varied range.

Peruvian Andes

There is marked consistency in the area of the overdeepenings predicted by all five models (median range of 0.022 km² to 0.024 km²) (Figure 5.6). The models also show a similar range, except for OGGM which has several very large overdeepenings with sizes greater than 1 km².

New Zealand Southern Alps

Farinotti's Ensemble, Huss and Farinotti and OGGM are consistent in their prediction of the largest overdeepenings (median values of 0.036 km², 0.034 km² and 0.033 km², respectively) (Figure 5.6), whereas GlabTop2 and Millan predict smaller overdeepenings (median values of 0.017 km² to 0.023 km²). OGGM shows the broadest range, as a consequence of having several very large predicted overdeepenings.

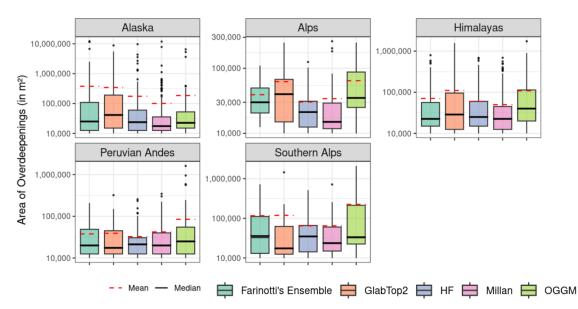


Figure 5.6: Area of individual overdeepenings with the outliers for all five regions. The black solid line represents the median and the red dotted line represents the mean. The black dots represent the outliers. The y-axis are logarithmic to fit the results and inconsistent to fit the variability between the regions and the overdeepenings.

Summary

The range of the predicted overdeepening areas changes from one region to another, with Alaska showing the largest individual overdeepenings as large as 12.11 km². The largest overdeepening for all other models is around 1 km² whereas for the Alps it is just 0.3 km². Across all regions and models, a few exceptionally large overdeepenings heavily influence the mean. The mean size of overdeepenings varies by region, ranging from 0.015 km² (Alps) to 0.37 km² (Alaska), with Alaska also showing the highest mean values across all models, from 0.010 km² to 0.376 km². This suggests that the very few large overdeepenings that exist are also highly variable in size.

While the mean shows a wide range across regions, the median values, ranging between 0.014 km² and 0.041 km², suggest that the majority of overdeepenings across all regions and models are relatively similar in size. OGGM displays a wide range of overdeepening sizes in all regions, except in

Alaska, where it has the narrowest range. In Alaska, GlabTop2 and Farinotti's Ensemble show the widest range of overdeepening sizes. Millan consistently estimates smaller overdeepenings for all of the regions. The Millan data also show a relatively narrow range for overdeepenings across all regions, yet in each region, it also identifies several exceptionally large overdeepenings that fall outside this typical range. Appendix B, C, D, and E represent the distribution of depth of all the pixels of the overdeepenings for comparison. Appendix F has the values of area and volume of the overdeepenings.

5.2.3. Volume of Overdeepenings

	Farinotti's Ensemble	GlabTop2	HF	Millan	OGGM
Alaska	3.700	6.86	1.957	2.740	0.585
Himalayas	1.427	0.553	2.020	0.413	0.589
Peruvian Andes	0.037	0.439	0.294	0.096	0.330
European Alps	0.0347	0.079	0.256	0.474	0.246
New Zealand Southern Alps	0.00319	0.017	0.0058	0.043	0.025

Table 5.1: Total volume of overdeepenings for all the ice thickness models across all the regions (in km³)

Among all the regions, in Alaska, the total volume of overdeepenings estimated by all the models is notably high (Table 5.1). Among all the models, GlabTop2 estimates a volume of 6.86 km³ for Alaska, accounting for 10.9% of the total glacier volume estimated by the model - the highest total volume across all regions and models. In contrast, the New Zealand Southern Alps have a lower total volume of overdeepenings across all models compared to other regions. The Farinotti Ensemble produces the smallest estimate, with a volume of 0.00319 km³, representing only 0.1% of the total glacier volume estimated by this model. This is represented pictorally in Appendix A.

The total estimated volume is not only inconsistent between regions but also between the models of a given region.

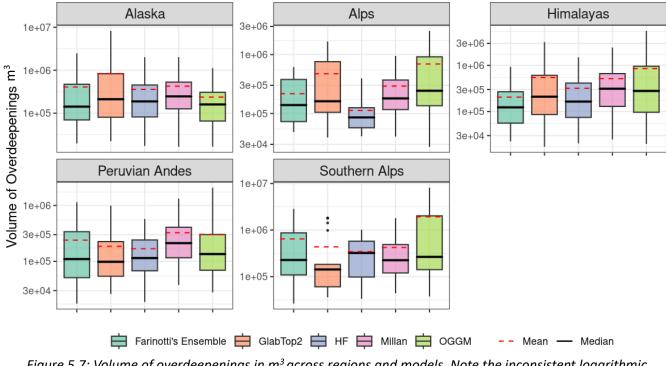
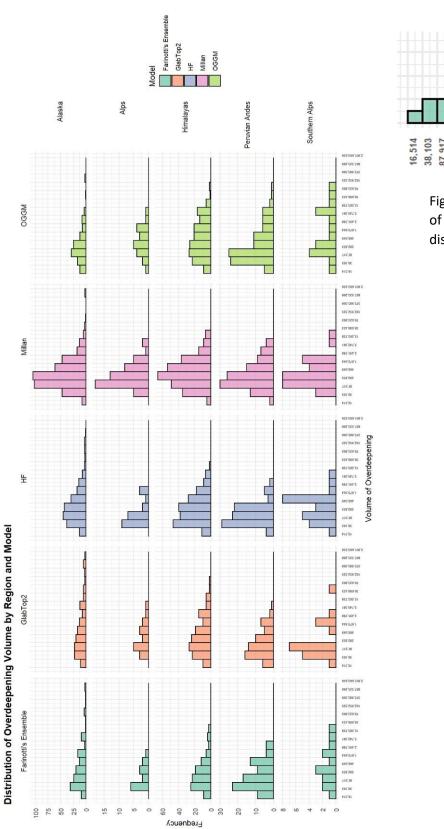


Figure 5.7: Volume of overdeepenings in m³ across regions and models. Note the inconsistent logarithmic scales across regions.



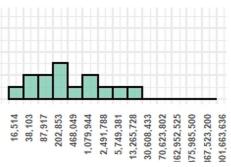


Figure 5.8 (a) : Zoomed version of Farinotti's ensemble volume distribution for Southern Alps

Range of volume in m ³
16,514
38,103
87,917
202,853
468,049
1,079,944
2,491,788
5,749,381
13,265,728
30,608,433
30,608,433
70,623,802
162,952,525
375,985,500
867,523,200
2,001.663.636

Figure 5.8:(b) Distribution of volume of the overdeepenings. The y-axis denotes the frequency, and the x-axis denotes the volume of the overdeepenings in m^3 . Each column represents a model, and each row represents a region.

Alaska

Huss and Farinotti and GlabTop2 both predict most of their overdeepenings to range from around $38 \times 10^{-6} \text{ km}^3$ to $470 \times 10^{-6} \text{ km}^3$ (Figure 5.7) but they also have a long tail (reaching around 0.163 km³) of overdeepenings that have higher volumes (Figure 5.8). Farinotti's Ensemble, Millan and OGGM also suggest most of the overdeepenings will be less than $5 \times 10^{-4} \text{ km}^3$ in volume (Figure 5.7) but have very few overdeepenings with volumes greater than 10^{-3} km^3 (Figure 5.8).

Alps

All the models except for OGGM, show a majority of the overdeepenings to be within 5 x 10⁻⁵ km³ (Figure 5.7). For OGGM the volume of estimated overdeepenings is more distributed (Figure 5.8), with fewer small features, and a greater number of larger features, compared to the other model results.

Himalayas

All the models except OGGM are left skewed suggesting that most of the overdeepenings have a low volume (Figure 5.8). OGGM again has a uniform spread throughout suggesting that there are overdeepenings with both high and low volume. The Himalayas show a slightly more balanced distribution of overdeepening volumes compared to other regions, indicating a relatively higher prevalence of larger overdeepenings.

Peruvian Andes

All the models except OGGM estimate that most of the overdeepenings will be small in volume (Figure 5.7). For OGGM, although the majority of overdeepenings are also predicted to be small, there is a high number of larger overdeepenings, reaching as large as around 0.163 km³.

New Zealand Southern Alps

Farinotti's Ensemble and OGGM show a normal distribution suggesting that both of these models predict overdeepenings that incorporate that there are more overdeepenings in the middle and fewer overdeepenings at the extreme ends (Figure 5.8).

GlabTop2 estimates that most of the overdeepenings will be small in volume (up to $2.02 \times 10^{-4} \text{ km}^3$) but three overdeepenings have a very high volume that increases the mean of the distribution. For Huss and Farinotti most of the overdeepenings are within $5 \times 10^{-4} \text{ km}^3$ and very few overdeepenings

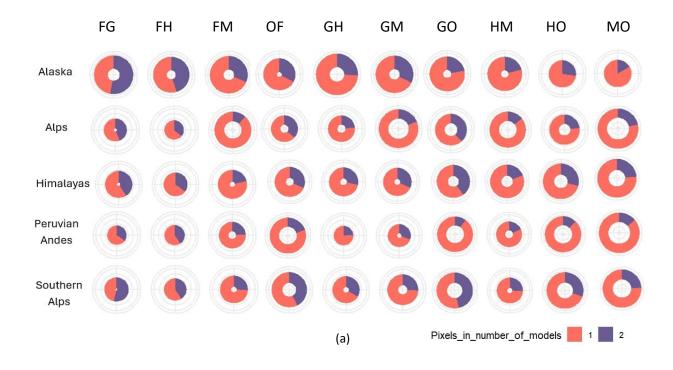
exceed 10⁻³ km³. As with other regions, OGGM shows considerable variation in predicted overdeepening volumes.

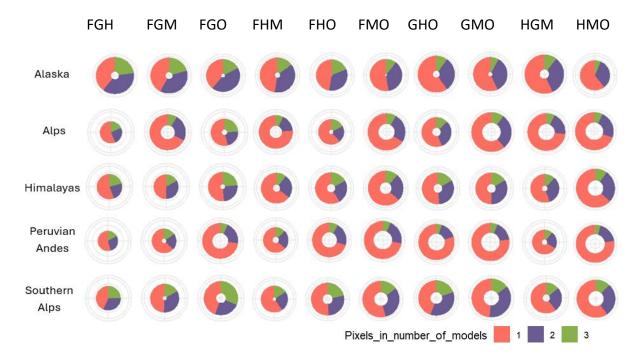
Summary

Figure 5.7 and 5.8 show that most of the overdeepenings across all models and regions are predicted to have a volume of less than 5 x 10^{-4} km³. As compared to other regions, all the models predict the volume of overdeepenings to be widely varying for Alaska. In contrast, for all other regions, the overdeepenings are consistently within a range of 1.6×10^{-5} km³ to 3.1×10^{-2} km³; the results from Huss and Farinotti and Millan are particularly consistent in this regard. Millan predicts mid-volumed overdeepenings with most of them within the range of 1.5×10^{-5} km³ to 50×10^{-5} km³. Among all the regions except for Alaska (where GlabTop2 produces the widest range), OGGM consistently has the widest range for all the regions where it predicts overdeepenings with both small and large volumes. For all the models the vast majority of the overdeepenings are relatively small, with only a few that are exceptionally large. In Alaska, the Himalayas and the New Zealand Southern Alps, there are overdeepenings with predicted higher volumes, irrespective of the model choice, as compared to the other regions.

5.3. PERCENTAGE OF COMMON PIXELS

Each model was compared with another and the percentage of common pixels was plotted using pie charts (Figure 5.9). The size of each pie reflects the total number of pixels for each combination of models, applied across all regions, to visualise the level of agreement when combining 2 models (Figure 5.9a), 3 models (Figure 5.9b), and 4 models (Figure 5.9c). In a combination of two models : the pixel score of 1 means there are that pixel falls in only one model and the a score of 2 means it falls in two models. The pixel score of 3 means that the pixel falls in 3 models and the pixel score of 4 means it is present in 4 models. The size of the pie chart represents the total number of pixels for the model combination; for instance, in Figure 5.9 (a) for Alaska the model combination of GlabTop2 and Huss and Farinotti has the highest number of pixels compared to other two-model combinations, hence the pie chart for this model has the largest diameter with a total of 122381 pixels while Millan and OGGM having a total of 21872 pixels has the smallest diameter (see Appendix H for reference and Figure 5.9(a)).





(b)

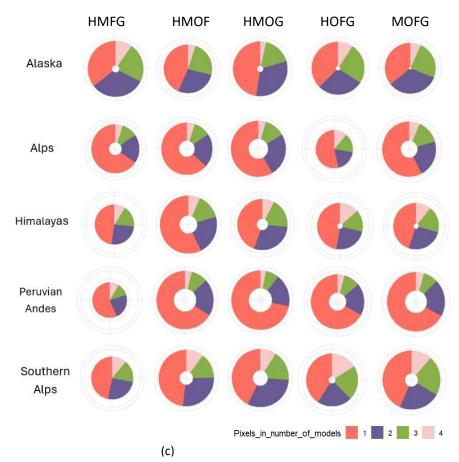


Figure 5.9: (a) Model combinations with 2 ice thickness models (b) Model combinations with 3 ice thickness models (c) Model combinations with 4 ice thickness models. The initial letters of the models have been used in the combination in each case. (H: Huss and Farinotti, F: Farinotti's Ensemble, M: Millan, O: OGGM, G: GlabTop2)

Table 5.2: Model combinations with the highest total pixels and the highest percentage of common
pixels (H: HF, F: Farinotti's Ensemble, M: Millan, O: OGGM, G: GlabTop2)

Region	2 combinations		3 combinations		4 combinations	
	Highest	Percentage	Highest	Percentage	Highest	Percentage
	Pixel	of common	Pixel	of common	Pixel	of common
	Total	pixels	Total	pixels	Total	pixels
Alaska	HG	FG (52.9%)	HGM	FGH	HFMG	HFMG
				(20.6%)		(9.6%)
Alps	MO	FG (43.8%)	GMO	FGO	HMOG	HOFG
				(24.6%)		(11.26%)
Himalayas	MO	FG (40.8%)	НМО	FGO	HMOF	HOFG
				(23.5%)		(13.5%)
Peruvian	MO	FH (40.3%)	НМО	FGH	HMOG	HMOF
Andes				(16.7%)		(8.6%)
New Zealand	MO	FG (53.6%)	НМО	FGO	HMOG	HOFG
Southern				(31.4%)		(16.4%)
Alps						

Table 5.3: Model combinations with the lowest total pixels and lowest percentage of common pixels (*H: HF, F: Farinotti's Ensemble, M: Millan, O: OGGM, G: GlabTop2*)

Region	egion 2 combinations		3 combinations		4 combinations	
	Lowest Pixel	Percentage	Lowest Pixel	Percentage	Lowest Pixel	Percentage
	Total	of common	Total	of common	Total	of common
		pixels		pixels		pixels
Alaska	ОН	MO (16.6%)	НМО	HMO (6.2%)	HMOF	HMOG
						(3.4%)
Alps	FH	FM	FGH	НМО	HOFG	HMOG
		(11.75%)		(6.91%)		(4.3%)
Himalayas	FH	FM (21.2%)	FGM	HMO (9.6%)	HMGF	HMGF
						(9.1%)
Peruvian	FG	GO (9.9%)	FGH	HMO (4.7%)	HMGF	HMOG
Andes						(3.07%)
New	FH	FM (25.7%)	FGH	НМО	HMGF	HMOG
Zealand				(11.96%)		(9.05%)
Southern						
Alps						

Millan and OGGM show the greatest consistency in their identification of overdeepenings (total pixels) for all regions except Alaska (Table 5.2). In contrast, Farinotti's Ensemble and GlabTop2 show the highest percentage of common pixels. The three model combinations for the highest and lowest percentage of common pixels are fairly consistent throughout. As seen from Table 5.3, Farinotti's Ensemble and Millan have the lowest common percentages for the Alps, Himalayas and the New Zealand Southern Alps. The agreement between Farinotti's Ensemble and GlabTop2 is the highest in the New Zealand Southern Alps (53.6%). The least overlap is between GlabTop2 and OGGM (9.9%) in the Peruvian Andes.

As it can be observed through the Figure 5.9 (a,b,c), it is can be observed that there is a disagreement between the models irrespective of the region.

When combined with the rest of the results in this chapter, it becomes clear that overdeepening locations and characteristics vary widely between regions, which may be expected, and between ice thickness models, with only around 50% agreement between any two models in any of the study

regions, and inconsistent patterns both within and between each region depending on the model choice.

6. DISCUSSION

This study has shown that the identification of subglacial overdeepenings using a GIS-based approach is largely insensitive to the choice of DEM, but is highly dependent on the ice thickness model that is used in the calculation. This raises questions about how robust previous studies that have presented future glacial lake locations using a single ice thickness model may be (e.g. Drenkhan et al. 2018; Viani et al. 2020; Magnin et al. 2020). This discussion will consider the influence of DEM and ice thickness model choice on the overdeepening characteristics, assess the uncertainties in doing so, and subsequently present a possible way forward for future studies pursuing a similar theme.

6.1. DEM AND ICE THICKNESS CHOICE

6.1.1. DEMs

There is an increasing amount of digital elevation data available for monitoring Earth's surfaces, along with those that are glacierised (Guth et al. 2021). Previous work has shown that uncertainties in DEMs are largely underestimated in the literature (Hugonnet et al. 2021), which makes the first key finding of this study, that DEM choice has little impact on overdeepening size and location, a little surprising.

There are relatively few studies that specifically address the influence of DEM choice on predicted overdeepening characteristics. Those that exist are locally focused; for example, DEM choice was found to affect the morphometry, location, accuracy and number of overdeepenings when applied with a single ice thickness model in the European Alps (Gharehchahi et al., 2020) and the Peruvian Andes (Drenkhan et al., 2018). For instance, for Gharehchahi et al. (2020), 125 overdeepenings are predicted by the swissALTI^{3D} and 113 by DHM25 L2. Findings like this are not necessarily widely applicable, since certain local DEMs like swissALTI^{3D}, DHM25L2 and Regione Autonoma Valle d'Aosta (RAVA) are only available for a given area, performing optimally within specific terrains. However, when globally available DEMs were analysed across multiple ice thickness models in the current study, the influence of the models became more pronounced in estimating overdeepenings, as discussed in Section 5.1.

Although the consistency of global DEMs is known to be variable (Li et al., 2018), the vertical differences (which are of the order of 1-10 m) are relatively small in comparison to the inconsistencies between ice thickness estimates. DEMs also tend to perform best over shallow-sloped terrain, making them more reliable over glacier tongues, as suggested by intercomparison studies of DEMs (Magnin et al. 2020; Gharechahi et al. 2020). These two factors together likely explain why there was little impact on overdeepening characteristics using different DEM choices in the current study.

Given the scale of analysis in this study, spanning five regions with over 150 glaciers each, the ALOS-PALSAR DEM was selected as an appropriate data source given its wide swath and a spatial resolution of 30 m (as mentioned in section 4.3.4) ALOS-PALSAR is an open-source DEM and it has been used in recent studies to estimate the overdeepenings (Furian et al. 2021, Furian et al. 2022). Highresolution DEMs have been shown elsewhere to provide more accurate estimates of both ice thickness (Ramsankaran et al. 2018) and overdeepenings location and size (Furian et al. 2021). This choice balances the need for robust data performance with the practical requirement of optimizing processing time. It seems likely that future studies could also adopt an approach of selecting the most appropriate choice of DEM for the scale of analysis, and anticipate that it plays only a minor role in the results that are derived.

6.1.2. Ice Thickness Models

The choice of ice thickness model appears to be a critical step when estimating overdeepening characteristics. When a single DEM is applied across all regions with various ice thickness models, OGGM consistently exhibits a wide range of overdeepening volumes and areas across all regions. Among the models, OGGM, GlabTop2, and Millan predict higher overdeepening volumes for all regions. GlabTop2, OGGM, and Farinotti's Ensemble produce higher mean values for the area. Although the range of median values of area across all regions and models is relatively similar, many large overdeepenings are found in Alaska, with the largest having an area of around 12 km² (Appendix G), predominantly predicted by Farinotti's Ensemble.

Previous studies have noted that HF tends to overestimate the number of overdeepenings compared to GlabTop2 in the Austrian Alps (Otto et al. 2022). A similar trend is observed in our study: HF predicts a greater number of overdeepenings and higher volumes of individual overdeepenings than GlabTop2 across all regions. However, the total area of overdeepenings predicted by GlabTop2 remains higher than that of HF across most regions.

OGGM consistently predicts higher ice thickness volumes and the greatest pixel depths across all regions except for Alaska where GlabTop2 and Farinotti's Ensemble produces the higher volume (see Appendix F for reference). This overestimation (with respect to (wrt) other ice thickness models) may stem from OGGM's methodology, which uses temperature and precipitation as the main driver of the model. This is appropriate for studying long-term glacier fluctuations across broad scales of time and space but has been shown to be somewhat deficient on the scale of individual glaciers and using only limited climate records (Parkes and Goosse, 2020). Furthermore, the model employs inverse distance interpolation to estimate missing values especially of climate data (i.e. away from the centreline), potentially introducing errors that might contribute to overestimation (wrt other models).

Farinotti's Ensemble, on the other hand, tends to underestimate (wrt other models) both the number and volume of overdeepenings across all regions compared to other models, except for the volume in Alaska. This underestimation (wrt other models) could be attributed to its weighted average approach, which has a tendency to smooth out extreme values (Farinotti et al., 2017). Additionally, the models with high weights in Farinotti's Ensemble have primarily been tested and validated on glaciers in regions such as the Himalayas and the Alps, where their performance is well-established. This could explain why the ensemble tends to overestimate (wrt other models) overdeepening areas, and more generally to behave differently when applied to Alaska.

There is no consistent trend of models across all regions in terms of producing overdeepenings with the highest volumes or areas. Each model behaves differently depending on the region. OGGM estimates the highest total area across Himalayas, Peruvian Andes and New Zealand Southern Alps, suggesting that it is sensitive (in predicting the total area of overdeepenings) to regions which are dominated by valley-type glaciers, whereas Farinotti's ensemble predicts the highest total area for glaciers in Alaska, suggesting it is sensitive to larger glaciers, and Millan is sensitive to the region which is dominated by cirque glaciers. While models like GlabTop2 and OGGM tend to overestimate (wrt other models) values across most regions, their outputs do not follow a uniform pattern. Particularly for models like OGGM, which rely on numerous inputs, even small changes in input parameters can impact the estimated glacier thickness. Previous studies have shown that minor shifts in initial parameter values can lead to large variations in final ice thickness outputs (Ramsankaran et al. 2018; Drenkhan et al. 2019; Pandit and Ramsankaran, 2020). For instance, Pandit and Ramsankaran, (2020), changed the shape factor from 0.6 to 0.9 for 31 times and averaged the ice thickness thus generated as the final output. This shows that even a small variability in shape factor can change the output.

The variability in glacier morphology across regions further complicates model predictions, and this may also account for the lack of any consistent trend in the model performance. For instance, the Alps and the Himalayas are characterized by predominantly flat-tongued valley glaciers, while glaciers in the tropical Andes are generally steeper (Drenkhan et al. 2018). Even within the European Alps, there are distinctions: the Swiss Alps host more valley glaciers, whereas cirque glaciers dominate the Austrian Alps (Cogley et al. 2010; Fischer et al. 2014; Otto et al. 2019; Otto et al. 2022). Similarly, glacier geography and type vary across the Himalayas, with the Eastern and Central Himalayas featuring more debris-covered glaciers compared to the Western Himalayas and the Karakoram. The general trend of glacier retreat in the Western, Central, and Eastern Himalayas contrasts with the Karakoram, which includes some advancing glaciers. The Alaskan glaciers have been reported to have long tunnel of troughs and is home to submarine landslides beneath the glaciers (Avidievitch and Coe, 2022). This might have attributed to the large and high voluminous overdeepenings in the Alaska as compared to other regions .

These regional differences highlight the importance of considering local glacier characteristics when interpreting model outputs. Hence if a similar study has to be done on one glacier or very few glaciers, then the local parameters of the glacier(s) like slope, aspect and debris cover can be taken into account.

Therefore, the finding that model outputs can vary both between and within regions leads to the recommendation that future studies should not assume that an ice thickness model performing well for one region can be considered representative elsewhere. Previous studies have compared the outputs of two ice thickness models (Otto et al. 2022). The findings of this study reveal that highly variable ranges of values for the area and depth of overdeepenings can extend both within and beyond the estimates provided by the two models alone. This broader spectrum of results, which incorporates five ice thickness models, adds a much more comprehensive analysis to the literature and provides a major advance in understanding the sensitivity of overdeepening estimates to the input datasets. Millan's dataset, which is not a part of Farinotti's ensemble was introduced into the study and it estimates the highest number of overdeepenings across all regions. The area and

volume of the majority of the overdeepenings estimated by Millan is similar to the area and volume of overdeepenings as predicted by other ice thickness models. This shows that recent models, which are not a part of any consensus may behave differently when compared with other ice thickness models that are a part of a consensus, in regard to some parameter(s).

These findings provide us with a chance to observe the differences in results from five different ice thickness models, they may also have value in a practical sense now that the effect of different models is better understood. In other words, future studies may decide to choose a specific ice thickness model because it fits best for their study purpose. Models that tend to overestimate values for area and volume of overdeepenings (like GlapTop2 or OGGM), may be deemed particularly suitable for predicting worst-case scenarios and serving as an upper limit for hazard planning. On the other hand, models that underestimate overdeepenings may help prioritize glaciers that require immediate attention, guiding more focused mitigation efforts. This assumes that the few(er) pixels defined as being overdeepening locations can be taken with high(er) confidence, but further work is required to establish whether this is actually the case when compared to field validation data. If field validation is not possible it will be best to stick to a model that is a combination of models with different approaches like Farinotti's Ensemble which is most likely to provide a solution with less uncertainty.

6.1.3. Uncertainties

Overdeepenings vary in size and contribute only a small fraction of the total glacier volume. In this study, they are shown to account for 0.1% to 10.9% of the total glacier volume (see Appendix A for reference). Alaska, with its substantial glacier volume ranging from 45 km³ to 62.99 km³, also has the highest total volume of overdeepenings, reaching up to 10.9% of the modelled ice volume. In contrast, the European Alps have a much lower proportion of overdeepenings, contributing only 0.1% to 1.7% of the glacier volume.

Despite their relatively small contribution to glacier volume, the potential downstream impact of these overdeepenings could be high, particularly if they evolve into glacial lakes, which is a likely scenario for those located towards a glacier terminus. One key purpose of identifying overdeepenings is to assess the future location, size and hazard potential posed by specific glaciers. The volume of overdeepenings has been shown to have a greater influence on their hazard potential than their area in previous work (Haeberli et al. 2016).

While overdeepenings detected in many cases have shown reliable performance regionally, significant uncertainties remain regarding their depth, volume, and area on a global scale. In most instances, the location of overdeepenings is accurately identified, but their morphometric characteristics are less well protected. The prediction of deeper and larger overdeepenings tend to be more robust (Cathala et al. 2021), probably because a high number of pixels are well within the modelled overdeepening position and those that are less certain around the periphery have a low impact, relative to the overall area and volume. Based on this rationale, the overdeepenings estimated by GlabTop2 in Alaska and the Himalayas in this study can be considered relatively high-confidence estimates, at least in terms of the location and the ultimate size and volume.

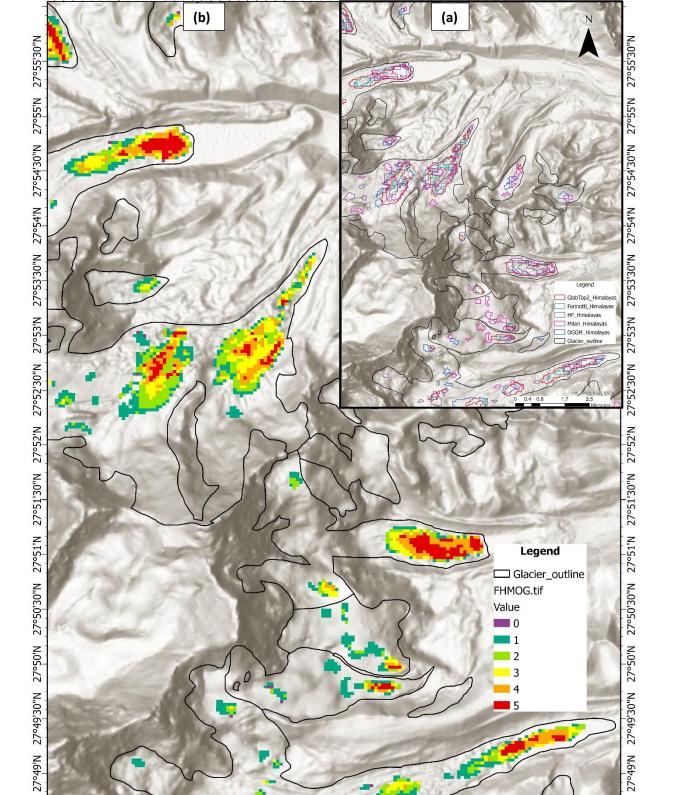
The accuracy of the overdeepening morphometry largely depends on the glacier type, and where along the glacier flowline the overdeepening is located (Sattar et al. 2019). The GlabTop2 model has shown particularly high uncertainties for depicting the morphometry of overdeepenings that are present beneath debris-covered glaciers (Linsbauer et al. 2016), most likely as a consequence of the methodological approach, which uses surface slope as a key input parameter. Debris-covered tongues tend to be very shallow in gradient, and inversions of the basal shear stress equation are increasingly uncertain as the slope approaches zero (Linsbauer et al. 2016). Overdeepenings located towards the termini of valley glaciers in the Austrian Alps were shown to be more accurate than the overdeepenings that were depicted beneath the cirque glaciers (Otto et al. 2022). These findings may be related to the ability of an individual model to correctly estimate ice thicknesses in both shallow and steep gradient terrain (Zorzut et al. 2020).

6.2. A POSSIBLE WAY FORWARD

6.2.1. Combining the models

Since overdeepenings can serve as a foundation for hazard mapping in a region, having confidence in the depiction of their initial morphometry is crucial. If the location and the morphometry of overdeepenings are highly uncertain themselves, these uncertainties will likely propagate during subsequent data processing and analysis. However, these uncertainties can potentially be reduced through various approaches, one of which involves combining multiple ice thickness models to achieve more reliable results. This method has been shown to provide more confident outcomes when combining two ice thickness models (Otto et al. 2022). It is reasonable to expect that combining a larger number of models would further enhance the confidence in overdeepenings estimates, thereby reducing overall uncertainties.

In this approach, pixels can be ranked according to the number of models that depict them as being an overdeepening location. Pixels with a rank of 5 can therefore be considered to be those with the highest confidence (Figure 6.1). For example, in the Himalayas, the number of pixels that are present in all five models accounts for 5.92% of the total pixels in the overdeepenings that are depicted. Focusing attention on these areas may provide an improved judgement about the likely location and morphometry of the overdeepenings. It can also lay the foundation for further analysis of the regions where the pixels have high confidence, as well as provide a basis for communicating the uncertainty in predictions of lake development to communities who may depend on those results for water resource planning and/or hazard mitigation and preparation.



0 0.30.6

88°10'30"E 88°11'E 88°11'30"E 88°12'E 88°12'30"E 88°13'E 88°13'30"E 88°14'E 88°14'30"E 88°15'E 88°15'30"E

Figure 6.1: (a) Outlines of the overdeepenings determined by different five ice thickness models in part of the Central Himalayas (b) Pixels ranked according to the models that they fall in with rank 5 being the pixel that is present in all five models whereas rank 1 being the pixel that is present in only ice thickness model.

1.2

2.5

Kilometers

1.8

88°10'30"E 88°11'E 88°11'30"E 88°12'E 88°12'30"E 88°13'E 88°13'30"E 88°14'E 88°14'30"E 88°15'E 88°15'30"E

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Section 5.4 showed that when two of the five models are used in combination, the number of pixels that are commonly identified as overdeepening locations rises to just over 50%. This low level of agreement highlights that overdeepenings predicted by different ice thickness models can vary significantly across regions. Since HF carries the highest weight (22%) in Farinotti's ensemble, it is perhaps not surprising that the highest overlap occurs between HF and Farinotti's Ensemble in some places. This overlap is most prominent in the Peruvian Andes, where the highest percentage (40.3%) is observed between these two models. Interestingly, Farinotti's Ensemble and GlabTop2 show the highest overlap in pixels across other regions, ranging from 53.6% in the New Zealand Southern Alps to 40.8% in the Himalayas. This difference in overlap percentages is likely due to GlabTop2's broader coverage, which results in more pixels than HF in all regions except the Peruvian Andes. Also, GlabTop2 is given the second-highest weight (19%) in Farinotti's ensemble, after HF, so again there may already be some level of correlation between these two datasets.

Overall, Farinotti's Ensemble, when combined with either HF, GlabTop2, or OGGM, gives strong agreement, most likely as a consequence of these three models being assigned the highest weightings in the ensemble, in that order. Additionally, the agreement between the HF model and GlabTop2 is noteworthy and aligns with findings from previous studies (Otto et al. 2022), further validating the consistency of these models in predicting overdeepenings.

In contrast, some models show minimal overlap, with Farinotti's Ensemble and Millan's model showing the least overlap in the Himalayas (21.2%), New Zealand Southern Alps (25.7%), and European Alps (11.75%). Farinotti's Ensemble consistently predicts smaller total overdeepening areas across all regions, except for in Alaska; Millan's model tends to show higher total area but in locations that are inconsistent with the ensemble outputs. In the European Alps, Millan's model estimates the highest total overdeepening area of all the models. This difference, between the ensemble outputs and those of Millan, are most likely to arise from the fundamental differences in how the ice thickness models are constructed. Specifically, Millan's model employs a velocity-based approach to derive its ice thicknesses. In contrast, the models within the ensemble rely on mass-conservation or shear-stress-based approaches, with the ensemble itself using a weighted method. This indicates that when assessing the potential for agreement or disagreement between two models, greater emphasis should be placed on the physical principles behind their functioning, rather than simply the number of pixels they generate.

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6.2.2. Implications of the findings

Glacier mass loss is occurring now at a rate that has not been experienced since records began, and is accelerating globally due to rising temperatures and changes in precipitation (Hock et al., 2019). One of the clearest impacts of this is the growth of glacial lakes and associated hazard potential (Allen et al., 2016; Shugar et al., 2020). The Himalayas (Rao et al. 2014; Riaz et al. 2014; Westoby et al. 2014; Banerjee et al. 2024) and the Peruvian Andes (Emmer et al. 2014; Huggel et al. 2020; Vilca et al. 2021) have experienced numerous GLOFs in recent decades (Vilímek et al. 2014). The number and size of glacial lakes have been increasing globally, with those in High Mountain Asia expanding at a particularly fast rate. Between 2008 and 2017, the total glacial lake area in this region grew by approximately 6.90%, equating to an increase of around 90.14 km² (Zheng et al. 2021) .

The central and eastern Himalayas are emerging as hotspots for GLOFs (Gardelle et al. 2011; Veh et al. 2020), while glacial lakes in the Cordillera of the Peruvian Andes have grown by 18.3% in area and 9.7% in volume between 1988 and 2016 (Drenkhan et al. 2018). Young proglacial lakes, particularly those located at the glacier terminus, tend to be more unstable than older lakes that have fully developed and are no longer directly connected to glaciers. The dynamics of glaciers are influenced by these lakes, and conversely, the proglacial lake geometry is influenced by the glacier dynamics (Carrivick et al. 2022). The glacial lakes in the early stages of formation, typically found at the glacier tongues where the gradient changes sharply, are particularly prone to causing GLOFs (Emmer et al. 2015b).

One way to prepare for, and ultimately mitigate, the risk of such disasters is to create an inventory of potential future glacial lakes. Mapping overdeepenings can play a crucial role in this process. The first step in overdeepening mapping involves identifying regions with high confidence for the evolution of overdeepenings, which is where combining models can be particularly useful.

Several studies have begun ranking overdeepenings based on their hazard potential in regions like the Himalayas (Pandit and Ramsankaran, 2020) and the Peruvian Andes (Drenkhan et al. 2019; Davies, 2021; Gegg, 2021; Haeberli and Drenkhan, 2022). The current study contributes to this effort by suggesting a ranking of individual pixels; based on their occurrence across different ice thickness models, it can support future hazard assessment efforts. The importance of subglacial overdeepenings extends beyond their potential to develop into glacial lakes. Subglacial overdeepenings are ubiquitous features in glacial systems, yet they remain enigmatic because their origins remain uncertain (Hooke, 1991; Alley et al. 2003; Cook and Swift, 2012). There have been numerous studies that have tried to answer questions regarding the evolution (Patton et al. 2016) sediment transfer (Swift et al. 2018), site and origin of formation of subglacial overdeepenings in a glacial system (Gegg and Preusser 2023; Lloyd et al. 2023). The estimation of area, volume and location of overdeepenings from this study when compared against the local climate, precipitation, bedrock geology (Magrani et al. 2020) and/or subglacial lithostratigraphic data (Preusser et al. 2010) might lead to a better understanding of the processes and factors that govern formation of subglacial overdeepenings (Huuse and Anderson 2000; Liebl et al. 2021; Swift et al. 2021; Evans 2023).

Apart from being important for hazard management and landscape evolution, overdeepenings also are a potential source of hydropower generation. These may act as reservoirs for meltwater and precipitation, accumulating water over time, after deglaciation (Cuellar and McKinney, 2017). A recent study in Mount Massif in the European Alps highlighted the high hydropower potential of future glacial lakes (Magnin et al. 2020). As glacial lakes increase in size and number, they can have both positive and negative impacts on downstream hydropower plants. While they may offer potential energy resources, they also pose risks to downstream areas, potentially leading to destruction and the loss of infrastructure and livelihoods (Rounce et al. 2017; Chettri et al. 2020).

Knowledge of future lake locations can be immensely valuable for mountain communities as well as businesses tasked with providing a sustainable power supply to remote regions (Farinotti et al., 2019a). By estimating the volume and area of overdeepenings, the approach developed in this study can help assess the potential hydropower capacity of these future lakes. Considering a range of values emphasizes the importance of using multiple volume estimates rather than relying on a single fixed value. It can also provide a means for communicating the range of power generation, and the longevity of that resource, to those who may be considering exploiting those locations for commercial benefit. This more nuanced approach – of ranking pixels – provides greater flexibility for decision-makers and stakeholders involved in policymaking for hydropower development. It allows for adjustments and refinements based on more varied projections of future glacial lake dynamics. By incorporating variability into hydropower assessments, this method ensures more adaptive and

resilient planning, helping to mitigate risks while optimising the potential benefits of glacial meltwater resources.

6.3. Limitations

There are a number of limitations within this study, which cannot be addressed within the scope of the work, but may prove fruitful for further investigation:

- Use of five ice thickness models produced estimates of overdeepening characteristics with greater confidence than has been possible in previous approaches. However, none of these results have been validated against existing glacial lake locations, known bed morphometries, or Ground Penetrating Radar (GPR) data, which limits the project to one of intercomparison, rather than being able to confirm their absolute accuracy.
- The ice thickness models used here were released in 2019 and rely on data collected prior to that year. Changes in ice thickness and bedrock topography occurring after 2019 have not been incorporated. Glacier dynamics, such as merging or splitting of glaciers over time can alter glacier outlines, bed topography, and ice thickness. These changes are not addressed in this study.
- Although it was possible to derive detailed insights into the location and volume of potential overdeepenings, it was not possible to evaluate the timing of when these features might develop. Predicting their temporal evolution remains a critical gap, but could be realized by combining these outputs with time-evolving glacier boundaries from a global dataset such as GloGEM (Huss and Hock, 2015). The estimation of overdeepening characteristics currently relies solely on an area threshold, without incorporating factors like slope and any evident surface features (e.g. crevasse fields), which affect and/or indicate the formation and distribution of overdeepenings (Frey et al. 2010). Including additional parameters such as slope could enhance the accuracy and reliability of our predictions.
- As with almost all previous studies, linking overdeepenings to potential future lakes is one thing, but linking those lakes, with reduced uncertainty, to a hazard is far more complex because then knowledge of the lake dam type and other catchment characteristics such as risk of avalanching must be incorporated.
- Combined, these points highlight the need for future studies to integrate updated datasets, validate findings with in-situ observations, and carefully consider the purpose of the work,

as well as the nuances of individual models to better capture the dynamics of glaciers and overdeepening evolution.

7. CONCLUSION

This study has shown for the first time that ice thickness model choice has a greater influence than DEM choice on the modelling of overdeepening location and size using a GIS-based approach. Future research can therefore focus on improving and refining ice thickness models providing a high-resolution, internally self-consistent and accurate glacier bed DEM that is used throughout the rest of the analysis. This work is the first to combine more than two ice thickness models to estimate overdeepenings across five diverse regions, providing a detailed assessment of their volume and area. The estimates reveal stark variations across all the models and regions, with the overdeepening area in Alaska being greatest at 12.11 km² as estimated by Farinotti's Ensemble, and volume ranging from 16.5 x 10^{-6} km³ as estimated by Millan to 2 km³ as estimated by GlabTop2. Despite the high uncertainty, these findings can offer insights into subglacial morphometry, as well as holding relevance for understanding hazard potential, sediment provenance, and glacier dynamics.

The results highlight significant differences in overdeepening characteristics derived from different ice thickness models. For the Himalayas, the overlap can be as high as 40.8% when combining two models but drops to as low as 5.9% with five models. Considering all regions, the agreement between any two models reached a maximum of just 53.6%. While no single model emerges as universally optimal, combining multiple models may provide a more reliable and flexible approach for estimating overdeepenings, balancing overestimation and underestimation.

It is impossible to eliminate uncertainties entirely when modelling overdeepenings, given the inherent complexities of glacier dynamics and ice thickness modelling. Therefore, the most practical approach would be to establish a range of estimates with varying confidence levels. This strategy provides a broader perspective while identifying regions with the highest reliability. Future research can aim to narrow this range and refine parameter estimates to improve accuracy further.

For small-scale studies requiring precision, local ice thickness models can be parameterized for enhanced accuracy. However, this study emphasizes the need to define clear objectives to ensure the correct model is employed, and to consider both overestimation and underestimation when analysing overdeepening locations and characteristics. Finally, this study introduces a novel ranking system for regions based on pixel occurrences across different models, offering a fresh perspective on regional overdeepening assessments. The range of estimates, model overlaps, and the proposed ranking scheme provides a new framework for identifying overdeepenings, paving the way for more robust methodologies in future research. This framework can improve our ability to estimate hazard potential, understand glacial system evolution, and assess the water storage capacity of future glacial lakes for hydropower generation.

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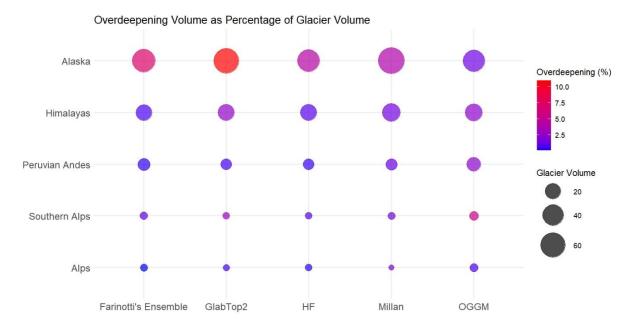
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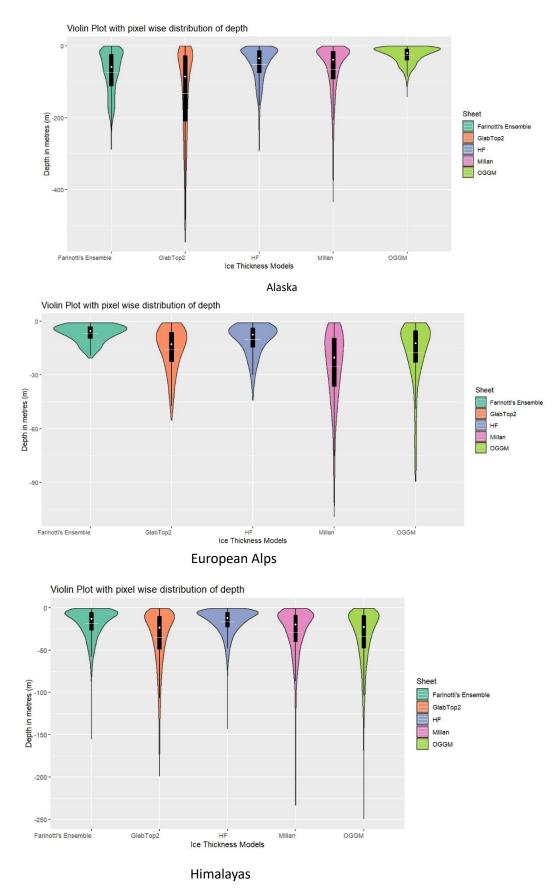
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APPENDIX A



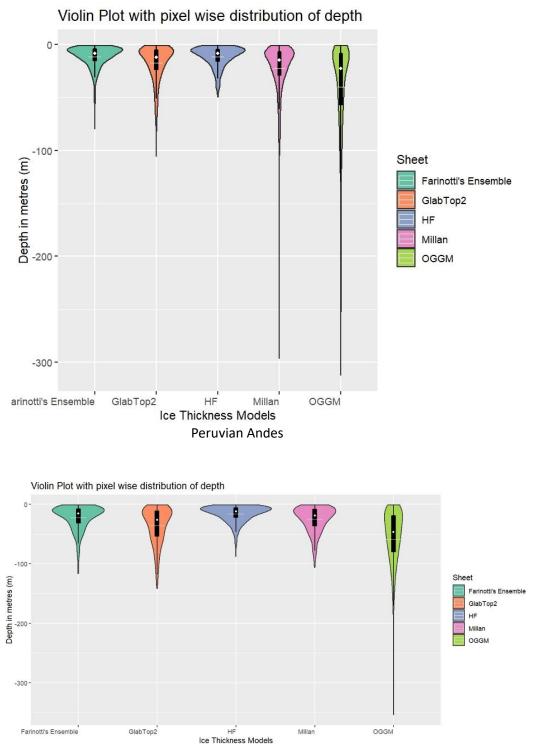
APPENDIX B

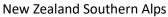


Nb: The violin plots represent the distribution of pixel depth of overdeepenings with the white dot representing the median whereas the line represents the mean of the pixel depths

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APPENDIX C





APPENDIX D

Value for boxplots for area of the overdeepenings (in $\ensuremath{\mathsf{m}}^2\xspace)$

	Region		Model		Median	Mean
	<fct></fct>		<fct></fct>		<dbl></dbl>	<dbl></dbl>
1	Alaska		Farinottis'	Ensemble	24985.	375944.
2	Alaska		GlabTop2		<u>41</u> 226.	343766.
3	Alaska		HF		23736.	174885.
4	Alaska		Millan		17490.	100590.
5	Alaska		OGGM		<u>22</u> 487.	185104.
6	Alps		Farinottis'	Ensemble	<u>29</u> 999.	<u>39</u> 281.
7	Alps		GlabTop2		<u>39</u> 999.	<u>62</u> 201.
8	Alps		HF		<u>21</u> 249.	the second s
9	Alps		Millan		<u>14</u> 999.	
10	Alps		OGGM		<u>34</u> 999.	<u>64</u> 385.
11	Himalayas		Farinottis'	Ensemble	22500.	<u>70</u> 480.
12	Himalayas		GlabTop2		28750.	<u>110</u> 363.
13	Himalayas		HF		25000.	<u>59</u> 824.
14	Himalayas		Millan		22500.	<u>49</u> 936.
	Himalayas		OGGM		40000.	108292.
			Farinottis'	Ensemble		<u>37</u> 804.
17					<u>17</u> 494.	<u>39</u> 045.
	Peruvian				<u>21</u> 243.	<u>32</u> 518.
	Peruvian		Millan		<u>19</u> 994.	<u>42</u> 156.
20	Peruvian				<u>24</u> 992.	<u>84</u> 667.
			Farinottis'	Ensemble	<u>36</u> 143.	<u>116</u> 262.
	Southern		GlabTop2		<u>17</u> 448.	<u>118</u> 595.
	Southern		HF		<u>34</u> 896.	<u>64</u> 703.
	Southern		Millan		<u>23</u> 680.	<u>64</u> 724.
25	Southern	Alps	OGGM		<u>33</u> 650.	<u>225</u> 579.

APPENDIX E

			Jume and are	a of the overbeepening	
	Region	HOUET		vorume_variance	
	<chr></chr>	<chr></chr>		<dbl></dbl>	<dbl></dbl>
	Alaska	Farinotti's	Ensemble	2.18e16	
2	Alaska	GlabTop2		4.05e16	1.02e12
3	Alaska	HF		3.75e15	6.32e11
4	Alaska	Millan		4.82e15	4.99e11
5	Alaska	OGGM		8.33e14	6.93e11
	Alps	Farinotti's	Ensemble	8.82e10	8.10e 8
7	Alps	GlabTop2		2.63e12	4.94e 9
8	Alps	HF		2.16e11	9.15e 8
9	Alps	Millan		3.28e12	2.41e 9
10	Alps	OGGM		4.96e12	3.63e 9
11	Himalayas	Farinotti's	Ensemble	1.20e13	1.70e10
12	Himalayas	GlabTop2		1.52e14	4.59e10
13	Himalayas	HF		5.62e12	1.01e10
14	Himalayas	Millan		1.22e13	7.65e 9
15	Himalayas	OGGM		1.08e14	3.57e10
16	Peruvian Andes	Farinotti's	Ensemble	4.78e11	1.73e 9
17	Peruvian Andes	GlabTop2		2.81e12	2.58e 9
18	Peruvian Andes	HF		6.06e11	1.68e 9
19	Peruvian Andes	Millan		3.31e12	3.30e 9
20	Peruvian Andes	OGGM		2.05e14	4.13e10
21	Southern Alps	Farinotti's	Ensemble	2.48e13	3.67e10
22	Southern Alps	GlabTop2		2.52e14	1.06e11
23	Southern Alps	HF		6.13e12	1.08e10
24	Southern Alps	Millan		2.27e13	1.76e10
25	Southern Alps	OGGM		1.15e15	2.40e11

Variance in different volume and area of the overdeepenings

p-value for ANOVA test of area and volume of overdeepenings across regions

	Region	Volume_p_value	Area_p_value
1	Alaska	9.074045e-04	1.144949e-02
2	Alps	3.298525e-01	5.767190e-02
3	Himalayas	8.485426e-05	1.269341e-05
4	Peruvian Andes	1.851389e-02	4.984693e-03
5	Southern Alps	8.900406e-02	3.069962e-01

APPENDIX F

Region	Model	Volume (in cu m)	Area (in sq m)
Alaska	Farinotti's Ensemble	20702.7558	9994.0762
Alaska	Farinotti's Ensemble	25059.4569	9994.0762
Alaska	Farinotti's Ensemble	27613.0276	9994.0762
Alaska	Farinotti's Ensemble	28046.4255	9994.0762
Alaska	Farinotti's Ensemble	28765.9087	9994.0762
Alaska	Farinotti's Ensemble	33032.4838	9994.0762
Alaska	Farinotti's Ensemble	35981.7854	9994.0762
Alaska	Farinotti's Ensemble	40443.5573	9994.0762
Alaska	Farinotti's Ensemble	41001.0882	9994.0762
Alaska	Farinotti's Ensemble	41594.3035	9994.0762
Alaska	Farinotti's Ensemble	42072.5356	9994.0762
Alaska	Farinotti's Ensemble	42946.6513	9994.0762
Alaska	Farinotti's Ensemble	57033.7605	9994.0762
Alaska	Farinotti's Ensemble	66287.4614	9994.0762
Alaska	Farinotti's Ensemble	66303.4736	9994.0762
Alaska	Farinotti's Ensemble	75889.4665	9994.0762
Alaska	Farinotti's Ensemble	77756.7984	9994.0762
Alaska	Farinotti's Ensemble	83902.6002	9994.0762
Alaska	Farinotti's Ensemble	84078.8873	9994.0762
Alaska	Farinotti's Ensemble	105236.0857	9994.0762
Alaska	Farinotti's Ensemble	112046.6241	9994.0762
Alaska	Farinotti's Ensemble	139397.0509	9994.0762
Alaska	Farinotti's Ensemble	174064.6129	9994.0762
Alaska	Farinotti's Ensemble	287042.0817	9994.0762
Alaska	Farinotti's Ensemble	19688.3424	12492.595
Alaska	Farinotti's Ensemble	35352.5807	12492.595
Alaska	Farinotti's Ensemble	48876.3641	12492.595
Alaska	Farinotti's Ensemble	64929.4711	12492.595
Alaska	Farinotti's Ensemble	68313.0857	12492.595
Alaska	Farinotti's Ensemble	69007.2543	12492.595
Alaska	Farinotti's Ensemble	77199.4201	12492.595
Alaska	Farinotti's Ensemble	80694.0528	12492.595
Alaska	Farinotti's Ensemble	82486.5084	12492.595
Alaska	Farinotti's Ensemble	95377.1222	12492.595
Alaska	Farinotti's Ensemble	119148.4327	12492.595
Alaska	Farinotti's Ensemble	119633.3748	12492.595
Alaska	Farinotti's Ensemble	120150.3413	12492.595
Alaska	Farinotti's Ensemble	144479.1827	12492.595
Alaska	Farinotti's Ensemble	157601.2877	12492.595
Alaska	Farinotti's Ensemble	38914.3124	14991.114
Alaska	Farinotti's Ensemble	40452.0972	14991.114
Alaska	Farinotti's Ensemble	71969.3658	14991.114

Alaska	Farinotti's Ensemble	94437.1276	14991.114
Alaska	Farinotti's Ensemble	99656.5071	14991.114
Alaska	Farinotti's Ensemble	111309.4512	14991.114
Alaska	Farinotti's Ensemble	139447.6801	14991.114
Alaska	Farinotti's Ensemble	230718.3496	14991.114
Alaska	Farinotti's Ensemble	252071.6597	14991.114
Alaska	Farinotti's Ensemble	62722.8323	17489.633
Alaska	Farinotti's Ensemble	68075.7996	17489.633
Alaska	Farinotti's Ensemble	82077.8151	17489.633
Alaska	Farinotti's Ensemble	102945.5732	17489.633
Alaska	Farinotti's Ensemble	115705.0391	17489.633
Alaska	Farinotti's Ensemble	142084.0569	17489.633
Alaska	Farinotti's Ensemble	187380.0847	17489.633
Alaska	Farinotti's Ensemble	278758.1125	17489.633
Alaska	Farinotti's Ensemble	47394.6984	19988.153
Alaska	Farinotti's Ensemble	49509.5338	19988.153
Alaska	Farinotti's Ensemble	148610.34	19988.153
Alaska	Farinotti's Ensemble	254880.6637	19988.153
Alaska	Farinotti's Ensemble	386205.7169	19988.153
Alaska	Farinotti's Ensemble	420859.8591	19988.153
Alaska	Farinotti's Ensemble	59643.2977	22486.672
Alaska	Farinotti's Ensemble	195762.8724	22486.672
Alaska	Farinotti's Ensemble	264498.9861	22486.672
Alaska	Farinotti's Ensemble	69700.2029	24985.191
Alaska	Farinotti's Ensemble	71950.4561	24985.191
Alaska	Farinotti's Ensemble	95762.3309	24985.191
Alaska	Farinotti's Ensemble	207671.7072	24985.191
Alaska	Farinotti's Ensemble	75528.0475	27483.71
Alaska	Farinotti's Ensemble	268155.5712	27483.71
Alaska	Farinotti's Ensemble	88340.8875	29982.229
Alaska	Farinotti's Ensemble	471374.9614	29982.229
Alaska	Farinotti's Ensemble	139546.8035	32480.748
Alaska	Farinotti's Ensemble	400873.5366	34979.267
Alaska	Farinotti's Ensemble	206624.3544	37477.786
Alaska	Farinotti's Ensemble	257203.2007	37477.786
Alaska	Farinotti's Ensemble	359136.8004	37477.786
Alaska	Farinotti's Ensemble	342920.2162	39976.305
Alaska	Farinotti's Ensemble	1077676.052	39976.305
Alaska	Farinotti's Ensemble	255420.1999	42474.824
Alaska	Farinotti's Ensemble	275700.2326	42474.824
Alaska	Farinotti's Ensemble	323477.6997	42474.824
Alaska	Farinotti's Ensemble	479761.409	42474.824
Alaska	Farinotti's Ensemble	279554.7594	44973.343
Alaska	Farinotti's Ensemble	641807.9988	44973.343
Alaska	Farinotti's Ensemble	177279.1077	49970.381
Alaska	Farinotti's Ensemble	270856.6069	49970.381
7 110310		2,0000000	13370.301

Alaska	Farinotti's Ensemble	655889.3131	49970.381
Alaska	Farinotti's Ensemble	1254096.141	49970.381
Alaska	Farinotti's Ensemble	911985.9919	54967.419
Alaska	Farinotti's Ensemble	499972.8178	62462.977
Alaska	Farinotti's Ensemble	1252102.084	67460.015
Alaska	Farinotti's Ensemble	796940.0469	72457.053
Alaska	Farinotti's Ensemble	946310.7394	72457.053
Alaska	Farinotti's Ensemble	1468123.182	72457.053
Alaska	Farinotti's Ensemble	476041.6899	77454.091
Alaska	Farinotti's Ensemble	821706.8611	87448.167
Alaska	Farinotti's Ensemble	593142.9962	94943.724
Alaska	Farinotti's Ensemble	650533.2958	99940.762
Alaska	Farinotti's Ensemble	1382114.744	107436.32
Alaska	Farinotti's Ensemble	1186497.354	109934.84
Alaska	Farinotti's Ensemble	2289063.133	112433.36
Alaska	Farinotti's Ensemble	1283527.854	127424.47
Alaska	Farinotti's Ensemble	2230989.95	129922.99
Alaska	Farinotti's Ensemble	1367834.878	139917.07
Alaska	Farinotti's Ensemble	1629187.84	142415.59
Alaska	Farinotti's Ensemble	4054933.264	142415.59
Alaska	Farinotti's Ensemble	4112528.978	147412.62
Alaska	Farinotti's Ensemble	1517417.847	149911.14
Alaska	Farinotti's Ensemble	1199530.4	157406.7
Alaska	Farinotti's Ensemble	2257641.023	157406.7
Alaska	Farinotti's Ensemble	1342746.599	167400.78
Alaska	Farinotti's Ensemble	1714938.863	167400.78
Alaska	Farinotti's Ensemble	2472218.732	242356.35
Alaska	Farinotti's Ensemble	4851749.025	257347.46
Alaska	Farinotti's Ensemble	6322811.291	289828.21
Alaska	Farinotti's Ensemble	6251043.07	294825.25
Alaska	Farinotti's Ensemble	9267353.977	312314.88
Alaska	Farinotti's Ensemble	5892090.181	332303.04
Alaska	Farinotti's Ensemble	9230973.563	379774.9
Alaska	Farinotti's Ensemble	7775309.274	407258.61
Alaska	Farinotti's Ensemble	10219098.2	447234.91
Alaska	Farinotti's Ensemble	8001102.565	454730.47
Alaska	Farinotti's Ensemble	35064320.05	592149.02
Alaska	Farinotti's Ensemble	11731343.03	707080.89
Alaska	Farinotti's Ensemble	20471921.53	861989.08
Alaska	Farinotti's Ensemble	110682238.7	1791438.2
Alaska	Farinotti's Ensemble	133380255.9	1796435.2
Alaska	Farinotti's Ensemble	136013258.6	1973830.1
Alaska	Farinotti's Ensemble	98521903.12	2258661.2
Alaska	Farinotti's Ensemble	255383278.7	2498519.1
Alaska	Farinotti's Ensemble	418408277.2	6718517.8
Alaska	Farinotti's Ensemble	1201792412	11475698
Ајаѕка	Farmotti s'Ensemble	1201/92412	114/5098

Alaska	Farinotti's Ensemble	1131865190	12115319
Alaska	GlabTop2	22388.7682	9994.0762
Alaska	GlabTop2	23463.5705	9994.0762
Alaska	GlabTop2	43451.113	9994.0762
Alaska	GlabTop2	47936.9795	9994.0762
Alaska	GlabTop2	51862.5702	9994.0762
Alaska	GlabTop2	53734.6295	9994.0762
Alaska	GlabTop2	55523.1202	9994.0762
Alaska	GlabTop2	61815.3196	9994.0762
Alaska	GlabTop2	64108.8821	9994.0762
Alaska	GlabTop2	75026.0257	9994.0762
Alaska	GlabTop2	104332.3855	9994.0762
Alaska	GlabTop2	112816.279	9994.0762
Alaska	GlabTop2	114771.1444	9994.0762
Alaska	GlabTop2	123706.888	9994.0762
Alaska	GlabTop2	143957.0312	9994.0762
Alaska	GlabTop2	270456.1485	9994.0762
Alaska	GlabTop2	22417.4377	12492.595
Alaska	GlabTop2	23173.2153	12492.595
Alaska	GlabTop2	28153.4787	12492.595
Alaska	GlabTop2	31422.1101	12492.595
Alaska	GlabTop2	32011.208	12492.595
Alaska	GlabTop2	32795.9601	12492.595
Alaska	GlabTop2	32985.2095	12492.595
Alaska	GlabTop2	34595.5831	12492.595
Alaska	GlabTop2	39294.3362	12492.595
Alaska	GlabTop2	52734.2459	12492.595
Alaska	GlabTop2	59386.7969	12492.595
Alaska	GlabTop2	69573.63	12492.595
Alaska	GlabTop2	72891.0607	12492.595
Alaska	GlabTop2	103161.2047	12492.595
Alaska	GlabTop2	104075.2747	12492.595
Alaska	GlabTop2	126566.8262	12492.595
Alaska	GlabTop2	135350.0723	12492.595
Alaska	GlabTop2	33319.789	14991.114
Alaska	GlabTop2	39981.7949	14991.114
Alaska	GlabTop2	77873.459	14991.114
Alaska	GlabTop2	79537.8168	14991.114
Alaska	GlabTop2	86559.7167	14991.114
Alaska	GlabTop2	87268.2201	14991.114
Alaska	GlabTop2	220454.0478	14991.114
Alaska	GlabTop2	336419.4002	14991.114
Alaska	GlabTop2	74918.3624	17489.633
Alaska	GlabTop2	83305.42	17489.633
Alaska	GlabTop2	86442.5986	17489.633
Alaska	GlabTop2	88685.5319	17489.633

Alasha	Clab Tar 2	00075 0055	17400 (22)
Alaska	GlabTop2	89075.9255	17489.633
Alaska	GlabTop2	99590.3232	17489.633
Alaska	GlabTop2	65806.3317	19988.153
Alaska	GlabTop2	179989.1408	19988.153
Alaska	GlabTop2	389767.906	19988.153
Alaska	GlabTop2	486621.6616	19988.153
Alaska	GlabTop2	73316.2237	22486.672
Alaska	GlabTop2	120319.6135	22486.672
Alaska	GlabTop2	209766.4129	22486.672
Alaska	GlabTop2	257805.8708	22486.672
Alaska	GlabTop2	296826.6264	22486.672
Alaska	GlabTop2	120781.071	24985.191
Alaska	GlabTop2	172701.2852	27483.71
Alaska	GlabTop2	296194.6768	29982.229
Alaska	GlabTop2	642636.3652	29982.229
Alaska	GlabTop2	659483.6792	29982.229
Alaska	GlabTop2	107960.606	32480.748
Alaska	GlabTop2	155319.0101	32480.748
Alaska	GlabTop2	196783.6906	32480.748
Alaska	GlabTop2	317389.6945	32480.748
Alaska	GlabTop2	116761.8468	34979.267
Alaska	GlabTop2	160686.3121	34979.267
Alaska	GlabTop2	204762.2074	34979.267
Alaska	GlabTop2	213007.2898	34979.267
Alaska	GlabTop2	233644.4717	34979.267
Alaska	GlabTop2	177071.8636	37477.786
Alaska	GlabTop2	306362.4477	37477.786
Alaska	GlabTop2	121567.9581	39976.305
Alaska	GlabTop2	479992.9002	39976.305
Alaska	GlabTop2	473367.1887	42474.824
Alaska	GlabTop2	774911.1729	42474.824
Alaska	GlabTop2	471212.0941	44973.343
Alaska	GlabTop2	275371.143	47471.862
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Alaska	GlabTop2	861850.3033	47471.862
Alaska	GlabTop2	388372.0964	49970.381
Alaska	GlabTop2	419545.0256	49970.381
Alaska	GlabTop2	480720.0083	49970.381
Alaska	GlabTop2	1356741.904	52468.9
Alaska	GlabTop2	1442646.797	52468.9
Alaska		263308.2856	54967.419
	GlabTop2		
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Alaska	GlabTop2	745602.9831	57465.938
Alaska	GlabTop2	699646.3361	59964.458
Alaska	GlabTop2	379457.3975	62462.977
Alaska	GlabTop2	511651.6864	64961.496

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Alaska	GlabTop2	989331.1995	64961.496
Alaska	GlabTop2	1076694.426	69958.534
Alaska	GlabTop2	923914.3464	72457.053
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Alaska	GlabTop2	1706627.901	77454.091
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Alaska	GlabTop2	848853.5513	109934.84
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Alaska	GlabTop2	8208755.667	159905.22
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Alaska	GlabTop2	7065474.275	194884.49
Alaska	GlabTop2	3289533.513	214872.64
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Alaska	GlabTop2	46231447.48	397264.53
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Alaska	GlabTop2	15749878.19	494706.77
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Alaska	GlabTop2	46743728.6	667104.59
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Alaska	HF	56076.3812	9994.0762
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Alaska	HF	86376.4147	9994.0762
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Alaska	HF	473170.162	9994.0762
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Alaska	HF	35399.855	12492.595
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Alaska	HF	730799.7456	69958.534
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Alps	GlabTop2	39371.1786	9999.6495
Alps	GlabTop2	67110.6849	9999.6495
Alps	GlabTop2	75592.4259	12481.42
•	GlabTop2	88301.3484	12491.42
Alps	•		
Alps	GlabTop2	124638.5758	14999.474
Alps	GlabTop2	152181.5557	17499.387
Alps	GlabTop2	161429.6107	19970.271
Alps	GlabTop2	129867.8847	34998.773

Alian	ClabTara		20000 500
Alps	GlabTop2	258695.1564	39998.598
Alps	GlabTop2	227300.9931	44998.423
Alps	GlabTop2	650474.6375	47498.335
Alps	GlabTop2	884441.6568	52498.16
Alps	GlabTop2	1000460.54	67497.634
Alps	GlabTop2	1572172.679	77497.284
Alps	GlabTop2	1666307.563	147494.83
Alps	GlabTop2	3111157.434	199992.99
Alps	GlabTop2	6392830.132	247491.33
Alps	HF	41227.1927	9999.6495
Alps	HF	50721.4888	9999.6495
Alps	HF	56929.157	9999.6495
Alps	HF	57306.9514	9999.6495
Alps	HF	124337.683	9999.6495
Alps	HF	65390.328	12481.42
Alps	HF	50819.752	12499.562
Alps	HF	57713.4313	12499.562
Alps	HF	88273.8836	17499.387
Alps	HF	83668.9424	19999.299
Alps	HF	97000.9947	19999.299
Alps	HF	76729.4884	22499.211
Alps	HF	124740.5009	22499.211
Alps	HF	143815.1497	22499.211
Alps	HF	125925.2131	24962.839
Alps	HF	137959.0318	27499.036
Alps	HF	395128.3474	32451.691
Alps	HF	252481.9953	32498.861
Alps	HF	667613.319	44998.423
Alps	HF	1418257.786	64997.722
Alps	HF	1431639.275	102496.41
Alps	HF	1352650.344	124995.62
Alps	Millan	40706.9789	9999.6495
Alps	Millan	57708.233	9999.6495
Alps	Millan	64515.8119	9999.6495
Alps	Millan	80124.1907	9999.6495
Alps	Millan	97027.8276	9999.6495
Alps	Millan	103843.948	9999.6495
Alps	Millan	109427.6514	9999.6495
Alps	Millan	117073.5179	9999.6495
Alps	Millan	122377.1819	9999.6495
Alps	Millan	140063.5867	9999.6495
Alps	Millan	147411.7207	9999.6495
Alps	Millan	160865.8101	9999.6495
Alps	Millan	217180.5578	9999.6495
Alps	Millan	125613.1776	12481.42
		-	-
Alps	Millan	259032.7275	12481.42

Alps	Millan	65867.1999	12499.562
Alps	Millan	129386.0962	12499.562
Alps	Millan	130814.3578	12499.562
Alps	Millan	109618.615	14999.474
	Millan	111542.8939	
Alps			14999.474
Alps	Millan	115273.0908	14999.474
Alps	Millan	134950.2762	14999.474
Alps	Millan	158224.655	14999.474
Alps	Millan	202143.2395	14999.474
Alps	Millan	224794.0886	14999.474
Alps	Millan	299316.8828	14999.474
Alps	Millan	355066.0608	14999.474
Alps	Millan	419727.0783	14999.474
Alps	Millan	458403.9036	17473.988
Alps	Millan	550877.6053	17499.387
Alps	Millan	345411.7084	19970.271
Alps	Millan	296929.5324	22499.211
Alps	Millan	336115.5135	22499.211
Alps	Millan	429295.3934	22499.211
Alps	Millan	510949.7336	24999.124
Alps	Millan	833256.0751	24999.124
Alps	Millan	270118.3607	27499.036
Alps	Millan	695662.8135	27499.036
Alps	Millan	1713505.676	27499.036
Alps	Millan	142783.6618	34947.975
Alps	Millan	301684.6881	39940.543
Alps	Millan	635922.3126	39998.598
Alps	Millan	869187.743	39998.598
Alps	Millan	922668.5416	39998.598
Alps	Millan	946827.5762	49998.248
Alps	Millan	1456599.802	52498.16
Alps	Millan	2348440.654	74888.518
Alps	Millan	2354238.686	77497.284
Alps	Millan	1499896.688	82497.109
Alps	Millan	5276652.391	182228.73
Alps	Millan	8256789.677	202492.9
Alps	Millan	9244668.943	259990.89
Alps	OGGM	27289.0882	9985.1357
Alps	OGGM	45326.1024	9999.6495
Alps	OGGM	56961.4615	19999.299
Alps	OGGM	253632.0372	22499.211
	OGGM	120714.4642	24962.839
Alps			
Alps	OGGM	140848.1848	24999.124
Alps	OGGM	88072.0604	29998.949
Alps	OGGM	191742.7379	29998.949
Alps	OGGM	216478.3241	32498.861

Alps	OGGM	216860.8615	32498.861
Alps	OGGM	233477.7479	34998.773
Alps	OGGM	345914.4491	34998.773
Alps	OGGM	540841.9479	54998.072
Alps	OGGM	744923.4988	59997.897
Alps	OGGM	822844.9014	74888.518
Alps	OGGM	1262996.847	87496.933
Alps	OGGM	1272077.076	99996.495
Alps	OGGM	2277532.927	114995.97
Alps	OGGM	2459400.353	142288.18
Alps	OGGM	2515645.662	162494.3
Alps	OGGM	10199414.49	247491.33
Himalayas	Farinotti's Ensemble	22504.625	9999.8854
Himalayas	Farinotti's Ensemble	27539.9677	9999.8854
Himalayas	Farinotti's Ensemble	27699.878	9999.8854
Himalayas	Farinotti's Ensemble	30006.9804	9999.8854
Himalayas	Farinotti's Ensemble	30839.4904	9999.8854
Himalayas	Farinotti's Ensemble	37104.0671	9999.8854
Himalayas	Farinotti's Ensemble	37947.5633	9999.8854
Himalayas	Farinotti's Ensemble	38258.839	9999.8854
Himalayas	Farinotti's Ensemble	40889.4241	9999.8854
Himalayas	Farinotti's Ensemble	42199.2236	9999.8854
Himalayas	Farinotti's Ensemble	43278.3128	9999.8854
Himalayas	Farinotti's Ensemble	47822.9384	9999.8854
Himalayas	Farinotti's Ensemble	49966.4686	9999.8854
Himalayas	Farinotti's Ensemble	63392.8284	9999.8854
Himalayas	Farinotti's Ensemble	128754.6284	9999.8854
Himalayas	Farinotti's Ensemble	38371.1424	12499.857
Himalayas	Farinotti's Ensemble	39807.8935	12499.857
Himalayas	Farinotti's Ensemble	41463.148	12499.857
Himalayas	Farinotti's Ensemble	42493.4096	12499.857
Himalayas	Farinotti's Ensemble	52959.5983	12499.857
Himalayas	Farinotti's Ensemble	58698.058	12499.857
Himalayas	Farinotti's Ensemble	60349.6504	12499.857
Himalayas	Farinotti's Ensemble	61190.7052	12499.857
Himalayas	Farinotti's Ensemble	84706.0608	12499.857
Himalayas	Farinotti's Ensemble	91512.6234	12499.857
Himalayas	Farinotti's Ensemble	34705.4129	14999.828
Himalayas	Farinotti's Ensemble	42428.7131	14999.828
Himalayas	Farinotti's Ensemble	46065.1461	14999.828
Himalayas	Farinotti's Ensemble	56956.1346	14999.828
Himalayas	Farinotti's Ensemble	69962.5773	14999.828
Himalayas	Farinotti's Ensemble	70122.4876	14999.828
Himalayas	Farinotti's Ensemble	115495.5029	14999.828
Himalayas	Farinotti's Ensemble	123144.3411	14999.828
Himalayas	Farinotti's Ensemble	149163.3301	14999.828

Himalayas Farinotti's Ensemble 273124.3122 67499.227 Himalayas Farinotti's Ensemble 408261.9241 67499.227 Himalayas Farinotti's Ensemble 458256.4685 69999.198 Himalayas Farinotti's Ensemble 498103.4241 69999.198 Himalayas Farinotti's Ensemble 1411032.418 82499.055 Himalayas Farinotti's Ensemble 585292.3666 84999.026 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 135387.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 243484.32 162498.14 Himalayas Farinotti's Ensemble 243484.32 162498.14 Himalayas Farinotti's Ensemble 243484.32 162498.14 Himalayas Farinotti's Ensemble 27577.907 234997.31 Himalayas Farinotti's Ensemble 175177.	r			
Himalayas Farinotti's Ensemble 458256.4685 69999.198 Himalayas Farinotti's Ensemble 498103.4241 69999.198 Himalayas Farinotti's Ensemble 693064.5693 72499.169 Himalayas Farinotti's Ensemble 1411032.418 82499.055 Himalayas Farinotti's Ensemble 585292.3666 84999.026 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1488275.185 142498.37 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 577257.907 234997.31 Himalayas Farinotti's Ensemble 175175.77.91 30996.45 Himalayas Farinotti's Ensemble 15219247.02 497494.3 Himalayas Farinotti's Ensemble 1	Himalayas	Farinotti's Ensemble	273124.3122	67499.227
Himalayas Farinotti's Ensemble 498103.4241 69999.198 Himalayas Farinotti's Ensemble 693064.5693 72499.169 Himalayas Farinotti's Ensemble 1411032.418 82499.055 Himalayas Farinotti's Ensemble 585292.3666 84999.026 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1488275.185 142498.37 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 3772577.907 234997.31 Himalayas Farinotti's Ensemble 4868165.661 319996.33 Himalayas Farinotti's Ensemble 15219247.02 497494.3 Himalayas Farinotti's Ensemble 18405990.29 584993.3 Himalayas Farinotti's Ensemble 1	Himalayas	Farinotti's Ensemble	408261.9241	67499.227
Himalayas Farinotti's Ensemble 693064.5693 72499.169 Himalayas Farinotti's Ensemble 1411032.418 82499.055 Himalayas Farinotti's Ensemble 585292.3666 84999.026 Himalayas Farinotti's Ensemble 543434.0115 8998.969 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1488275.185 142498.37 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 244843.32 162498.14 Himalayas Farinotti's Ensemble 5772577.907 234997.31 Himalayas Farinotti's Ensemble 4868165.661 319996.43 Himalayas Farinotti's Ensemble 11611276.88 404995.36 Himalayas Farinotti's Ensemble 1161276.88 404995.36 Himalayas Farinotti's Ensemble 112	Himalayas	Farinotti's Ensemble	458256.4685	69999.198
Himalayas Farinotti's Ensemble 1411032.418 82499.055 Himalayas Farinotti's Ensemble 585292.3666 84999.026 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1488275.185 142498.37 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 243497.31 114998.68 Himalayas Farinotti's Ensemble 243493.32 162498.14 Himalayas Farinotti's Ensemble 243493.32 162498.14 Himalayas Farinotti's Ensemble 875577.371 309996.45 Himalayas Farinotti's Ensemble 11611276.88 404995.36 Himalayas Farinotti's Ensemble 15219247.02 497494.3 Himalayas Farinotti's Ensemble 1108924	Himalayas	Farinotti's Ensemble	498103.4241	69999.198
Himalayas Farinotti's Ensemble 585292.3666 84999.026 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 272577.907 234997.31 Himalayas Farinotti's Ensemble 875777.371 309996.45 Himalayas Farinotti's Ensemble 14611276.88 404995.36 Himalayas Farinotti's Ensemble 15219247.02 497494.3 Himalayas Farinotti's Ensemble 11089240.63 532493.9 Himalayas Farinotti's Ensemble 19254011.58 794990.89 Himalayas GlabTop2 26012.8856	Himalayas	Farinotti's Ensemble	693064.5693	72499.169
Himalayas Farinotti's Ensemble 544348.0115 89998.969 Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 27577.907 234997.31 Himalayas Farinotti's Ensemble 8755777.371 309996.45 Himalayas Farinotti's Ensemble 1468165.661 319996.33 Himalayas Farinotti's Ensemble 11611276.88 404995.36 Himalayas Farinotti's Ensemble 15219247.02 497494.3 Himalayas Farinotti's Ensemble 18405990.29 584993.3 Himalayas GlabTop2 2612.8856 9999.8854 Himalayas GlabTop2 2612.8856 999	Himalayas	Farinotti's Ensemble	1411032.418	82499.055
Himalayas Farinotti's Ensemble 503528.1666 94998.912 Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1488275.185 142498.37 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 5772577.907 234997.31 Himalayas Farinotti's Ensemble 875777.371 309996.45 Himalayas Farinotti's Ensemble 11611276.88 404995.36 Himalayas Farinotti's Ensemble 11611276.88 404994.3 Himalayas Farinotti's Ensemble 112519247.02 497494.3 Himalayas Farinotti's Ensemble 18405990.29 584993.3 Himalayas GlabTop2 17136.0341	Himalayas	Farinotti's Ensemble	585292.3666	84999.026
Himalayas Farinotti's Ensemble 923184.0032 94998.912 Himalayas Farinotti's Ensemble 1353887.076 97498.883 Himalayas Farinotti's Ensemble 1486184.145 114998.68 Himalayas Farinotti's Ensemble 1488275.185 142498.37 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 2434843.32 162498.14 Himalayas Farinotti's Ensemble 2772577.907 234997.31 Himalayas Farinotti's Ensemble 8755777.371 30996.45 Himalayas Farinotti's Ensemble 1468165.661 319996.33 Himalayas Farinotti's Ensemble 15219247.02 497494.3 Himalayas Farinotti's Ensemble 11089240.63 532493.9 Himalayas Farinotti's Ensemble 19254011.58 794990.89 Himalayas GlabTop2 17136.0341 9999.8854 Himalayas GlabTop2 26371.7682 9999.8854 Himalayas GlabTop2 31205.6972 9999.8854 <td>Himalayas</td> <td>Farinotti's Ensemble</td> <td>544348.0115</td> <td>89998.969</td>	Himalayas	Farinotti's Ensemble	544348.0115	89998.969
HimalayasFarinotti's Ensemble1353887.07697498.883HimalayasFarinotti's Ensemble1486184.145114998.68HimalayasFarinotti's Ensemble1488275.185142498.37HimalayasFarinotti's Ensemble2434843.32162498.14HimalayasFarinotti's Ensemble4640924.612202497.68HimalayasFarinotti's Ensemble5772577.907234997.31HimalayasFarinotti's Ensemble8755777.37130996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop235630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop25630.4662999.8854HimalayasGlabTop25630.67279999.8854Himala	Himalayas	Farinotti's Ensemble	503528.1666	94998.912
HimalayasFarinotti's Ensemble1486184.145114998.68HimalayasFarinotti's Ensemble1488275.185142498.37HimalayasFarinotti's Ensemble2434843.32162498.14HimalayasFarinotti's Ensemble4640924.612202497.68HimalayasFarinotti's Ensemble5772577.907234997.31HimalayasFarinotti's Ensemble8755777.371309996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.8840495.36HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854Hima	Himalayas	Farinotti's Ensemble	923184.0032	94998.912
HimalayasFarinotti's Ensemble1488275.185142498.37HimalayasFarinotti's Ensemble2434843.32162498.14HimalayasFarinotti's Ensemble4640924.612202497.68HimalayasFarinotti's Ensemble5772577.907234997.31HimalayasFarinotti's Ensemble8755777.371309996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble115219247.02497494.3HimalayasFarinotti's Ensemble11809240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop231205.6972999.8854HimalayasGlabTop233047.7171999.8854HimalayasGlabTop233047.7171999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop256300.7279999.8854HimalayasGlabTop25630.607279999.8854HimalayasGlabTop25630.607279999.8854HimalayasGla	Himalayas	Farinotti's Ensemble	1353887.076	97498.883
HimalayasFarinotti's Ensemble2434843.32162498.14HimalayasFarinotti's Ensemble4640924.612202497.68HimalayasFarinotti's Ensemble5772577.907234997.31HimalayasFarinotti's Ensemble8755777.371309996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble115219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.7171999.8854HimalayasGlabTop233047.7171999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.67279999.8854HimalayasGlabTop256306.7279999.8854HimalayasGlabTop263856.6903999.8854HimalayasGlabTop273500.1345999.8854HimalayasGlabTop2 <td< td=""><td>Himalayas</td><td>Farinotti's Ensemble</td><td>1486184.145</td><td>114998.68</td></td<>	Himalayas	Farinotti's Ensemble	1486184.145	114998.68
HimalayasFarinotti's Ensemble4640924.612202497.68HimalayasFarinotti's Ensemble5772577.907234997.31HimalayasFarinotti's Ensemble8755777.371309996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.7171999.8854HimalayasGlabTop233611.6755999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop255630.4662999.8854HimalayasGlabTop263856.6903 <td>Himalayas</td> <td>Farinotti's Ensemble</td> <td>1488275.185</td> <td>142498.37</td>	Himalayas	Farinotti's Ensemble	1488275.185	142498.37
HimalayasFarinotti's Ensemble5772577.907234997.31HimalayasFarinotti's Ensemble8755777.371309996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble115219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.496999	Himalayas	Farinotti's Ensemble	2434843.32	162498.14
HimalayasFarinotti's Ensemble8755777.371309996.45HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.67279999.8854HimalayasGlabTop260437.54999.8854HimalayasGlabTop263856.6903999.8854HimalayasGlabTop273500.1345999.8854HimalayasGlabTop282855.496999.8854HimalayasGlabTop28285.496999.8854HimalayasGlabTop289524.1208999.8854	Himalayas	Farinotti's Ensemble	4640924.612	202497.68
HimalayasFarinotti's Ensemble4868165.661319996.33HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop25630.46629999.8854HimalayasGlabTop25630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop2825785.159999.8854HimalayasGlabTop285785.159999.8854Himalay	Himalayas	Farinotti's Ensemble	5772577.907	234997.31
HimalayasFarinotti's Ensemble5471315.489337496.13HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282785.15999.8854HimalayasGlabTop285785.15999.8854HimalayasGlabTop289524.1208999.8854HimalayasGlabT	Himalayas	Farinotti's Ensemble	8755777.371	309996.45
HimalayasFarinotti's Ensemble11611276.88404995.36HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop28978	Himalayas	Farinotti's Ensemble	4868165.661	319996.33
HimalayasFarinotti's Ensemble15219247.02497494.3HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.15 <td>Himalayas</td> <td>Farinotti's Ensemble</td> <td>5471315.489</td> <td>337496.13</td>	Himalayas	Farinotti's Ensemble	5471315.489	337496.13
HimalayasFarinotti's Ensemble11089240.63532493.9HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop28285.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	Farinotti's Ensemble	11611276.88	404995.36
HimalayasFarinotti's Ensemble18405990.29584993.3HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	Farinotti's Ensemble	15219247.02	497494.3
HimalayasFarinotti's Ensemble19254011.58794990.89HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.159999.8854HimalayasGlabTop28285.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.2371999.8854	Himalayas	Farinotti's Ensemble	11089240.63	532493.9
HimalayasGlabTop217136.03419999.8854HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	Farinotti's Ensemble	18405990.29	584993.3
HimalayasGlabTop226012.88569999.8854HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.1208999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	Farinotti's Ensemble	19254011.58	794990.89
HimalayasGlabTop226371.76829999.8854HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.2371999.8854	Himalayas	GlabTop2	17136.0341	9999.8854
HimalayasGlabTop231205.69729999.8854HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	26012.8856	9999.8854
HimalayasGlabTop231422.97989999.8854HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	26371.7682	9999.8854
HimalayasGlabTop233047.71719999.8854HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	31205.6972	9999.8854
HimalayasGlabTop233611.67559999.8854HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.2371999.8854	Himalayas	GlabTop2	31422.9798	9999.8854
HimalayasGlabTop244982.39489999.8854HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	33047.7171	9999.8854
HimalayasGlabTop255630.46629999.8854HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	33611.6755	9999.8854
HimalayasGlabTop255767.18339999.8854HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	44982.3948	9999.8854
HimalayasGlabTop256306.72799999.8854HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	55630.4662	9999.8854
HimalayasGlabTop260437.549999.8854HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	55767.1833	9999.8854
HimalayasGlabTop263856.69039999.8854HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	56306.7279	9999.8854
HimalayasGlabTop273500.13459999.8854HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	60437.54	9999.8854
HimalayasGlabTop282855.4969999.8854HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	63856.6903	9999.8854
HimalayasGlabTop285785.159999.8854HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	73500.1345	9999.8854
HimalayasGlabTop289524.12089999.8854HimalayasGlabTop291820.23719999.8854	Himalayas	GlabTop2	82855.496	9999.8854
Himalayas GlabTop2 91820.2371 9999.8854	Himalayas	GlabTop2	85785.15	9999.8854
	Himalayas	GlabTop2	89524.1208	9999.8854
Himplayas ClabTon? 107000 602 0000 8854	Himalayas	GlabTop2	91820.2371	9999.8854
ן דעטרטאט אטאראט אטערטע אטאראט אין אטאראטא אוויויז אוויויז אין אין אראטאט אין אין ארא אין אין ארא אין אין ארא א	Himalayas	GlabTop2	107099.603	9999.8854
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Himalayas	GlabTop2	32549.6759	12499.857
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Himalayas	GlabTop2	49390.3033	12499.857
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Himalayas	GlabTop2	139696.8858	12499.857
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Himalayas	GlabTop2	195706.9863	19999.771
Himalayas	GlabTop2	284361.9764	19999.771
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Himalayas	GlabTop2	218265.3216	22499.742
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Himalayas	GlabTop2	286201.555	24999.714
Himalayas	GlabTop2	147340.8412	27499.685
Himalayas	GlabTop2	204354.3482	27499.685
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Himalayas	GlabTop2	236525.6104	27499.685
Himalayas	GlabTop2	256760.9743	27499.685
Himalayas	GlabTop2	320961.8989	27499.685
Himalayas	GlabTop2	358075.7315	27499.685
Himalayas	GlabTop2	114131.9932	29999.656
Himalayas	GlabTop2	535186.7395	29999.656
Himalayas	GlabTop2	87191.3839	32499.628
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Himalayas	GlabTop2	286250.3825	32499.628
Himalayas	GlabTop2	322882.0429	32499.628
Himalayas	GlabTop2	246564.5579	34999.599
Himalayas	GlabTop2	931180.7377	34999.599
Himalayas	GlabTop2	1630285.277	37499.57
Himalayas	GlabTop2	181385.8613	39999.542
Himalayas	GlabTop2	454640.7873	39999.542
Himalayas	GlabTop2	709277.7625	42499.513
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Himalayas	GlabTop2	364137.6738	47499.456
Himalayas	GlabTop2	477158.8398	47499.456
Himalayas	GlabTop2	833375.4618	47499.456
Himalayas	GlabTop2	1144482.737	47499.456
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Himalayas	GlabTop2	384949.2029	54999.37
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Himalayas	GlabTop2	1344014.143	54999.37
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Himalayas	GlabTop2	686515.5721	57499.341
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Himalayas	GlabTop2	2758842.904	117498.65
Himalayas	GlabTop2	1442717.846	122498.6
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Himalayas	GlabTop2	2707208.974	139998.4
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Himalayas	GlabTop2	9861427.545	392495.5
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Himalayas	GlabTop2	44045849.13	667492.35
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Himalayas	GlabTop2	25037900.4	737491.55
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Himalayas	HF	96314.8145	17499.8
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Himalayas	HF	153292.9215	17499.8
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Himalayas	HF	2663863.523	177497.97
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Himalayas	HF	2660036.663	227497.39
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Himalayas	HF	6261246.624	334996.16
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Himalayas	Millan	154514.8313	9999.8854
Himalayas	Millan	157637.3541	9999.8854
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Himalayas	Millan	177939.8559	9999.8854
Himalayas	Millan	216783.405	9999.8854
Himalayas	Millan	217462.1082	9999.8854
Himalayas	Millan	225112.167	9999.8854
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Himalayas	Millan	423282.504	9999.8854
Himalayas	Millan	494337.5981	9999.8854
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innulayas	Ivinian	51/40.0/4/	12755.057

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Himalayas	Millan	65510.7241	12499.857
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Himalayas	Millan	128237.0562	12499.857
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Himalayas	Millan	222892.9541	12499.857
Himalayas	Millan	223012.5817	12499.857
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Himalayas	Millan	239548.0368	12499.857
Himalayas	Millan	285171.2933	12499.857
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Himalayas	Millan	418747.6438	12499.857
Himalayas	Millan	505613.1037	12499.857
Himalayas	Millan	507058.3996	12499.857
Himalayas	Millan	560502.6115	12499.857
Himalayas	Millan	615198.0298	12499.857
Himalayas	Millan	50165.4409	14999.828
Himalayas	Millan	74404.6651	14999.828
Himalayas	Millan	75336.0509	14999.828
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, Himalayas	Millan	129366.1936	14999.828
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Himalayas	Millan	157283.3543	14999.828
Himalayas	Millan	171761.9481	14999.828
Himalayas	Millan	186642.1488	14999.828
Himalayas	Millan	192882.3116	14999.828
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		273030.7370	17333.020

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Himalayas	Millan	380569.3703	14999.828
Himalayas	Millan	412724.7636	14999.828
Himalayas	Millan	504867.2626	14999.828
Himalayas	Millan	582473.7954	14999.828
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Himalayas	Millan	1092183.532	14999.828
Himalayas	Millan	63851.8075	17499.8
Himalayas	Millan	64670.8899	17499.8
Himalayas	Millan	142727.857	17499.8
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Himalayas	Millan	256582.7537	17499.8
Himalayas	Millan	291958.325	17499.8
Himalayas	Millan	303647.6442	17499.8
Himalayas	Millan	340613.7734	17499.8
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Himalayas	Millan	451060.506	17499.8
Himalayas	Millan	473040.2347	17499.8
Himalayas	Millan	527365.7842	17499.8
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Himalayas	Millan	198490.1575	19999.771
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Himalayas	Millan	455507.4766	19999.771
Himalayas	Millan	525943.6813	19999.771
Himalayas	Millan	718259.5932	19999.771
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Himalayas	Millan	179634.1724	22499.742
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Himalayas	Millan	254939.7061	22499.742
Himalayas	Millan	279977.2611	22499.742
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Himalayas	Millan	377898.5025	22499.742
Himalayas	Millan	390370.2834	22499.742
Himalayas	Millan	458003.7859	22499.742
Himalayas	Millan	467653.3335	22499.742
Himalayas	Millan	482048.9205	22499.742
Himalayas	Millan	528802.5353	22499.742
Himalayas	Millan	559999.6876	22499.742
Himalayas	Millan	562129.7902	22499.742
Himalayas	Millan	674555.2599	22499.742
Himalayas	Millan	792886.4237	22499.742
Himalayas	Millan	224142.9398	24999.714
Himalayas	Millan	603149.828	24999.714
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Himalayas	Millan	1296897.983	24999.714
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Himalayas	Millan	187561.3277	27499.685
Himalayas	Millan	195883.9862	27499.685
Himalayas	Millan	212305.9172	27499.685
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Himalayas	Millan	253062.2862	27499.685
Himalayas	Millan	342686.5035	27499.685
Himalayas	Millan	385492.4096	27499.685
Himalayas	Millan	422899.2076	27499.685
Himalayas	Millan	423070.1041	27499.685
Himalayas	Millan	492541.9644	27499.685
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Himalayas	Millan	1042496.601	29999.656
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Himalayas	Millan	227762.2831	32499.628
Himalayas	Millan	322044.6502	32499.628
Himalayas	Millan	618879.6283	32499.628
Himalayas	Millan	695043.3064	32499.628
Himalayas	Millan	757556.018	32499.628
-	Millan		
Himalayas		783267.3933	32499.628
Himalayas	Millan	1073633.94	32499.628

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Himalayas	Millan	1406806.392	32499.628
Himalayas	Millan	1652257.681	32499.628
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Himalayas	Millan	1232136.762	37499.57
Himalayas	Millan	1792530.733	37499.57
Himalayas	Millan	2094827.074	37499.57
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Himalayas	Millan	1592906.555	44999.484
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, Himalayas	Millan	1471351.551	47499.456
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Himalayas	Millan	807914.3277	57499.341
Himalayas	Millan	1523295.536	57499.341
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Himalayas	Millan	443013.7232	59999.313
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Himalayas	Millan	682079.5878	72499.169
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Himalayas	Millan	2234018.594	82499.055
Himalayas	Millan	4850873.379	87498.998
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	Millan	3709821.999	94998.912
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Himalayas	Millan	1339522.007	137498.42
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Himalayas	Millan	9167663.521	167498.08
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Himalayas	Millan	3941965.335	194997.77
Himalayas	Millan	7915484.264	209997.59
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Himalayas	Millan	23612136.7	212497.57
Himalayas	Millan	19435962.63	214997.54
Himalayas	Millan	13977961.8	269996.91
Himalayas	Millan	10405796.31	287496.71
Himalayas	Millan	6388436.329	329996.22
Himalayas	Millan	6673212.119	344996.05
Himalayas	Millan	9728431.022	364995.82
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Himalayas	Millan	19523637.4	489994.39
Himalayas	Millan	23542684.37	494994.33
Himalayas	Millan	23941293.09	559993.58

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Himalayas	OGGM	19860.6123	9999.8854
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Himalayas	OGGM	32094.3589	9999.8854
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Himalayas	OGGM	51102.9301	9999.8854
Himalayas	OGGM	67417.4405	9999.8854
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Himalayas	OGGM	58059.6375	14999.828
Himalayas	OGGM	59636.7679	14999.828
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Himalayas	OGGM	75039.4235	14999.828
Himalayas	OGGM	76695.8986	14999.828
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Himalayas	OGGM	66139.379	17499.8
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Himalayas	OGGM	74382.6927	17499.8
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Himalayas	OGGM	90136.9068	17499.8
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Himalayas	OGGM	107163.0789	17499.8
Himalayas	OGGM	129443.0971	17499.8
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Himalayas	OGGM	262001.3928	17499.8
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Himalayas	OGGM	226606.2905	22499.742
Himalayas	OGGM	86662.8255	24999.714
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Himalayas	OGGM	417152.2031	24999.714
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Himalayas	OGGM	104448.2662	27499.685
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Himalayas	OGGM	336550.0992	27499.685
Himalayas	OGGM	403169.209	27499.685
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Himalayas	OGGM	319885.2511	29999.656
Himalayas	OGGM	359243.931	29999.656
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Himalayas	OGGM	1462594.327	29999.656
Himalayas	OGGM	155204.5207	32499.628
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Himalayas	OGGM	201014.5427	34999.599
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Himalayas	OGGM	244125.621	34999.599
Himalayas	OGGM	293789.3587	34999.599
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Himalayas	OGGM	166840.1295	37499.57
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Himalayas	OGGM	389976.0008	39999.542
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Himalayas	OGGM	760158.5272	42499.513
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. Innulu yus	00011	551,54.5445	17333.707

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Himalayas	OGGM	708583.1904	44999.484
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Himalayas	OGGM	1646117.615	49999.427
Himalayas	OGGM	434356.5959	52499.399
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Himalayas	OGGM	737928.5573	54999.37
Himalayas	OGGM	914941.9101	59999.313
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Himalayas	OGGM	548229.8029	62499.284
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Himalayas	OGGM	804703.9153	69999.198
Himalayas	OGGM	815975.7588	72499.169
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Himalayas	OGGM	1169354.279	74999.141
Himalayas	OGGM	1282215.534	74999.141
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Himalayas	OGGM	1573827.184	79999.084
Himalayas	OGGM	1098045.281	82499.055
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Himalayas	OGGM	935614.2807	89998.969
Himalayas	OGGM	2727909.42	92498.94
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Himalayas	OGGM	5215252.749	117498.65
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Himalayas	OGGM	2839819.759	124998.57
Himalayas	OGGM	2279781.01	134998.45
Himalayas	OGGM	2439389.777	134998.45
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Himalayas	OGGM	4232212.254	169998.05
Himalayas	OGGM	5976910.282	172498.02
Himalayas	OGGM	7223241.218	172498.02
Himalayas	OGGM	4426650.944	177497.97
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Himalayas	OGGM	6436826.888	182497.91
Himalayas	OGGM	7315102.959	184997.88
Himalayas	OGGM	10062626.41	189997.82
Himalayas	OGGM	4420089.74	202497.68
Himalayas	OGGM	7863397.458	219997.48
Himalayas	OGGM	4238346.217	224997.42
Himalayas	OGGM	12444342.29	227497.39
Himalayas	OGGM	5253064.816	237497.28
Himalayas	OGGM	6946364.507	242497.22
Himalayas	OGGM	3157007.287	244997.19
Himalayas	OGGM	8116541.53	244997.19
Himalayas	OGGM	12921746.49	244997.19
Himalayas	OGGM	11688060.67	254997.08
Himalayas	OGGM	11068798.97	269996.91
Himalayas	OGGM	8052808.13	274996.85
Himalayas	OGGM	5083982.28	294996.62
Himalayas	OGGM	15265203.52	297496.59
Himalayas	OGGM	5590597.57	299996.56
Himalayas	OGGM	14174293.17	329996.22
Himalayas	OGGM	6396874.953	362495.85
Himalayas	OGGM	11675579.12	364995.82
Himalayas	OGGM	11327847.51	394995.47
Himalayas	OGGM	11777358.96	467494.64
Himalayas	OGGM	29851141.15	502494.24
Himalayas	OGGM	25830116.66	534993.87
Himalayas	OGGM	14996323.55	547493.73
Himalayas	OGGM	22155871.65	577493.38
Himalayas	OGGM	33899496.87	762491.26
Himalayas	OGGM	73921246.59	897489.72
Himalayas	OGGM	94019289.64	1732480.2
Peruvian	Farinotti's Ensemble	17590.7626	9996.7761
Andes		1,000,010	555617761
Peruvian	Farinotti's Ensemble	24356.1589	9996.7761
Andes			
Peruvian	Farinotti's Ensemble	26599.0879	9996.7761
Andes			
Peruvian	Farinotti's Ensemble	28835.9154	9996.7761
Andes			
Peruvian	Farinotti's Ensemble	34798.348	9996.7761
Andes			
Peruvian	Farinotti's Ensemble	36730.0981	9996.7761
Andes			
Peruvian	Farinotti's Ensemble	37005.888	9996.7761
Andes			
Peruvian	Farinotti's Ensemble	39916.3264	9996.7761
Andes			
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Peruvian Andes	Farinotti's Ensemble	44261.8488	9996.7761
Peruvian Andes	Farinotti's Ensemble	44926.9175	9996.7761
Peruvian	Farinotti's Ensemble	50177.9097	9996.7761
Peruvian	Farinotti's Ensemble	50362.1764	9996.7761
Peruvian Andes	Farinotti's Ensemble	51293.2726	9996.7761
Peruvian Andes	Farinotti's Ensemble	56846.9015	9996.7761
Peruvian Andes	Farinotti's Ensemble	62135.7232	9996.7761
Peruvian Andes	Farinotti's Ensemble	73268.6076	9996.7761
Peruvian Andes	Farinotti's Ensemble	94637.4486	9996.7761
Peruvian Andes	Farinotti's Ensemble	180123.7957	9996.7761
Peruvian	Farinotti's Ensemble	35057.0536	12495.97
Andes Peruvian	Farinotti's Ensemble	37690.4817	12495.97
Andes Peruvian	Farinotti's Ensemble	40369.0612	12495.97
Andes Peruvian	Farinotti's Ensemble	48326.7	12495.97
Andes Peruvian	Farinotti's Ensemble	56599.1787	12495.97
Andes Peruvian	Farinotti's Ensemble	66702.1217	12495.97
Andes Peruvian	Farinotti's Ensemble	70482.6408	12495.97
Andes Peruvian	Farinotti's Ensemble	143950.1588	12495.97
Andes Peruvian	Farinotti's Ensemble	443405.5878	12495.97
Andes Peruvian	Farinotti's Ensemble	49287.0837	14995.164
Andes Peruvian	Farinotti's Ensemble	57865.86	14995.164
Andes Peruvian	Farinotti's Ensemble	61353.5048	14995.164
Andes Peruvian	Farinotti's Ensemble	65595.3009	14995.164
Andes			

Peruvian Andes	Farinotti's Ensemble	66472.7035	14995.164
Peruvian Andes	Farinotti's Ensemble	83043.2873	14995.164
Peruvian	Farinotti's Ensemble	132341.3537	14995.164
Peruvian	Farinotti's Ensemble	328766.0448	14995.164
Andes Peruvian	Farinotti's Ensemble	49997.3038	17494.358
Andes Peruvian	Farinotti's Ensemble	92327.4026	17494.358
Andes Peruvian	Farinotti's Ensemble	117213.1759	17494.358
Andes Peruvian	Farinotti's Ensemble	118211.3891	17494.358
Andes Peruvian	Farinotti's Ensemble	80893.1018	19993.552
Andes Peruvian	Farinotti's Ensemble	82613.7383	19993.552
Andes Peruvian	Farinotti's Ensemble	111251.9636	19993.552
Andes Peruvian	Farinotti's Ensemble	108642.9417	22492.746
Andes Peruvian	Farinotti's Ensemble	118429.8245	22492.746
Andes Peruvian	Farinotti's Ensemble	120869.2234	22492.746
Andes Peruvian	Farinotti's Ensemble	148481.1683	22492.746
Andes Peruvian	Farinotti's Ensemble	100959.8726	24991.94
Andes Peruvian	Farinotti's Ensemble	155192.871	24991.94
Andes	Farinotti's Ensemble	488738.8684	24991.94
Andes			
Peruvian Andes	Farinotti's Ensemble	152763.2346	27491.134
Peruvian Andes	Farinotti's Ensemble	187893.5068	27491.134
Peruvian Andes	Farinotti's Ensemble	373912.6181	27491.134
Peruvian Andes	Farinotti's Ensemble	207309.8525	32489.522
Peruvian Andes	Farinotti's Ensemble	315376.8081	34988.716

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Peruvian Andes	Farinotti's Ensemble	443353.1145	37487.91
Peruvian	Farinotti's Ensemble	441088.2199	42486.298
Andes			
Peruvian	Farinotti's Ensemble	215102.7495	44985.492
Andes			
Peruvian Andes	Farinotti's Ensemble	254992.2291	47484.686
Peruvian	Farinotti's Ensemble	556459.9486	47484.686
Andes		550+55.5+00	47404.000
Peruvian	Farinotti's Ensemble	171459.5977	49983.88
Andes		1/1-35.55//	45565.66
Peruvian	Farinotti's Ensemble	1144416.088	49983.88
Andes		11111101000	15566166
Peruvian	Farinotti's Ensemble	511347.544	54982.269
Andes			
Peruvian	Farinotti's Ensemble	643392.3625	59980.657
Andes			
Peruvian	Farinotti's Ensemble	628040.868	69977.433
Andes			
Peruvian	Farinotti's Ensemble	633923.9805	69977.433
Andes			
Peruvian	Farinotti's Ensemble	1071908.953	69977.433
Andes			
Peruvian	Farinotti's Ensemble	651033.9411	79974.209
Andes			70074.000
Peruvian	Farinotti's Ensemble	885544.4942	79974.209
Andes Peruvian	Farinotti's Ensemble	542993.8324	84972.597
Andes	Familiotti s Ensemble	542995.6524	64972.597
Peruvian	Farinotti's Ensemble	747771.5426	84972.597
Andes		/ 4/ / 1.5420	04372.337
Peruvian	Farinotti's Ensemble	983066.7547	84972.597
Andes			
Peruvian	Farinotti's Ensemble	3023152.245	87471.791
Andes			
Peruvian	Farinotti's Ensemble	1280868.664	92470.179
Andes			
Peruvian	Farinotti's Ensemble	862181.6673	97468.567
Andes			
Peruvian	Farinotti's Ensemble	2684703.044	99967.761
Andes			
Peruvian	Farinotti's Ensemble	1499664.071	107465.34
Andes		2005742.44	107420 55
Peruvian	Farinotti's Ensemble	2805743.11	187439.55
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Peruvian Andes	Farinotti's Ensemble	3046729.847	189938.75
Peruvian Andes	Farinotti's Ensemble	2132794.85	207433.1
Peruvian	GlabTop2	26027.9831	9996.7761
Andes			
Peruvian	GlabTop2	28823.7123	9996.7761
Andes			
Peruvian	GlabTop2	31200.8754	9996.7761
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Peruvian	GlabTop2	37138.9018	9996.7761
Andes			
Peruvian	GlabTop2	37984.5763	9996.7761
Andes			
Peruvian	GlabTop2	40508.1765	9996.7761
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Peruvian	GlabTop2	47500.5504	9996.7761
Andes		40544.0640	0000 7764
Peruvian	GlabTop2	49514.0612	9996.7761
Andes		E 4722 405	0000 7764
Peruvian	GlabTop2	54732.105	9996.7761
Andes Peruvian	ClabTan2	74702 0117	9996.7761
Andes	GlabTop2	74783.0117	9990.7701
Peruvian	GlabTop2	84050.0427	9996.7761
Andes		04050.0427	5550.7701
Peruvian	GlabTop2	105522.6101	9996.7761
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Peruvian	GlabTop2	114385.7186	9996.7761
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Peruvian	GlabTop2	124186.0248	9996.7761
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Peruvian	GlabTop2	36343.2599	12495.97
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Peruvian	GlabTop2	40051.7808	12495.97
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Peruvian	GlabTop2	45821.4045	12495.97
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Peruvian	GlabTop2	52393.9919	12495.97
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Peruvian	GlabTop2	56604.0599	12495.97
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Peruvian	GlabTop2	67416.0028	12495.97
Andes	ClahTan2	00427 2022	12405.07
Peruvian	GlabTop2	88427.2932	12495.97
Andes			

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Peruvian Andes	GlabTop2	53776.6026	14995.164
Peruvian Andes	GlabTop2	71450.3463	14995.164
Peruvian	GlabTop2	77291.9682	14995.164
Andes Peruvian	GlabTop2	80507.484	14995.164
Andes			
Peruvian Andes	GlabTop2	98315.4617	14995.164
Peruvian Andes	GlabTop2	125812.6975	14995.164
Peruvian	GlabTop2	132863.6462	14995.164
Andes Peruvian	GlabTop2	152265.3483	14995.164
Andes			
Peruvian Andes	GlabTop2	227346.1156	14995.164
Peruvian Andes	GlabTop2	65273.1392	17494.358
Peruvian	GlabTop2	95213.4347	17494.358
Andes			
Peruvian Andes	GlabTop2	181353.8678	17494.358
Peruvian	GlabTop2	90520.1241	19993.552
Andes Peruvian	GlabTop2	131847.1284	19993.552
Andes	GlabTopz	131047.1204	19995.552
Peruvian	GlabTop2	255333.9158	19993.552
Andes			
Peruvian Andes	GlabTop2	128602.3252	22492.746
Peruvian	GlabTop2	225522.973	22492.746
Andes Peruvian	ClahTan2	165242 1957	27401 124
Andes	GlabTop2	165342.1857	27491.134
Peruvian	GlabTop2	222406.3024	32489.522
Andes Peruvian	GlabTop2	337359.4649	32489.522
Andes		557555.4043	52703.322
Peruvian Andes	GlabTop2	267711.5159	34988.716
Andes Peruvian	GlabTop2	486879.1166	37487.91
Andes			
Peruvian Andes	GlabTop2	448925.048	39987.104
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Peruvian Andes	GlabTop2	226362.546	44985.492
Peruvian Andes	GlabTop2	574752.3892	44985.492
Peruvian Andes	GlabTop2	789656.2283	44985.492
Peruvian	GlabTop2	459957.867	47484.686
Andes Peruvian	GlabTop2	992158.0611	49983.88
Andes Peruvian	GlabTop2	354080.1468	57481.463
Andes Peruvian	GlabTop2	1366696.698	72476.627
Andes Peruvian	GlabTop2	1457962.431	74975.821
Andes			
Peruvian Andes	GlabTop2	3315290.698	74975.821
Peruvian Andes	GlabTop2	2055236.854	87471.791
Peruvian Andes	GlabTop2	1489395.165	99967.761
Peruvian	GlabTop2	714569.3595	102466.95
Andes Peruvian	GlabTop2	1434245.715	117462.12
Andes Peruvian	GlabTop2	2224508.437	132457.28
Andes Peruvian	GlabTop2	2297264.514	134956.48
Andes Peruvian		4119992.123	149951.64
Andes	GlabTop2		
Peruvian Andes	GlabTop2	11900308.94	319896.83
Peruvian Andes	HF	25209.1553	9996.7761
Peruvian Andes	HF	30458.9271	9996.7761
Peruvian	HF	39686.9082	9996.7761
Andes Peruvian	HF	43949.4496	9996.7761
Andes Peruvian	HF	45563.9191	9996.7761
Andes Peruvian	HF	52442.8042	9996.7761
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Peruvian Andes	HF	56832.2578	9996.7761
Peruvian	HF	57284.9927	9996.7761
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Peruvian Andes	HF	59313.1472	9996.7761
Peruvian	HF	64722.7796	9996.7761
Andes		04722.7750	5550.7701
Peruvian	HF	68506.9596	9996.7761
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Peruvian	HF	69669.9146	9996.7761
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Peruvian	HF	71194.0813	9996.7761
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Peruvian	HF	78742.9163	9996.7761
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Peruvian	HF	82694.2788	9996.7761
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Peruvian	HF	18713.4474	12495.97
Andes		10/13.44/4	12433.37
Peruvian	HF	31685.3383	12495.97
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Peruvian	HF	40987.7582	12495.97
Andes		40987.7582	12495.97
Peruvian	HF	42969.541	12495.97
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		43263.6356	12405 07
Peruvian	HF	43203.0350	12495.97
Andes		44910 5202	12405 07
Peruvian	HF	44819.5303	12495.97
Andes		50700 2227	12405.07
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Peruvian	HF	63157.1224	12495.97
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Peruvian	HF	81807.1137	12495.97
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Peruvian	HF	109990.1635	12495.97
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Peruvian	HF	76262.027	14995.164
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Peruvian	HF	80572.1604	14995.164
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Peruvian	HF	87662.1591	14995.164
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Peruvian Andes	HF	92423.807	14995.164
Peruvian	HF	96190.9027	14995.164
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Andes			
Peruvian	HF	118479.8572	14995.164
Andes Peruvian	HF	66087.0857	17494.358
Andes		00007.0037	17-13-1.550
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Peruvian Andes	HF	102769.5917	17494.358
Peruvian	HF	116111.2363	17494.358
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Peruvian	HF	113038.4968	19993.552
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Peruvian	HF	116958.1312	19993.552
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Peruvian Andes	HF	143170.381	19993.552
Peruvian	HF	149166.9822	19993.552
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Peruvian	HF	159861.7754	19993.552
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Peruvian	HF	204955.8754	19993.552
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Peruvian	HF	282815.2876	19993.552
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Peruvian	HF	72643.809	22492.746
Andes		05002.0100	22402 746
Peruvian Andes	HF	85682.8169	22492.746
Peruvian	HF	96172.5981	22492.746
Andes		50172.5501	22432.740
Peruvian	HF	106453.7063	22492.746
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Peruvian	HF	119580.5765	22492.746
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Peruvian	HF	196028.0905	22492.746
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Peruvian	HF	203814.8859	22492.746
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Peruvian Andes		240581.5933	22492.746
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Peruvian Andes	HF	266774.3181	22492.746
Peruvian	HF	451344.9219	22492.746
Andes Peruvian	HF	86789.6377	24991.94
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Peruvian Andes	HF	157253.9738	24991.94
Peruvian	HF	181990.8694	24991.94
Andes		181990.8094	24991.94
Peruvian	HF	276928.5141	24991.94
Andes		270928.5141	24991.94
Peruvian	HF	148389.645	27491.134
Andes		140303.045	27451.154
Peruvian	HE	190608.6956	27491.134
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Peruvian	HF	243966.732	27491.134
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Peruvian	HF	466447.4733	27491.134
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Peruvian	HF	149841.8134	29990.328
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Peruvian	HF	194651.5813	29990.328
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Peruvian	HF	279915.832	29990.328
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Peruvian	HF	378350.884	29990.328
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Peruvian	HF	442323.1732	32489.522
Andes		450207 6500	22400 522
Peruvian Andes	HF	450307.6588	32489.522
Peruvian	HF	164779.623	34988.716
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Peruvian	HF	224288.0197	34988.716
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Peruvian	HF	362245.2382	34988.716
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Peruvian	HF	270979.5049	37487.91
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Peruvian	HF	296911.0835	37487.91
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Peruvian	HF	315378.0284	39987.104
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Peruvian	HF	303077.3078	47484.686
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Peruvian Andes	HF	425767.2331	47484.686
Peruvian	HF	456050.4357	49983.88
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Peruvian	HF	510564.1053	49983.88
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Peruvian	HF	410253.4374	52483.075
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Peruvian	HF	574254.5029	52483.075
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Peruvian	HF	926121.0081	64979.045
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Peruvian	HF	1109904.512	72476.627
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Peruvian	HF	1118275.836	102466.95
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Peruvian	HF	2384476.379	117462.12
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Peruvian	HF	2382976.619	119961.31
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Peruvian	HF	2235928.094	154950.03
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Peruvian	HF	3686499.11	232425.04
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Peruvian	HF	5266414.399	254917.79
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Peruvian	Millan	37701.4645	9996.7761
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Peruvian	Millan	37824.7158	9996.7761
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Peruvian	Millan	42613.2106	9996.7761
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Peruvian	Millan	44070.2602	9996.7761
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Peruvian	Millan	45703.0344	9996.7761
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Peruvian	Millan	45867.7762	9996.7761
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Peruvian	Millan	50663.5929	9996.7761
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Peruvian	Millan	62437.1397	9996.7761
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Peruvian	Millan	63680.6352	9996.7761
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Peruvian Millan 97115.8974 9996.776	71
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Peruvian Millan 107317.6855 9996.776	61
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Peruvian Millan 118579.9226 9996.776	61
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Peruvian Millan 253641.3464 9996.776	61
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Peruvian Millan 136468.4408 12495.97	/
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Peruvian Millan 193414.1873 12495.97	7
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Peruvian Millan 93496.4592 14995.16	54
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Peruvian Millan 102137.4713 14995.16	54
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Peruvian Millan 144322.3532 14995.16	64
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Peruvian Andes	Millan	148584.8946	14995.164
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Peruvian Andes	Millan	189720.3102	14995.164
Peruvian	Millan	198589.5203	14995.164
Andes	willian	198389.3203	14995.104
Peruvian	Millan	212034.8912	14995.164
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Peruvian	Millan	219517.8296	14995.164
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Peruvian	Millan	222071.9375	14995.164
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Peruvian	Millan	294294.7398	14995.164
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Peruvian	Millan	641912.127	14995.164
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Peruvian	Millan	113504.6551	17494.358
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Peruvian	Millan	151279.3381	17494.358
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Peruvian	Millan	158802.5467	17494.358
Andes			
Peruvian	Millan	174428.6109	17494.358
Andes			
Peruvian	Millan	427944.2654	17494.358
Andes		622462 7025	47404 250
Peruvian	Millan	633462.7035	17494.358
Andes Peruvian	Millan	118997.2685	19993.552
Andes	ivinan	118997.2085	19993.332
Peruvian	Millan	144906.8815	19993.552
Andes		111300.0013	199999.592
Peruvian	Millan	150035.8427	19993.552
Andes			
Peruvian	Millan	163027.2585	19993.552
Andes			
Peruvian	Millan	215236.9836	19993.552
Andes			
Peruvian	Millan	235686.9315	19993.552
Andes			
Peruvian	Millan	328007.0123	19993.552
Andes			
Peruvian	Millan	828931.8922	19993.552
Andes			

	5 a:11	201506.0002	22402 746
Peruvian Andes	Millan	201506.0602	22492.746
Peruvian	Millan	243928.9025	22492.746
Andes	Willidii	243928.9023	22492.740
Peruvian	Millan	257596.3698	22492.746
Andes	Windh	237330.3030	22432.740
Peruvian	Millan	269451.6773	22492.746
Andes			
Peruvian	Millan	127893.3253	24991.94
Andes			
Peruvian	Millan	154302.045	24991.94
Andes			
Peruvian	Millan	187729.9853	24991.94
Andes			
Peruvian	Millan	274712.4319	24991.94
Andes			
Peruvian	Millan	347818.7383	24991.94
Andes			
Peruvian	Millan	351106.2523	24991.94
Andes			
Peruvian	Millan	352111.7874	24991.94
Andes			
Peruvian	Millan	458980.399	24991.94
Andes			
Peruvian	Millan	509939.3068	24991.94
Andes			
Peruvian	Millan	158994.1353	27491.134
Andes			
Peruvian	Millan	576009.3081	27491.134
Andes			
Peruvian	Millan	971372.528	27491.134
Andes			
Peruvian	Millan	388119.4622	29990.328
Andes			
Peruvian	Millan	283867.1944	32489.522
Andes			
Peruvian	Millan	480193.0404	32489.522
Andes			
Peruvian	Millan	915919.22	32489.522
Andes			
Peruvian	Millan	259964.9907	37487.91
Andes			
Peruvian	Millan	441044.2888	39987.104
Andes			
Peruvian	Millan	549338.2219	39987.104
Andes			

Demusian	NA:II	744262 2050	20007 104
Peruvian Andes	Millan	711362.3859	39987.104
Peruvian	Millan	279590.0093	42486.298
Andes		275550.0055	42400.250
Peruvian	Millan	1035316.75	42486.298
Andes			
Peruvian	Millan	1745859.088	42486.298
Andes			
Peruvian	Millan	1211503.827	44985.492
Andes			
Peruvian	Millan	469669.0906	49983.88
Andes			
Peruvian	Millan	629115.9608	49983.88
Andes			
Peruvian	Millan	1181985.759	57481.463
Andes			
Peruvian	Millan	823890.7933	62479.851
Andes			
Peruvian	Millan	1323722.276	62479.851
Andes		10207221270	02 17 5100 2
Peruvian	Millan	6479987.872	67478.239
Andes		0175507.072	07 17 0.235
Peruvian	Millan	1094667.727	87471.791
Andes	iviniari	1034007.727	0/4/1./51
Peruvian	Millan	2559460.231	92470.179
Andes		2000 1001201	52
Peruvian	Millan	3672256.877	94969.373
Andes		00722001077	5 15 05 10 7 0
Peruvian	Millan	1215588.203	99967.761
Andes			
Peruvian	Millan	903065.6992	102466.95
Andes			
Peruvian	Millan	5601141.657	109964.54
Andes			
Peruvian	Millan	3567263.881	124959.7
Andes			
Peruvian	Millan	2256120.556	127458.9
Andes			
Peruvian	Millan	2956851.605	129958.09
Andes			
Peruvian	Millan	2279867.781	134956.48
Andes			`
Peruvian	Millan	2486220.911	149951.64
Andes	-		
Peruvian	Millan	5593736.818	157449.22
Andes	-		
		1	1

	5 a:11	5470050 504	242424 40
Peruvian Andes	Millan	5173353.591	212431.49
Peruvian	Millan	7456338.042	222428.27
Andes			
Peruvian	Millan	5833073.696	299903.28
Andes			
Peruvian	Millan	11811480.16	344888.78
Andes			
Peruvian	OGGM	27923.1238	9996.7761
Andes			
Peruvian	OGGM	30056.225	9996.7761
Andes			
Peruvian	OGGM	30083.0718	9996.7761
Andes			
Peruvian	OGGM	32450.4724	9996.7761
Andes			
Peruvian	OGGM	39634.4349	9996.7761
Andes			
Peruvian	OGGM	42516.8061	9996.7761
Andes			
Peruvian	OGGM	46753.721	9996.7761
Andes			
Peruvian	OGGM	46802.5334	9996.7761
Andes			
Peruvian	OGGM	58753.0251	9996.7761
Andes			
Peruvian	OGGM	61003.276	9996.7761
Andes			
Peruvian	OGGM	67392.8169	9996.7761
Andes			
Peruvian	OGGM	69525.9181	9996.7761
Andes			
Peruvian	OGGM	70325.2208	9996.7761
Andes			
Peruvian	OGGM	73584.6677	9996.7761
Andes			
Peruvian	OGGM	402870.5644	9996.7761
Andes			
Peruvian	OGGM	32653.0438	12495.97
Andes			
Peruvian	OGGM	40748.5775	12495.97
Andes			
Peruvian	OGGM	43726.1329	12495.97
Andes			
Peruvian	OGGM	50432.9544	12495.97
Andes			

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Peruvian Andes	OGGM	54300.1154	12495.97
Peruvian	OGGM	72470.5251	12495.97
Andes			
Peruvian Andes	OGGM	105073.5362	12495.97
Peruvian	OGGM	107886.3497	12495.97
Andes			
Peruvian	OGGM	112904.2627	12495.97
Andes			
Peruvian	OGGM	117240.0227	12495.97
Andes		11, 2101022,	22 100107
Peruvian	OGGM	170130.6805	12495.97
Andes		1/0100000	12 100.07
Peruvian	OGGM	564573.787	12495.97
Andes		50-57575707	12.00.07
Peruvian	OGGM	60727.486	14995.164
Andes		00727.400	14555.104
Peruvian	OGGM	61958.7784	14995.164
Andes		01000.7704	14555.104
Peruvian	OGGM	64333.5008	14995.164
Andes		04353.5000	14555.104
Peruvian	OGGM	78275.5378	14995.164
Andes		78275.5578	14999.104
Peruvian	OGGM	102100.862	14995.164
Andes		102100.002	14555.104
Peruvian	OGGM	106071.7494	14995.164
Andes		100071.7454	14555.104
Peruvian	OGGM	128064.1687	14995.164
Andes		12000 11007	1 1000.101
Peruvian	OGGM	136411.0862	14995.164
Andes			
Peruvian	OGGM	492879.3788	14995.164
Andes			
Peruvian	OGGM	63214.4769	17494.358
Andes			
Peruvian	OGGM	102952.6381	17494.358
Andes			
Peruvian	OGGM	103688.4848	17494.358
Andes			1. 10 1.000
Peruvian	OGGM	69392.9043	19993.552
Andes			
Peruvian	OGGM	83897.504	19993.552
Andes			10000.002
Peruvian	OGGM	144900.78	19993.552
Andes		21.000070	
, 11005			

De la la la	00004	100010 0110	40000 550
Peruvian Andes	OGGM	180040.8146	19993.552
Peruvian	OGGM	643857.3005	19993.552
Andes		010007.0000	19999.992
Peruvian	OGGM	75659.194	22492.746
Andes			
Peruvian	OGGM	92172.4233	22492.746
Andes			
Peruvian	OGGM	150307.9717	22492.746
Andes			
Peruvian	OGGM	297513.9165	22492.746
Andes			
Peruvian	OGGM	72791.4665	24991.94
Andes			
Peruvian	OGGM	111181.1856	24991.94
Andes			
Peruvian	OGGM	177081.5639	24991.94
Andes			
Peruvian	OGGM	301959.5043	24991.94
Andes			
Peruvian	OGGM	447549.7591	24991.94
Andes			
Peruvian	OGGM	135155.3876	27491.134
Andes			
Peruvian	OGGM	155458.8984	27491.134
Andes			
Peruvian	OGGM	201528.0258	27491.134
Andes			
Peruvian	OGGM	262898.6149	27491.134
Andes			
Peruvian	OGGM	198312.51	29990.328
Andes			
Peruvian	OGGM	199338.7904	29990.328
Andes			
Peruvian	OGGM	229797.7175	29990.328
Andes			
Peruvian	OGGM	264199.4649	29990.328
Andes			
Peruvian	OGGM	184448.5729	32489.522
Andes			
Peruvian	OGGM	186902.6154	32489.522
Andes			
Peruvian	OGGM	198074.5496	32489.522
Andes			
Peruvian	OGGM	941268.711	32489.522
Andes			
L	1		

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Peruvian Andes	OGGM	352765.8734	37487.91
Peruvian	OGGM	270884.3208	39987.104
Andes		27000 1.0200	33307.101
Peruvian	OGGM	326307.121	39987.104
Andes			
Peruvian	OGGM	646717.7062	42486.298
Andes			
Peruvian	OGGM	1726735.617	47484.686
Andes			
Peruvian	OGGM	260714.2607	49983.88
Andes			
Peruvian	OGGM	839003.1071	49983.88
Andes			
Peruvian	OGGM	518266.6993	54982.269
Andes			
Peruvian	OGGM	1399623.091	54982.269
Andes			
Peruvian	OGGM	983939.2761	77475.015
Andes			
Peruvian	OGGM	2434922.757	79974.209
Andes			
Peruvian	OGGM	568726.5005	87471.791
Andes			
Peruvian	OGGM	1019838.343	87471.791
Andes			
Peruvian	OGGM	629522.3238	92470.179
Andes			
Peruvian	OGGM	1324530.121	102466.95
Andes			
Peruvian	OGGM	4134547.976	107465.34
Andes			
Peruvian	OGGM	2077455.031	109964.54
Andes			
Peruvian	OGGM	2629719.555	127458.9
Andes			
Peruvian	OGGM	7527018.373	152450.84
Andes			
Peruvian	OGGM	11053254.23	152450.84
Andes			
Peruvian	OGGM	3743932.98	159948.42
Andes			
Peruvian	OGGM	1491018.177	164946.81
Andes			
Peruvian	OGGM	8314285.235	169945.19
Andes			

Den in	00004	0005067.000	402427.04
Peruvian Andes	OGGM	9005967.689	192437.94
Peruvian	OGGM	3303426.848	207433.1
Andes			
Peruvian	OGGM	5247314.114	214930.69
Andes			
Peruvian	OGGM	12165338.21	237423.43
Andes			
Peruvian	OGGM	3163716.045	239922.63
Andes			
Peruvian	OGGM	10977006.84	244921.01
Andes			
Peruvian	OGGM	23457721.86	562318.65
Andes			
Peruvian	OGGM	20657523.84	587310.6
Andes			50701010
Peruvian	OGGM	40940600.56	949693.73
Andes			5 15 65 617 6
Peruvian	OGGM	132516485.3	1609481
Andes		10201010010	1000 101
Southern	Farinotti's Ensemble	26431.4328	9970.3595
Alps		20101.1020	3370.3333
Southern	Farinotti's Ensemble	70689.204	12462.949
Alps		70005.201	12 102.5 15
Southern	Farinotti's Ensemble	202145.9357	12462.949
Alps			
Southern	Farinotti's Ensemble	234748.9018	12462.949
Alps			
Southern	Farinotti's Ensemble	124256.1532	14955.539
Alps			
Southern	Farinotti's Ensemble	73343.0576	19940.719
Alps			
Southern	Farinotti's Ensemble	220582.9464	22433.309
Alps			
Southern	Farinotti's Ensemble	1533437.217	49851.798
Alps			
Southern	Farinotti's Ensemble	466728.0227	54836.977
Alps			
Southern	Farinotti's Ensemble	795624.2229	102196.19
Alps			
Southern	Farinotti's Ensemble	1138956.572	114659.13
Alps			
Southern	Farinotti's Ensemble	2851299.793	219347.91
Alps			
Southern	Farinotti's Ensemble	11151098.77	261721.94
Alps			
	1	1	

Southern	Farinotti's Ensemble	16604225.92	720358.48	
Alps		10004223.92	720338.48	
Southern Alps	GlabTop2	44301.8905	9970.3595	
Southern Alps	GlabTop2	56122.5233	9970.3595	
Southern	GlabTop2	69677.8064	9970.3595	
Alps Southern	ClahTan2	142270 0507	0070 2505	
Alps	GlabTop2	142279.0507	9970.3595	
Southern	GlabTop2	36088.0887	12462.949	
Alps				
Southern Alps	GlabTop2	101538.3507	12462.949	
Southern	GlabTop2	167433.7615	12462.949	
Alps				
Southern	GlabTop2	173071.9073	12462.949	
Alps				
Southern	GlabTop2	41382.408	17448.129	
Alps			17110 100	
Southern	GlabTop2	60568.8387	17448.129	
Alps	ClahTan2	102461.0560	17440 120	
Southern Alps	GlabTop2	183461.8569	17448.129	
Southern	GlabTop2	142577.8451	19940.719	
Alps	Glabiopz	142377.0431	15540.715	
Southern	GlabTop2	178358.0113	24925.899	
Alps				
Southern	GlabTop2	1422319.8	42374.028	
Alps				
Southern	GlabTop2	982266.0184	92225.826	
Alps				
Southern	GlabTop2	1794084.469	124629.49	
Alps		4040400 740	407000.44	
Southern	GlabTop2	1813122.719	137092.44	
Alps Southern	ClahTon2	3164173.07	226025 60	
Alps	GlabTop2	51041/5.0/	226825.68	
Southern	GlabTop2	69662293.74	1443209.5	
Alps		05002255.74	1773203.3	
Southern	HF	33429.671	9970.3595	
Alps				
Southern	HF	42554.7651	9970.3595	
Alps				
Southern	HF	59830.3725	9970.3595	
Alps				

		75000 0700	0070 0505
Southern Alps	HF	75060.9729	9970.3595
Southern	HF	127744.6228	9970.3595
Alps			
Southern	HF	177524.0039	12462.949
Alps			
Southern	HF	69253.6523	14955.539
Alps			
Southern	HF	108783.9614	14955.539
Alps			
Southern	HF	98072.0929	17448.129
Alps		4 4 4 2 4 7 6 6 2	47440.400
Southern	HF	144317.668	17448.129
Alps		200400 0072	27410 400
Southern Alps	HF	390488.6073	27418.489
Southern	HF	322423.1342	34896.258
Alps		522425.1542	54690.256
Southern	HF	644059.1187	34896.258
Alps		044033.1107	54050.250
Southern	HF	353706.7802	52344.388
Alps		333700.7002	52511.500
Southern	HF	657375.8531	52344.388
Alps			
Southern	HF	1585321.851	52344.388
Alps			
Southern	HF	545046.067	59822.157
Alps			
Southern	HF	592794.894	64807.337
Alps			
Southern	HF	489091.9578	69792.517
Alps			
Southern	HF	576031.6794	77270.286
Alps			
Southern	HF	721264.5952	79762.876
Alps		1017527.000	124500.05
Southern	HF	1017527.098	134599.85
Alps Southern	HF	3223806.427	171988.7
Alps	'''	3223000.427	1/1300./
Southern	HF	12179916.31	513473.52
Alps			515775.52
Southern	Millan	61705.596	9970.3595
Alps			
Southern	Millan	87813.5885	9970.3595
Alps	-		
Alps			

Southern	Millan	97334.5395	9970.3595
Alps	IVIIIdii	97334.3395	9970.3393
Southern	Millan	115863.44	9970.3595
Alps			
Southern	Millan	127647.256	9970.3595
Alps			
Southern	Millan	111577.7798	12462.949
Alps			
Southern	Millan	247749.4985	12462.949
Alps			
Southern	Millan	44126.9345	14955.539
Alps			
Southern	Millan	383072.3005	14955.539
Alps			
Southern	Millan	117422.2215	17448.129
Alps			
Southern	Millan	198460.6027	17448.129
Alps			
Southern	Millan	226466.9434	19940.719
Alps			
Southern	Millan	186999.3141	22433.309
Alps			
Southern	Millan	241502.0496	22433.309
Alps			
Southern	Millan	251157.3362	22433.309
Alps		247040.400	0.4005.000
Southern	Millan	217018.409	24925.899
Alps		222050.246	27442 402
Southern	Millan	223950.316	27418.489
Alps	N Aillen	172522 0702	22402.000
Southern	Millan	173532.8782	32403.668
Alps Southern	Millan	522053.964	34896.258
Alps	Willian	522055.904	54690.256
Southern	Millan	542503.4245	34896.258
Alps	Willian	542505.4245	54690.256
Southern	Millan	400847.0648	42374.028
Alps		-1000-1.0040	72377.020
Southern	Millan	697076.5541	54836.977
Alps		057070.5541	5-050.577
Southern	Millan	1796482.735	62314.747
Alps		1, 50402.755	52514.747
Southern	Millan	1372974.005	64807.337
Alps		10,20,4.000	0.007.007
Southern	Millan	1469206.17	69792.517
Alps		1,00200.17	00,02.01
, up3	L		

Southern	Millan	2421722 002	72205 107
Alps	willan	2421722.982	72285.107
Southern Alps	Millan	1073883.91	89733.236
Southern	Millan	2399083.073	132107.26
Alps Southern	Millan	6104118.329	256736.76
Alps	Willdi	0104118.529	250750.70
Southern	Millan	26013926.46	715373.3
Alps			
Southern	OGGM	37772.5342	9970.3595
Alps			
Southern	OGGM	40742.5257	12462.949
Alps			
Southern	OGGM	125241.6877	14955.539
Alps			
Southern	OGGM	159212.9615	14955.539
Alps		446007 6070	22422 202
Southern	OGGM	116087.6879	22433.309
Alps		101216 0200	22422 200
Southern	OGGM	181216.0309	22433.309
Alps Southern	OGGM	624460.4005	22422 200
Alps	OGGIM	624460.4005	22433.309
Southern	OGGM	245423.0407	24925.899
Alps		243423.0407	24525.055
Southern	OGGM	265447.7386	27418.489
Alps			
Southern	OGGM	443339.9107	39881.438
Alps			
Southern	OGGM	1344137.308	72285.107
Alps			
Southern	OGGM	3030189.974	92225.826
Alps			
Southern	OGGM	8100118.4	181959.06
Alps			
Southern	OGGM	6426822.609	224333.09
Alps			
Southern	OGGM	7708948.834	309081.15
Alps			440755.4
Southern	OGGM	59467533.2	418755.1
Alps	0000	20000117.01	446172 50
Southern	OGGM	20606117.01	446173.59
Alps Southern	OGGM	137190375.3	2103745.9
		12/1202/2.2	2103/43.9
Alps			

APPENDIX G

Table for total area and volume for Model_DEM combination of Himalayas

	1	
Model_DEM combination	Total_Volume (km ³)	Total_Area (km²)
OGGM_SRTM	0.670567729	15.25889461
OGGM_Copernicus	0.654524703	15.16643161
OGGM_ASTER	0.608397652	15.43132562
OGGM_ALOS	0.601324991	15.88614364
GlabTop2_ALOS	0.548224712	18.07776672
GlabTop2_Copernicus	0.510565309	19.16483177
GlabTop2_ASTER	0.472580287	18.11025372
GlabTop2_SRTM	0.44862059	20.37434782
Millan_SRTM	0.440641251	14.9515176
Millan_Copernicus	0.432667236	14.28428457
Millan_ALOS	0.401195427	13.82696755
Millan_ASTER	0.38348985	13.75949455
HF_ASTER	0.234557764	12.74240151
HF_Copernicus	0.234557764	12.07266948

HF_SRTM	0.229930942	13.09226152
HF_ALOS	0.195967344	11.96021448
Ensemble_ASTER	0.145437667	7.4870043
Ensemble_Copernicus	0.144676712	7.117152285
Ensemble_ALOS	0.132925872	7.327068293
Ensemble_SRTM	0.123354348	7.651938306

APPENDIX H

Table for pixel distribution in Alaska

	HF_I	Milan	OGGM_HF	OGGM_Millan	OGGM_GlabTop2	HF_GlabTop2	Millan_GlabTop2	GlabTop_Farinotti	HF_Farinotti	Farinotti_Millan	OGGM_Farinotti
1		21301	54608	18854	79121	85794	20677	48087	56654	21462	68819
2	2	5578	22326	3018	25035	36587	9471	69536	50354	8901	29645
Total		26879	76934	21872	104156	122381	30148	117623	107008	30363	98464

		No of Overdeepenings	Area of Overdeepenings (sq km)	s No of Overdeepenings	Area of Overdeepenings No of (sq km) Overo	No of Overdeepenings	Area of Overdeepenings (sq km)		4 Pixels	20 Pixels	xels
								No of Overdeenenings	Area of Overdeenenings	No of Area of Overdeeneines Overdeenenin	Area of Overdeenenin
	Farinotti's Ensemble	_0	524 52.05	135 135	50.75	49	49.13				5.61%
	Millan's		1640 45.72	72 413	41.44	81	31.43	3 74.82%	9.36%	92.06%	31.26%
ALASKA	Huss_Farinotti		750 40.34	34 220	38.27	66	35.44	1 70.67%	5.13%	91.20%	12.15%
	GlabTop2			52 148	50.88	68	49.21	1 70.16%	2.15%	86.29%	5.37%
	OGGM		512 22.93	33 117	21.66	32	19.96	77.15%		93.75%	12.95%
	Farinotti's Ensemble	ē	38 0.63	33 14	0.55	4	0.3	63.16%	12.70%	89.47%	52.38%
	Millan's		204 2.	2.3 52	1.77	80	0.98	3 74.51%	23.04%	96.08%	57.39%
EUROPEAN ALPS	Huss_Farinotti		79 0.88	38 22	0.67	4	0.34	1 72.15%	23.86%	94.94%	61.36%
	GlabTop2		54 1.19	17	1.06	9	0.67	68.52%	10.92%	88.89%	43.70%
	OGGM		56 1.48	18 21	1.35	6	1.04	4 62.50%	8.78%	83.93%	29.73%
	Farinotti's Ensemble	ē	60 1.79	13	1.63	9	1.52	2 78.33%	8.94%	%00.06	15.08%
NICHA ZEALAND	Millan's		117 2.24	30	1.94	10	1.52	2 74.36%	13.39%	91.45%	32.14%
	Huss_Farinotti		92 1.83	33 24	1.55	11	1.33	3 73.91%	15.30%	88.04%	27.32%
3001 DENN ALF3	GlabTop2		51 2.38	38 15	2.25	S	2.02	2 70.59%	5.46%	90.20%	15.13%
	OGGM		55 4.2	.2 18	4.06	80	3.85	5 67.27%	3.33%	85.45%	8.33%
	Farinotti's Ensemble	ē	354 3.92	32 79	2.99	20	1.91	1 77.68%	23.72%	94.35%	51.28%
	Millan's		419 5.62	52 106	4.47	22	2.84	4.70%	20.46%	94.75%	49.47%
ANDES	Huss_Farinotti		375 3.84	34 88	2.86	12	1.32	2 76.53%	25.52%	96.80%	65.63%
	GlabTop2		216 2.94	94 61	2.38	13	1.47	7 71.76%	19.05%	93.98%	50.00%
	OGGM		360 9.18	18	8.3	28	6.93	3 72.78%	9.59%	92.22%	24.51%
	Farinotti's Ensemble	ē	329 8.1	.1 104	7.33	30	5.86	68.39%	9.51%	90.88%	27.65%
	Millan's		996 16.33	33 277	13.83	62	9.38	3 72.19%	15.31%	93.78%	42.56%
HIMALAYAS	Huss_Farinotti		636 13.46	16 200	11.96	57	8.93	68.55%	11.14%	91.04%	33.66%
	GlabTop2		416 16.94	94 144	15.89	55	14.13	65.38%	6.20%	86.78%	16.59%
	OGGM		330 18.67	57 167	18.08	73	15.86	5 49.39%	3.16%	77.88%	15.05%

APPENDIX I